

3402 Waves and Wave Scattering Notes

Based on the 2015 spring lectures by Prof V
Smyshlyaev

The Author(s) has made every effort to copy down all the content on the board during lectures. The Author(s) accepts no responsibility whatsoever for mistakes on the notes nor changes to the syllabus for the current year. The Author(s) highly recommends that the reader attends all lectures, making their own notes and to use this document as a reference only.

3402 (Waves and Wave Scattering)

<i>Year:</i>	2014–2015
<i>Code:</i>	MATH3402
<i>Level:</i>	Advanced
<i>Value:</i>	Half unit (= 7.5 ECTS credits)
<i>Term:</i>	2
<i>Structure:</i>	3 hour lectures per week
<i>Assessment:</i>	100% examination
<i>Normal Pre-requisites:</i>	MATH7402
<i>Lecturer:</i>	Prof V Smyshlyaev

Course Description and Objectives

Modelling the propagation and scattering of acoustic and electromagnetic waves has proved a major challenge to mathematicians and physicists for many centuries, and its practical importance can be observed in many applications prevalent throughout our modern world. These include the mitigation of aircraft, rail and traffic noise in urban areas, sonar detection, wireless and fibre optic communication, baggage screening, medical diagnostics and the workings of the cochlea. This course aims to provide an introduction to linear and nonlinear wave theory and the approximate methods used to tackle wave transmission and scattering in inhomogeneous media.

Recommended Texts

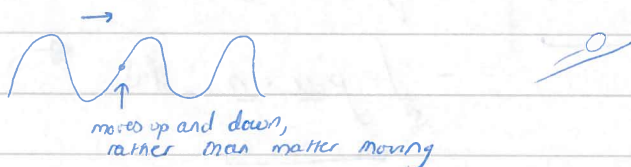
- (i) Pierce, A.D, *Acoustics: an introduction to its physical principles and applications*, Acoustic Society of America 1989.
- (ii) Billingham, J. and King, A.C. *Wave Motion*, CUP 2001.

Detailed Syllabus

- Acoustic waves - governing equations, plane acoustic waves, spherically symmetric waves, causality and the Sommerfeld radiation condition, acoustic energy and intensity.
- Electromagnetic (EM) waves - governing equations, plane EM waves, Poynting's vector.
- Impedance and surface boundary conditions, interfacial boundary conditions.
- Plane wave reflection and transmission at interfaces - reflection by acoustically soft and hard boundaries and by a perfect conductor, reflection and transmission between two insulators.
- Radiation from vibrating bodies - a radially pulsating sphere, a transversely oscillating sphere.
- Green's functions, monopoles, dipoles, quadrupoles, multipole expansions.
- Kirchoff-Helmholtz integral theorem, acoustic scattering by air bubbles in water, acoustic scattering by a fixed rigid sphere.
- Introduction to the WKB approximation, slowly varying waveguides, optic fibres.

100% Exam, Office 1-2pm Monday

0. Introduction: The course is about mainly waves (acoustic/sound, electromagnetic), and their scattering by various "obstacles"



- water waves
- sound (acoustic) - disturbing air & sending disturbances
- radio / mobiles / light (electromagnetic waves)
- elastic, etc.

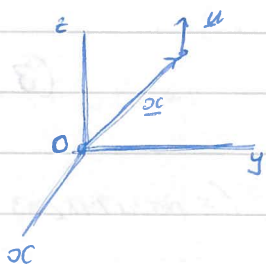
Common feature: waves ^(move) propagate (then transfer energy/information; they scatter when hit obstacles; disturbances, not matter, move)
 These are described by partial differential equations (PDE)

1. Governing equations

1.1 Acoustic (sound) waves

Acoustic (sound) waves are small amplitude mechanical disturbances propagating in a fluid, typically a compressible gas (e.g. air) or liquid (e.g. water)

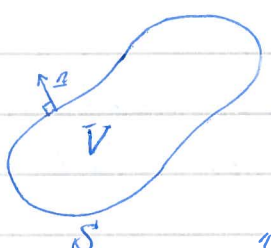
The governing equations are derived from basic conservation laws of "inviscid fluid":



- $\underline{x} = (x, y, z)$, t time
- $\rho(\underline{x}, t)$ fluid density
- $\underline{u}(\underline{x}, t)$ fluid velocity
- $p(\underline{x}, t)$ fluid pressure

Conservation of mass (q. MATH12301) - won't heavily rely on this i.e. will derive

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{u}) = 0 \quad (1)$$



$$\int_V (1) dV$$

$$\frac{d}{dt} \left(\int_V \rho dV \right) = - \int_V \nabla \cdot (\rho \mathbf{u}) =$$

rate of change " " " " " " " " " " " "

$$= - \int_S \rho \mathbf{u} \cdot \mathbf{n} dS \quad \leftarrow \text{Divergence theorem}$$

mass outflow rate

Newton's 2nd law holds for continuous mechanics

Conservation of momentum (\equiv Newton's 2nd Law for continuous media; neglecting gravity, viscosity etc)

nonlinear \rightarrow

$$\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} = - \frac{1}{\rho} \nabla p \quad (2)$$

acceleration \rightarrow

$$\mathbf{a} = \frac{d\mathbf{u}}{dt} = \frac{d}{dt} \mathbf{u}(\mathbf{x}(t), t) = \sum_j \frac{\partial \mathbf{u}}{\partial x_j} \frac{\partial x_j}{\partial t} + \frac{\partial \mathbf{u}}{\partial t}$$

acc. of a material point

$$= \frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \text{ etc.}$$

Additionally, physically, an equation of state is held connecting one unknown to another
e.g. pressure p to density ρ

Under the "adiabatic" assumption (no heat exchange)
let

$$\rho = P(\rho) \quad (3)$$

We will be considering small disturbances (= perturbations) of a uniform state at rest:

$$\rho = \rho_0 + \tilde{\rho}, \quad p = p_0 + \tilde{p}, \quad \mathbf{u} = \frac{0}{\rho_0} + \tilde{\mathbf{u}}$$

ρ_0, ρ_0 constant; $\tilde{\rho}, \tilde{p}, \tilde{u}$ "small"

Plugging in to (1) - (3)

$$(1) \Rightarrow \frac{\partial}{\partial t} (\rho_0 + \tilde{\rho}) + \nabla \cdot ((\rho_0 + \tilde{\rho}) \tilde{u}) = 0 \quad (1')$$

neglect to good approximation
diff. a constant

$$(2) \Rightarrow \frac{\partial \tilde{u}}{\partial t} + \tilde{u} \cdot \nabla \tilde{u} = - \frac{1}{\rho_0 + \tilde{\rho}} \nabla (\rho_0 + \tilde{p}) \quad (2')$$

small so neglect *small so neglect* *diff. a constant*

$$(3) \Rightarrow \rho_0 + \tilde{\rho} = \underline{P}(\rho_0 + \tilde{p}) \quad (3')$$

Since $\tilde{\rho}, \tilde{p}, \tilde{u}$ small, neglecting in (1'), (2') the higher order smallness terms, like, $\nabla \cdot (\tilde{\rho} \tilde{u})$

$$(1') \Rightarrow \frac{\partial \tilde{\rho}}{\partial t} + \rho_0 \nabla \cdot \tilde{u} = 0 \quad (4)$$

$$(2') \Rightarrow \frac{\partial \tilde{u}}{\partial t} = - \frac{1}{\rho_0} \nabla \tilde{p} \quad (5)$$

Use Taylor expansion for (3) (about ρ_0):

$$P(\rho_0 + \tilde{p}) = \underline{P}(\rho_0) + \frac{dP}{d\rho}(\rho_0) \tilde{p} + O(\tilde{p}^2) \quad \leftarrow \text{eventually neglect}$$

\Rightarrow To leading orders of smallness,

$$\rho_0 + \tilde{\rho} = \underline{P}(\rho_0) + \frac{dP}{d\rho}(\rho_0) \tilde{p}$$

$\underline{P}(\rho_0) = \rho_0$ $\frac{dP}{d\rho}(\rho_0) = \rho'$

$$\Rightarrow \tilde{\rho} = \underline{P}' \tilde{p}, \quad \text{where } \underline{P}' = \frac{dP}{d\rho}(\rho_0) \text{ is a constant}$$

characterising the medium's physical properties; physically $\underline{P}' > 0$
 $\Leftrightarrow \underline{P}' = c^2$, where $c := (\underline{P}')^{1/2} > 0$ is the wave speed (will soon see describes the speed of waves)

$$\tilde{\rho} = \frac{1}{c^2} \tilde{p} \quad (6)$$

So eliminating \tilde{p} in (4) - (5) via (6):

$$\frac{1}{c^2} \frac{\partial \tilde{p}}{\partial t} + \rho_0 \nabla \cdot \tilde{u} = 0 \quad (7)$$

$$\frac{\partial u}{\partial t} = -\frac{1}{\rho_0} \nabla \tilde{p} \quad (8)$$

could be some other non-linear equations

(7)-(8) main equations of ("linear") acoustics
(A (vector) PDE system)

To start analysing (7)-(8):

Dropping 'wiggles' (" \sim ") henceforth, differentiate (7) in t and use (8):

$$\frac{1}{c^2} \frac{\partial^2 p}{\partial t^2} + \rho_0 \nabla \cdot \frac{\partial u}{\partial t} = 0 \quad \Rightarrow$$

$$\frac{1}{c^2} \frac{\partial^2 p}{\partial t^2} + \rho_0 \nabla \cdot \left(-\frac{1}{\rho_0} \nabla p \right) = 0$$

$$\frac{1}{c^2} \frac{\partial^2 p}{\partial t^2} - \nabla^2 p = 0 \quad \text{WAVE EQN} \quad (9)$$

where $\nabla^2 p := \nabla \cdot (\nabla p) = \frac{\partial^2 p}{\partial x^2} + \frac{\partial^2 p}{\partial y^2} + \frac{\partial^2 p}{\partial z^2}$
 $=: \Delta p$
 (= Laplacian of p)

(9) is the wave equation (a scalar PDE).

Plane waves: For solutions of (9)

$p(t, x, y, z)$, let p does not depend on y & z
 in reality depends on only 2 of the variables

$$p = p(t, x); \quad \text{so (9)} \Rightarrow$$

$$\frac{1}{c^2} \frac{\partial^2 p}{\partial t^2} - \frac{\partial^2 p}{\partial x^2} = 0 \quad (9')$$

$$\Rightarrow p(x, t) = f(x - ct) + g(x + ct) \quad \text{solves (9')}$$

$\forall f, g \in C^2$
 functions diff. twice continuously

$$\frac{1}{c^2} [(-c^2) f'' + c^2 g''] - (f'' + g'') = 0$$

Decided $p = p(t, x)$, but could choose y, z

More generally, for any direction \underline{n} , $|\underline{n}| = 1$,

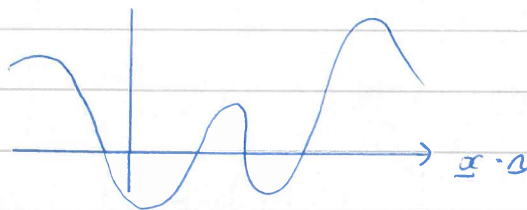
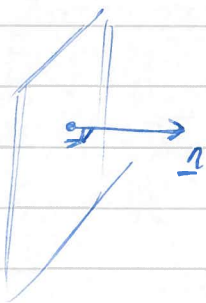
$$p = f(\underline{x} \cdot \underline{n} - ct) + g(\underline{x} \cdot \underline{n} + ct) \quad (10)$$

\uparrow
Projection i.e. works for any direction

solves (9) $\forall f, g \in C^2$

In (10) $f(\underline{x} \cdot \underline{n} - ct)$ describes a 'plane wave' moving in positive \underline{n} direction

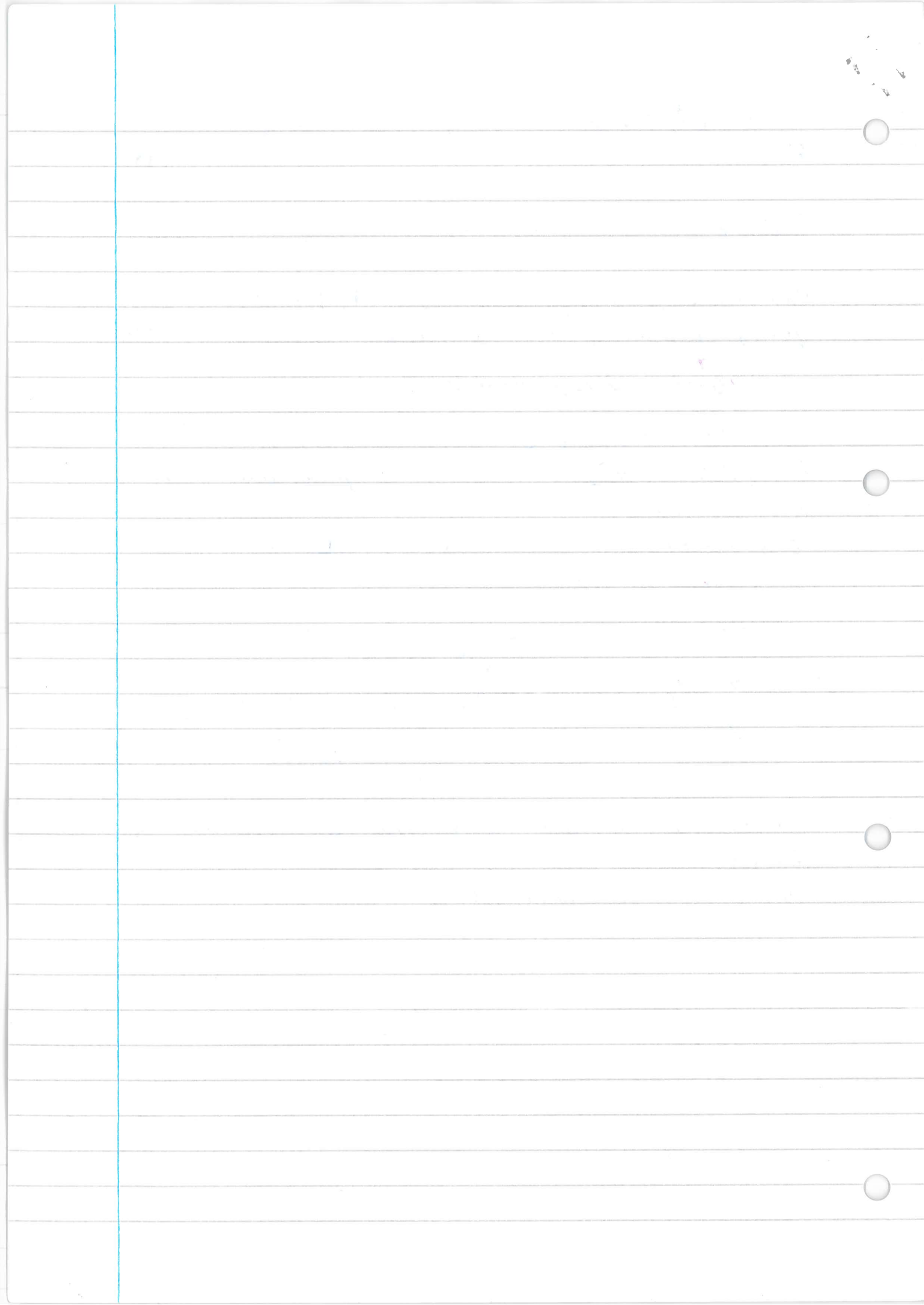
$f(\underline{x} \cdot \underline{n} - ct)$ is constant \forall plane \perp to \underline{n} , which plane moves with speed c



2D cross section

Similarly $g(\underline{x} \cdot \underline{n} + ct)$ describes plane wave in -ve \underline{n} direction

Thus c is the "phase velocity"

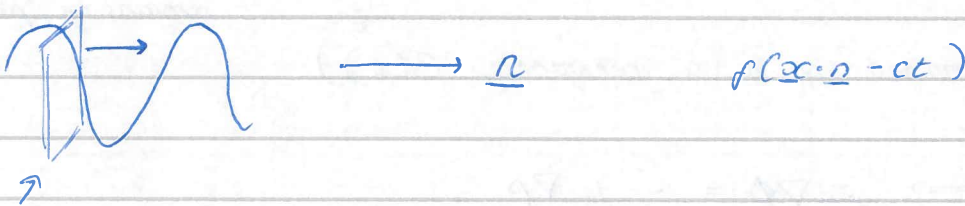


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$$\begin{cases} \frac{1}{c^2} \frac{\partial p}{\partial t} + \rho_0 \nabla \cdot \underline{u} = 0 & (7) \\ \frac{\partial \underline{u}}{\partial t} = -\frac{1}{\rho_0} \nabla p & (8) \end{cases}$$

$$\Leftrightarrow \frac{1}{c^2} \frac{\partial^2 p}{\partial t^2} - \nabla^2 p = 0 \quad (9)$$

a priori
we don't know
this but
it is
true



Remark: If p solves (9), \underline{u} has to be recovered from (7)-(8), this depends on chosen initial/boundary conditions e.g.

If $\forall t < t_0, \underline{u} \equiv 0, p \equiv 0 \Rightarrow$ from (8)

$$\underline{u}(\underline{x}, t) = -\frac{1}{\rho_0} \int_{t_0}^t \nabla p(\underline{x}, t') dt'$$

Then (7) is also satisfied (check!) \square

To eliminate p to end up with equation for \underline{u} is subtle since p scalar, \underline{u} vector

Let at $t=0$, $\nabla \times \underline{u} = 0$ (i.e. fluid is initially irrotational) \Rightarrow taking curl of (8):

$$\nabla \times \left(\frac{\partial \underline{u}}{\partial t} \right) = \frac{\partial}{\partial t} (\nabla \times \underline{u}) = -\frac{1}{\rho_0} \nabla \times (\nabla p) = \underline{0}$$

can interchange

when curl meets gradient they kill each other

$$\Rightarrow \nabla \times \underline{u} = \underline{c} \Rightarrow \underline{c} = 0 \quad (t=0)$$

\Rightarrow fluid remains irrotational $\forall t$

$$\iff \nabla \times \underline{u} = 0$$

$\Rightarrow \exists \Phi(x, t)$ a (scalar) velocity potential s.t. $\underline{u} = \nabla \Phi$

value of integral does not depend on path
 $\oint \underline{u} \cdot d\underline{r} = 0$ so circulation is 0

(Φ determined up to a constant, $C(t)$)
done w.r.t. \underline{x} but t remains a parameter

$$\text{Then (8)} \Rightarrow \frac{\partial}{\partial t} \nabla \Phi = - \frac{1}{\rho_0} \nabla p$$

$$\iff \nabla \left(\frac{\partial \Phi}{\partial t} + \frac{1}{\rho_0} p \right) = 0$$

$$\Rightarrow \frac{\partial \Phi}{\partial t} + \frac{1}{\rho_0} p = A(t)$$

constant, may depend on t

without loss of generality (wlog) $A(t) \equiv 0$

We can do this since Φ determined up to a constant

Re-defining $\Phi \rightarrow \Phi - \int_0^t A(t') dt' = C(t)$

$$\Rightarrow \frac{\partial \Phi}{\partial t} = - \frac{1}{\rho_0} p \quad (11)$$

Differentiating (11) in t , via (7)

$$\frac{\partial^2 \Phi}{\partial t^2} = - \frac{1}{\rho_0} \frac{\partial p}{\partial t} = - \frac{1}{\rho_0} \left(- \rho_0 c^2 \nabla \cdot \overbrace{(\nabla \Phi)}^{\underline{u}} \right)$$

$= \nabla^2 \Phi$

$$\Rightarrow \frac{1}{c^2} \frac{\partial^2 \Phi}{\partial t^2} - \nabla^2 \Phi = 0 \quad (12)$$

i.e. Φ solves the same wave equation as p , cf (9)

\langle Check 'reverse' argument: Φ solves (12) $\Rightarrow \underline{u} = \nabla \Phi$,
 and p via (11), $p = -\rho_0 \frac{\partial \Phi}{\partial t}$ solve (7) and (8) \rangle

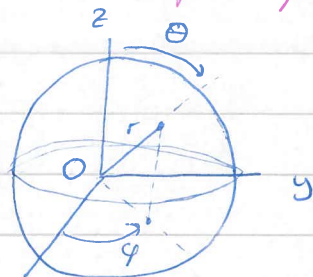
Spherically symmetric waves

Seek solutions of wave equation (9) / (12) in spherical

$$x = (r, \theta, \varphi)$$

different from vel. potential ϕ

Remember that Laplace is always combination of 2nd derivatives



$$\nabla^2 p = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 \frac{\partial p}{\partial r})$$

$$+ \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} (\sin \theta \frac{\partial p}{\partial \theta})$$

$$+ \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2 p}{\partial \varphi^2}$$

Seek spherically symmetric solutions of (12) (9)

i.e. $p = p(t, r, \theta, \varphi) = p(t, r)$

assume no dependence on spherical coordinates

$$\Rightarrow \frac{1}{c^2} \frac{\partial^2 p}{\partial t^2} - \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 \frac{\partial p}{\partial r}) = 0$$

Try $p = \frac{q}{r} \Rightarrow \frac{\partial p}{\partial r} = \frac{1}{r} \frac{\partial q}{\partial r} - \frac{q}{r^2}$

$$\Rightarrow r^2 \frac{\partial p}{\partial r} = r \frac{\partial q}{\partial r} - q$$

$$\Rightarrow \frac{\partial}{\partial r} (r^2 \frac{\partial p}{\partial r}) = \frac{\partial q}{\partial r} + r \frac{\partial^2 q}{\partial r^2} - \frac{\partial q}{\partial r} = r \frac{\partial^2 q}{\partial r^2}$$

$$\frac{1}{c^2} \frac{1}{r} \frac{\partial^2 q}{\partial t^2} - \frac{1}{r} \frac{\partial^2 q}{\partial r^2} = 0$$

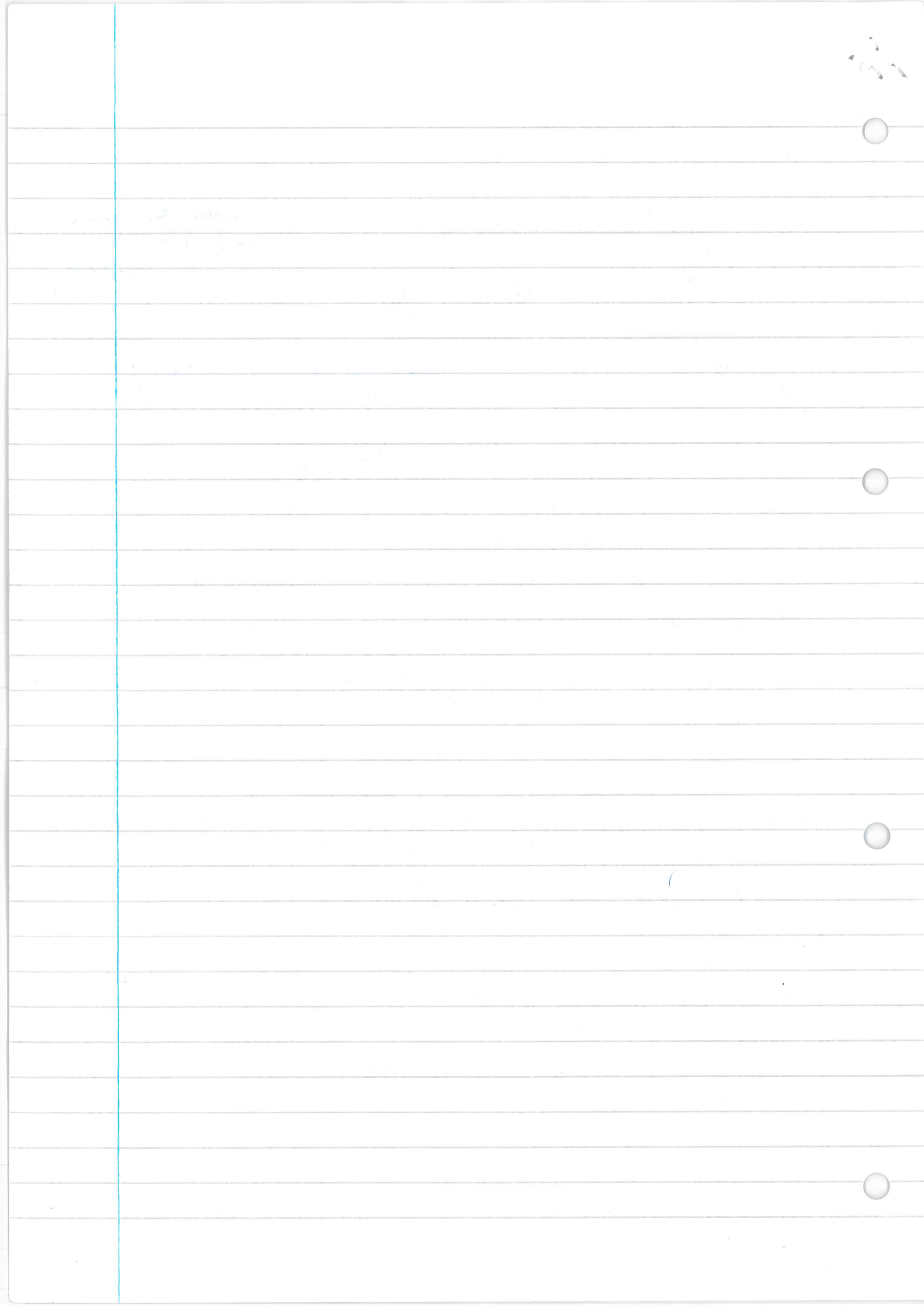
$\Rightarrow q$ solves 1-dimensional wave equation

$$\frac{1}{c^2} \frac{\partial^2 q}{\partial t^2} - \frac{\partial^2 q}{\partial r^2} = 0$$

$$\Rightarrow q = f(r - ct) + g(r + ct)$$

$$\Rightarrow p = \frac{f(r - ct)}{r} + \frac{g(r + ct)}{r} \quad (13)$$

Which is a general form for a radially symmetric solution of (9) or (12)



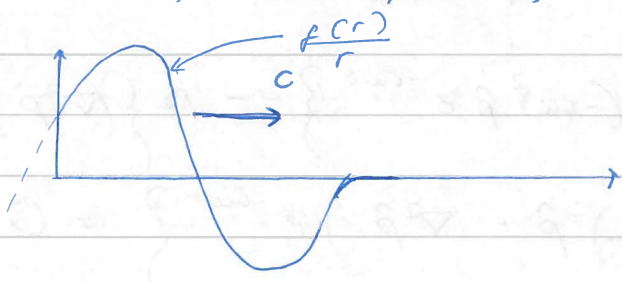
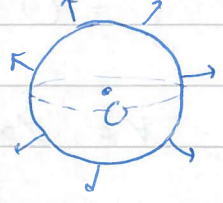
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$$p(r, t) = \frac{f(r-ct)}{r} + \frac{g(r+ct)}{r} \quad (13)$$

$$\forall f, g \in C^2$$

In (13), f -part describes a spherical wave travelling away from centre / origin $r=0$, with speed c ;

see waves moving away when throw stone \rightarrow



g -part travels towards 0
(f "outgoing", g "incoming")

Time-harmonic (T-H) waves

time dep. sinusoidal

Seek solutions of (9) with following time dependence:

$$p(x, t) = |\hat{p}(x)| \cos(\omega t - \psi(x))$$

modulus since allowed to be complex



i.e. "pure tone" solution with angular frequency $\omega \Rightarrow$

$$T = \frac{2\pi}{\omega} \text{ time period}$$

Phase $\psi(x)$, amplitude $|\hat{p}(x)| \iff$

$$p(x, t) = \text{Re} \left\{ \underbrace{|\hat{p}(x)| e^{i\psi(x)}}_{=: \hat{p}(x)} e^{-i\omega t} \right\}$$

$$p(x, t) = \text{Re} \left\{ \hat{p}(x) e^{-i\omega t} \right\} \quad (14)$$

where $\hat{p}(\underline{x}) := |\hat{p}(\underline{x})| e^{+i\phi(\underline{x})}$ is called "complex amplitude"
 i.e. incorporating both the amplitude and the phase.

Plug (14) in to (9)

$$\frac{1}{c^2} \frac{\partial^2 \hat{p}}{\partial t^2} - \nabla^2 \hat{p} = 0 \quad \Rightarrow$$

can take outside

$$\frac{1}{c^2} \operatorname{Re} \left\{ (-\omega)^2 \hat{p} e^{-i\omega t} \right\} - \operatorname{Re} \left\{ (\nabla^2 \hat{p}) e^{-i\omega t} \right\} = 0$$

$$- \operatorname{Re} \left\{ \left(\left(\frac{\omega}{c} \right)^2 \hat{p} + \nabla^2 \hat{p} \right) e^{-i\omega t} \right\} = 0 \quad \forall t \quad \forall \underline{x}$$

$$=: S'$$

$$(\operatorname{Re} S') \cos \omega t + (\operatorname{Im} S') \sin \omega t = 0 \quad \forall t$$

$$\Rightarrow (\text{varying } t) \quad \operatorname{Re} S = 0, \operatorname{Im} S = 0 \rightarrow$$

$$S = 0$$

$$\nabla^2 \hat{p} + k^2 \hat{p} = 0$$

$$k := \frac{\omega}{c} \quad (15)$$

(15) is called Helmholtz equation ("reduced wave equation" for T-H waves);

$k := \frac{\omega}{c}$ is wavenumber

Time dependence must solve this equation

For velocity $\underline{u}(\underline{x}, t)$ ⁱⁿ ~~for~~ T-H case also seek in form

$$\underline{u}(\underline{x}, t) = \operatorname{Re} \left\{ \underline{\hat{u}}(\underline{x}) e^{-i\omega t} \right\} \rightarrow$$

$$\text{via (8)} \quad \leftrightarrow \quad \frac{\partial \underline{u}}{\partial t} = -\frac{1}{\rho_0} \nabla p$$

p has same form as before

$$\text{Re} \left\{ \underbrace{(-i\omega \hat{u} + \frac{1}{\rho_0} \nabla \hat{p})}_{=: S} e^{-i\omega t} \right\} = 0 \quad \forall t$$

Repeating same argument $\Rightarrow S=0 \Rightarrow$

$$\hat{u} = -\frac{i}{\rho_0 \omega} \nabla \hat{p} \quad (16)$$

Plane T-H waves:

For a plane wave travelling in a positive \underline{n} direction, which is also T-H:

$$p(\underline{x}, t) = \underset{\text{(10) plane}}{f(\underline{x} \cdot \underline{n} - ct)} = \underset{\text{14 T-H}}{\text{Re} \left\{ \hat{p}(\underline{x}) e^{-i\omega t} \right\}}$$

$$= \text{Re} \left\{ e^{i \frac{\omega}{c} (-ct + \underline{x} \cdot \underline{n})} A \right\} \quad A \in \mathbb{C}$$

(a complex constant)

$$p(\underline{x}, t) = \text{Re} \left\{ \underbrace{A e^{ik \underline{x} \cdot \underline{n}}}_{=: \hat{p}(\underline{x})} e^{-i\omega t} \right\} \quad (17)$$

$$\iff \hat{p}(\underline{x}) = A e^{ik \underline{x} \cdot \underline{n}} = A e^{i \underline{k} \cdot \underline{x}}$$

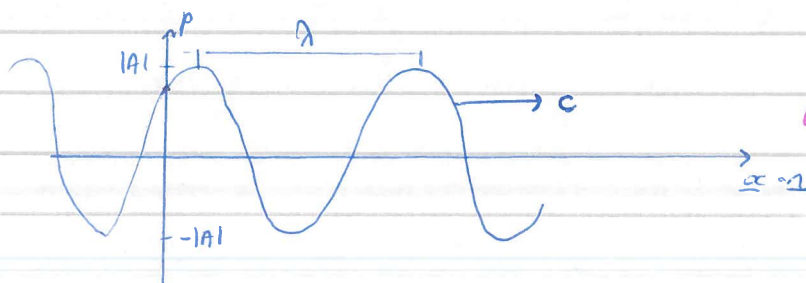
where $\underline{k} := k \underline{n}$, $|\underline{k}| = 1$

is called wavevector

could have + or -

$$\text{So } (17) \iff (A = |A| e^{-i\psi})$$

$$p(\underline{x}, t) = |A| \cos(k \underline{x} \cdot \underline{n} - \psi - \omega t)$$



draw
before ↑ function of t,
now function of x

$\lambda = \frac{2\pi}{k}$ is the wave-length (= spatial period)

$\left(T = \frac{2\pi}{\omega}, f = \frac{1}{T} = \frac{\omega}{2\pi} \right)$ frequency, $A = |A| e^{-i\phi}$
 ↑
 complex amplitude ← conventional amplitude

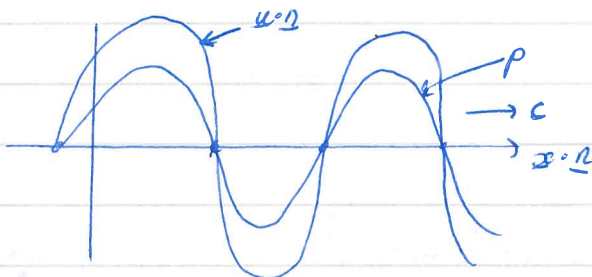
Also plugging (17) in to (16)

$$\begin{aligned} \underline{u} &= \frac{-i}{\rho_0 \omega} \nabla \hat{p} = \frac{-i}{\rho_0 \omega} \nabla (A e^{i\mathbf{k} \cdot \mathbf{x}}) \\ &= \frac{-i}{\rho_0 \omega} A i \mathbf{k} e^{i\mathbf{k} \cdot \mathbf{x}} = \frac{A}{\rho_0 \omega} \mathbf{k} e^{i\mathbf{k} \cdot \mathbf{x}} \\ &= \frac{A}{\rho_0 c} e^{i\mathbf{k} \cdot \mathbf{x}} \Rightarrow \end{aligned}$$

$$\underline{u} = \text{Re} \left\{ \frac{A \mathbf{k}}{\rho_0 c} e^{i\mathbf{k} \cdot \mathbf{x} - i\omega t} \right\}$$

$$= \frac{p_0}{\rho_0 c} \quad (18)$$

ie. for plane T-H waves, velocity \underline{u} and pressure p are "in phase", and \underline{u} is in the \underline{n} -direction (ie. sound wave "longitudinal")



Remark: Every acoustic field, described by wave equation (9) can be expressed as a 'superposition' of T-H waves, via Fourier Transform (FT) in time:

Assuming sufficient decay of $p(\underline{x}, t)$ for $t \rightarrow \pm \infty$, seek $p(\underline{x}, t)$, a solution to (9), as a FT of $p^*(\underline{x}, \omega)$:

$$p(\underline{x}, t) = (2\pi)^{-1/2} \int_{-\infty}^{\infty} p^*(\underline{x}, \omega) e^{-i\omega t} d\omega \quad (19)$$

\Rightarrow the inverse FT gives:

$$p^*(\underline{x}, \omega) = (2\pi)^{-1/2} \int_{-\infty}^{\infty} p(\underline{x}, t) e^{i\omega t} dt$$

Then (9) \Rightarrow

$$0 = \frac{1}{c^2} \frac{\partial^2 p}{\partial t^2} - \nabla^2 p = (2\pi)^{-1/2} \int_{-\infty}^{\infty} \left[\frac{(-i\omega)^2}{c^2} p^* - \nabla^2 p^* \right] e^{-i\omega t} d\omega$$

$$= - (2\pi)^{-1/2} \int_{-\infty}^{\infty} \left(\nabla^2 p^* + k^2 p^* \right) e^{-i\omega t} d\omega = 0 \quad \forall t$$

$k = \frac{\omega}{c}$

\Rightarrow (by FT inversion of zero)

$$\nabla^2 p^* + k^2 p^* = 0$$

i.e. p^* , the inverse FT of p , solves Helmholtz equation (15). So (19), for real-valued $p(\underline{x}, t)$ gives

$$p(\underline{x}, t) = \text{Re } p(\underline{x}, t) = \int_{-\infty}^{\infty} \underbrace{\text{Re} \left\{ (2\pi)^{-1/2} p^*(\underline{x}, \omega) e^{-i\omega t} \right\}}_{=: \hat{p}} d\omega$$

i.e. p is a superposition (= integral in ω) of T-H waves

$$\text{Re} \left\{ (2\pi)^{-1/2} p^*(\underline{x}, \omega) e^{-i\omega t} \right\} \quad \blacksquare$$

Causality and Sommerfeld Radiation Condition

Causality: the 'effect' / consequence can only follow the 'cause' / source which can only effect future but not past.

How the wave equation (9) treats future ($t \uparrow$) and past ($t \downarrow$) equally:

$$(9): \frac{1}{c^2} \frac{\partial^2 p}{\partial t^2} - \nabla^2 p = 0$$

Let $t \rightarrow -t$ ($t^2 = -t$) \Rightarrow

$$\frac{1}{c^2} \frac{\partial^2 p}{\partial (t^2)^2} - \nabla^2 p = 0$$

not discriminating between past + future

so an additional mathematical condition needs to be introduced, called causality condition

It means the solution to (9) is required to be identically zero for $t \leq t_0$ i.e. before some time t_0 when the "source" switches on

Consider spherically symm. waves (13)

$$p(r, t) = \frac{f(r-ct)}{r} + \frac{g(r+ct)}{r}$$

$p, g \in C^2$
twice differentiable

We require: $\exists t_0$ s.t. outside a "source" region $r \leq r_0$
i.e. $\forall r > r_0, \forall t \leq t_0$ it holds $p(r, t) \equiv 0$

\Rightarrow (Exercise: Exam 2013 Q 1 (d))

$g(\xi)$ is a constant so could be added to $f \Rightarrow$

$$p(x, t) = \frac{f(r-ct)}{r}$$

is a general "causal" solution of (9) i.e. only outgoing wave ~~is~~ remains (and no incoming wave)

Consider now radially symmetric T-H waves

$$(13) \Rightarrow \overset{\text{TH}}{p(\omega, t)} = \text{Re} \left\{ A \frac{e^{\frac{i\omega}{c}(-ct+r)}}{r} + B \frac{e^{\frac{-i\omega}{c}(ct+r)}}{r} \right\} \quad \text{This is our guess}$$

$\left. \begin{array}{l} \text{this is } (r-ct) \\ \text{radially symmetric} \end{array} \right\}$

$$= \text{Re} \left\{ \left(A \frac{e^{ikr}}{r} + B \frac{e^{-ikr}}{r} \right) e^{-i\omega t} \right\} \quad A, B \in \mathbb{C}$$

$\left. \begin{array}{l} \text{this is } \hat{p} \end{array} \right\}$

$$\Leftrightarrow \hat{p} = A \frac{e^{ikr}}{r} + B \frac{e^{-ikr}}{r} \quad (13')$$

We expect physically (13') not to contain an incoming part

$\Leftrightarrow B=0$. This is achieved by the Sommerfeld radiation condition

\uparrow
means same as outgoing

$$r \frac{\partial \hat{p}}{\partial r} - ikr \hat{p} \rightarrow 0 \quad \text{as } r \rightarrow \infty \quad (20)$$

uniformly with respect to a direction (holds not only for spher. symm solutions but \forall with 'bounded' sources)

Check: (13') \Rightarrow

$$r \frac{\partial \hat{p}}{\partial r} = A i k e^{ikr} - \frac{A}{r} e^{ikr} - B i k e^{-ikr} - \frac{B}{r} e^{-ikr}$$

$$-i k r \hat{p} = -i k A e^{ikr} - i k B e^{-ikr}$$

$$r \frac{\partial \hat{p}}{\partial r} - i k r \hat{p} = -2 i k B e^{-ikr} - \frac{A e^{ikr} + B e^{-ikr}}{r}$$

$\not\rightarrow 0$ as $r \rightarrow \infty$
unless $B=0$

$(|e^{\pm ikr}|=1) \rightarrow 0$
as $r \rightarrow \infty$

$\Rightarrow B=0$ as desired \square

Remark: Sommerfeld radiation condition (20) ensures, physically, the T-H waves are outgoing; mathematically, it ensures uniqueness of solutions of boundary-value problems per (15)

[Often require additionally
 $p(\infty, t) = O\left(\frac{1}{r}\right)$ as $r := |\underline{x}| \rightarrow \infty$ most decay not slower than $\frac{1}{r}$ (20')]

One can see that FTs of causal solutions to (9) satisfying (20) (& (20'))

Back to radially-symmetric time T-H outgoing waves
 $\Rightarrow p = \text{Re} \left\{ \hat{p} e^{-i\omega t} \right\} = \text{Re} \left\{ \frac{A e^{i(kr - i\omega t)}}{r} \right\}$
 $= \hat{p}$

$$(16) \Rightarrow \underline{\hat{u}} = \frac{-i}{\rho_0 \omega} \nabla \hat{p} = \frac{-i}{\rho_0 \omega} \nabla \left(A \frac{e^{i(kr - i\omega t)}}{r} \right)$$

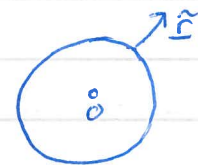
$$= \frac{-iA}{\rho_0 \omega} \left(\frac{ik}{r} - \frac{1}{r^2} \right) e^{i(kr - i\omega t)} \underline{\tilde{r}}$$

\leftarrow unit vector in radial direction

$\underline{\tilde{r}} = \frac{\underline{x}}{|\underline{x}|} = \frac{\underline{x}}{r}$
 is a unit vector in radial direction

$$\Rightarrow \underline{\hat{u}} = \hat{u}_r \underline{\tilde{r}}, \quad \hat{u}_r = \frac{A e^{i(kr - i\omega t)}}{r} \left(ik - \frac{1}{r} \right) \frac{1}{\rho_0 \omega}$$

$= \hat{p}$



26/01/15

Sph symm + T-H + outgoing

$$p = \text{Re} \left\{ A \frac{e^{i k r}}{r} e^{-i \omega t} \right\}$$

$$u = \text{Re} \left\{ \hat{u} e^{-i \omega t} \right\}$$

connects the two \rightarrow (16) $\leftrightarrow \hat{u} = \frac{-i}{\omega c} \nabla \hat{p} = \hat{u}_r \hat{r}$

$$\hat{u}_r = \frac{A}{r} e^{i k r} \left(i k - \frac{1}{r} \right) \frac{-i}{\omega c}$$

← accidentally missed this factor beforehand

$$= \underbrace{\frac{A}{r} e^{i k r}}_{\hat{p}} \left(\frac{1 + i}{k r} \right) \frac{k}{\omega c}$$

$$\hat{u} = \hat{u}_r \hat{r}, \hat{u}_r = \frac{\hat{p}}{\omega c} \left(\frac{1 + i}{k r} \right)$$

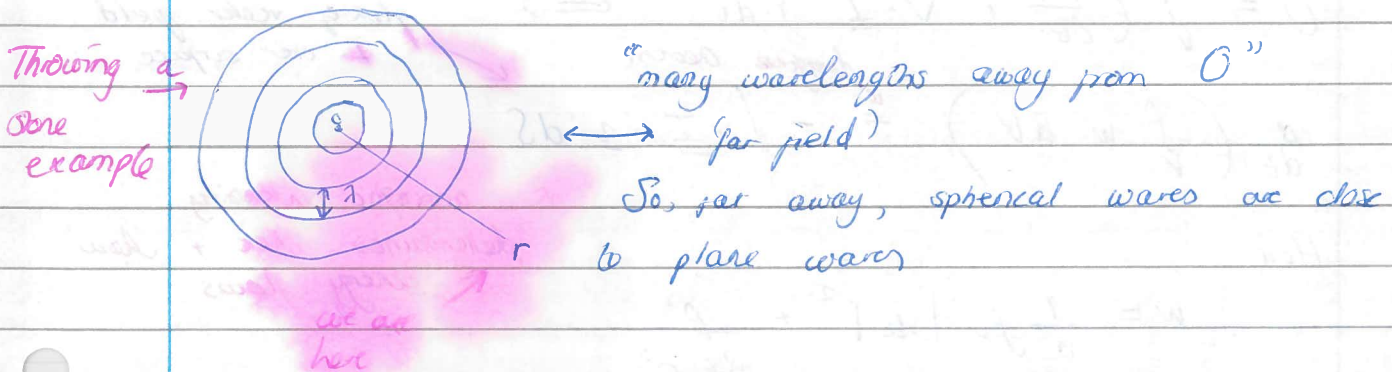
k large, can neglect

unit vector \hat{r} radial direction

(21)

Interpretation: The 1st part of (21) is similar to plane T-H waves, \hat{p} is in-phase with p ; 2nd term out-of-phase with p by $\pi/2$ ($i = e^{i\pi/2}$)

\hookrightarrow $k r$ "large", $k r \gg 1$, the 2nd term is negligible
 $\iff k = 2\pi/\lambda$ (λ wavelength)
 $\iff k r = 2\pi \frac{r}{\lambda} \gg 1 \iff r \gg \lambda$



many wavelengths away from O
 far field
 So, far away, spherical waves are close to plane waves

Acoustic energy, Intensity (Kirchhoff, 1876)

Going back main acoustics equations (7) & (8):

$$(7) \Rightarrow \frac{1}{\rho_0 c^2} \frac{\partial p}{\partial t} + \nabla \cdot \underline{u} = 0 \quad (7')$$

multiply by p

$$(8) \quad \rho_0 \frac{\partial \underline{u}}{\partial t} + \nabla p = 0 \quad (8')$$

dot multiply by \underline{u}

$$(7') \times p + \underline{u} \cdot (8') \Rightarrow$$

$\nabla \cdot (p\underline{u})$

$$\frac{1}{\rho_0 c^2} p \frac{\partial p}{\partial t} + \rho_0 \underline{u} \cdot \frac{\partial \underline{u}}{\partial t} + p \nabla \cdot \underline{u} + \underline{u} \cdot \nabla p = 0$$

$$\frac{\partial}{\partial t} \left(\frac{1}{2} \frac{p^2}{\rho_0 c^2} + \frac{1}{2} \rho_0 \underline{u} \cdot \underline{u} \right) + \nabla \cdot (p\underline{u}) = 0 \quad (22)$$

$\underbrace{\hspace{10em}}_W$
mechanical (energy)

$\underbrace{\hspace{10em}}_I$
intensity

$$\frac{\partial W}{\partial t} + \nabla \cdot \underline{I} = 0$$

Interpretation of 'conservation law' (22):



Integrate (22) over a volume V with boundary S (unit normal \underline{n}):

$$0 = \int_V \left(\frac{\partial W}{\partial t} + \nabla \cdot \underline{I} \right) dV \iff$$

divergence theorem

flux of vector field \underline{I} over surface

$$\frac{d}{dt} \left(\int_V W dV \right) = - \int_S \underline{I} \cdot \underline{n} dS$$

Here

$$W = \underbrace{\frac{1}{2} \rho_0 |\underline{u}|^2}_{\text{kinetic energy density}} + \underbrace{\frac{p^2}{2\rho_0 c^2}}_{\text{potential energy (elastic energy) density}}$$

kinetic energy density

potential energy (elastic energy) density

acoustic intensity determines direction + how energy flows

= Acoustic energy density (total mechanical energy)

Hence $\underline{I} := p \underline{u} :=$ acoustic energy flux, or acoustic intensity

\underline{I} hence represents energy transported in unit time per unit area

For plane T-H waves

$$(18) \Rightarrow \underline{u} = \frac{p}{\rho_0 c} \underline{n} \quad (|\underline{n}| = 1, \text{ direction}) \Rightarrow$$

$$\underline{I} = p \underline{u} = \frac{\rho^2}{\rho_0 c} \underline{n} \quad \text{i.e. the energy flows in direction of propagation } \underline{n}, \text{ as expected}$$

For spherically symmetric T-H waves:

$$p = \operatorname{Re} \left(\frac{A}{r} e^{i(kr - \omega t)} \right) = \operatorname{Re} \left(\frac{|A| e^{i\psi + ikr - i\omega t}}{r} \right)$$

$$= \frac{|A|}{r} \cos(\omega t - kr - \psi) \quad A = |A| e^{i\psi} \text{ complex amplitude;}$$

$$(21) \Rightarrow \underline{u} = \operatorname{Re} \left\{ \hat{u}_r e^{-i\omega t} \right\} \hat{r}$$

$$= \operatorname{Re} \left[\frac{|A|}{r \rho_0 c} e^{i(kr - \omega t + \psi)} \left(1 + \frac{i}{kr} \right) \right] \hat{r}$$

$$\frac{|A|}{r \rho_0 c} \cos(\omega t - kr - \psi) \hat{r} + \frac{|A|}{kr^2 \rho_0 c} \sin(\omega t - kr - \psi) \hat{r}$$

$$\Rightarrow \underline{I} = p \underline{u} = p u_r \hat{r} =: I_r \quad \text{where}$$

$$I_r = p u_r = \frac{|A|^2}{r^2 \rho_0 c} \cos^2(\omega t - kr - \psi)$$

$$+ \frac{|A|^2}{kr^3 \rho_0 c} \cos(\omega t - kr - \psi) \sin(\omega t - kr - \psi)$$

Time harmonic - everything repeats itself i.e. periodic

For time-period average $\langle I_r \rangle := \frac{1}{T} \int_0^T I_r dt$

$\left(T = \frac{2\pi c}{\omega} \right)$; since

$$\int_0^T \cos^2(\omega t - kr - \varphi) dt = \frac{T}{2} \quad (\text{exercise})$$

$$\int_0^T \cos(\omega t - kr - \varphi) \sin(\omega t - kr - \varphi) dt = 0 \quad (\text{check})$$

$$\Rightarrow \langle I_r \rangle = \frac{|A|^2}{\rho_0 c r^2} \cdot \frac{1}{T} \cdot \frac{T}{2}$$

$$\langle I_r \rangle = \frac{|A|^2}{2\rho_0 c r^2} > 0 \quad (23)$$

silly case when $t=0$

The energy flows in radial direction (\vec{E}) away from the origin O , as expected



This finishes the acoustic waves section of chapter 1

1.2 Electromagnetic (EM) waves

The EM fields are described by the Maxwell's equations (ME)
 For electric field $\underline{E}(\underline{x}, t)$, magnetic field $\underline{B}(\underline{x}, t)$, given
 electric charge density ρ and electric current density \underline{j}
 in a uniform electromagnetic medium, the MEs are:

$$\nabla \cdot \underline{E} = \frac{\rho}{\epsilon} \quad (24), \quad \nabla \cdot \underline{B} = 0 \quad (25)$$

Faraday's Law \rightarrow $\nabla \times \underline{E} = -\frac{\partial \underline{B}}{\partial t} \quad (26), \quad \nabla \times \underline{B} = \mu \epsilon \frac{\partial \underline{E}}{\partial t} + \mu \underline{j} \quad (27)$

current density
↓

where ϵ, μ are "electric permittivity" and "magnetic permeability"
 of the medium (i.e. medium's physical characteristics)

Remark (24) relates to Coulomb's Law / Gauss Law

(25) 'Gauss Law' for \underline{B}

(26) = Faraday's Law of EM induction ('-' per Lenz Law)

(27) without $\mu \epsilon \frac{\partial \underline{E}}{\partial t}$ ('Faraday's term') = Ampere's Law



In an electromagnetic medium without "sources" ($\rho = 0, \underline{j} = 0$)

MEs are

$$\nabla \cdot \underline{E} = 0 \quad (28), \quad \nabla \cdot \underline{B} = 0 \quad (29)$$

$$\nabla \times \underline{E} = -\frac{\partial \underline{B}}{\partial t} \quad (30), \quad \nabla \times \underline{B} = \mu \epsilon \frac{\partial \underline{E}}{\partial t} \quad (31)$$

To eliminate \underline{B} , take curl of (30) and use (31)

$$\nabla \times (\nabla \times \underline{E}) = -\frac{\partial}{\partial t} (\nabla \times \underline{B}) = -\mu \epsilon \frac{\partial^2 \underline{E}}{\partial t^2}$$

Also (vector calculus) $\nabla \times \nabla \times \underline{E} = \nabla(\nabla \cdot \underline{E}) - \nabla^2 \underline{E}$

gradient of divergence \rightarrow divergence of gradient = Laplace

$$\underline{a} \times (\underline{b} \times \underline{c}) = (\underline{a} \cdot \underline{c}) \underline{b} - (\underline{a} \cdot \underline{b}) \underline{c}$$

$$\nabla \cdot \underline{E} = 0 \quad \text{by (28)}$$

$$\Rightarrow -\nabla^2 \underline{E} = -\mu \epsilon \frac{\partial^2 \underline{E}}{\partial t^2}$$

$$\frac{1}{c^2} \frac{\partial^2 \underline{E}}{\partial t^2} - \nabla^2 \underline{E} = 0 \quad (32)$$

$c = (\epsilon \mu)^{-1/2}$ wave speed

Similarly (exercise) $\nabla \times$ (31) & (30)

$$\Rightarrow \frac{1}{c^2} \frac{\partial^2 \underline{B}}{\partial t^2} - \nabla^2 \underline{B} = 0 \quad (33)$$

So both \underline{E} and \underline{B} solve (vector) wave equations (32), (33) of (9)/(12)
↑ vector ↑ scalar

For vacuum $c = c_0 = (\epsilon_0 \mu_0)^{-1/2} \approx 3 \times 10^8 \text{ ms}^{-1}$ = speed of light;
 in any other media it can only be smaller, physically

[Exercise (Exam 2010 Q1(a)): Show that for 'full' MEs (24)-(27), \underline{E} solves

$$\frac{1}{c^2} \frac{\partial^2 \underline{E}}{\partial t^2} - \nabla^2 \underline{E} = -\frac{1}{\epsilon} \nabla \rho - \mu \frac{\partial \underline{j}}{\partial t};$$

Derive similar equation for \underline{B}]

Similarly to acoustics (32)-(33) admit T-H solutions:

$$\underline{E}(\underline{x}, t) = \text{Re} \left\{ \hat{\underline{E}}(\underline{x}) e^{-i\omega t} \right\} \quad (34)$$

$$\underline{B}(\underline{x}, t) = \text{Re} \left\{ \hat{\underline{B}}(\underline{x}) e^{-i\omega t} \right\} \quad (35)$$

where $\hat{\underline{E}}(\underline{x}), \hat{\underline{B}}(\underline{x})$ are complex-valued vector field.

As in acoustics, (34) ^{subst.} \rightarrow (32), (35) \rightarrow (33)

yields (vector) Helmholtz equations for $\underline{\hat{E}}, \underline{\hat{B}}$ of (15):

$$\nabla^2 \underline{\hat{E}} + k^2 \underline{\hat{E}} = 0 \quad (36)$$

$$\nabla^2 \underline{\hat{B}} + k^2 \underline{\hat{B}} = 0 \quad (37)$$

$k := \frac{\omega}{c}$ the wavenumber

Plane T-H EM waves

Similarly acoustics, solutions of MEs include plane T-H (EM) waves, in any direction \underline{n} - given \underline{n} , choose x, y, z

Choosing WLOG \underline{n} along z -direction they have the form:

$$\underline{E}(x, t) = \text{Re} \left\{ \underbrace{\underline{E}_0}_{\underline{\hat{E}}} e^{ikz - i\omega t} \right\} \quad (38)$$

$$\underline{B}(x, t) = \text{Re} \left\{ \underbrace{\underline{B}_0}_{\underline{\hat{B}}} e^{ikz - i\omega t} \right\} \quad (39)$$

where $\underline{E}_0, \underline{B}_0$ are complex vector constants

By the above construction, $\forall \underline{E}_0, \underline{B}_0 \in \mathbb{C}^3$, (32) and (33) are satisfied

- 3D space

three complex numbers
↓

Now, (28) $\iff \nabla \cdot \underline{E} = 0$, $\underline{E}_0 = (E_x, E_y, E_z)$

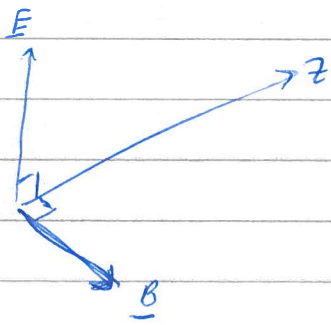
$$\implies \nabla \cdot \underline{E} = \text{Re} \left\{ E_z ik e^{ikz - i\omega t} \right\} = 0 \quad \forall z, \forall t$$

$$\implies E_z = 0$$

Similarly (29) $\iff \nabla \cdot \underline{B} = 0$, $\underline{B}_0 = (B_x, B_y, B_z)$

$$\implies B_z = 0$$

So both \underline{E} and \underline{B} must be perpendicular (\perp) to propagation direction (z) i.e. EM waves are (in this sense) transverse (in contrast to acoustics):



$$\underline{B} = \text{Re} \{ \underline{B}_0 e^{ikz - i\omega t} \}$$

Now plug (38)-(39) in to (30) $\leftrightarrow \nabla \times \underline{E} = -\frac{\partial \underline{B}}{\partial t}$

$$\begin{aligned} \nabla \times \underline{E} &= \text{Re} \left\{ e^{-i\omega t} \begin{vmatrix} \hat{i} & \hat{j} & \hat{k} \\ \frac{\partial}{\partial x} & \frac{\partial}{\partial y} & \frac{\partial}{\partial z} \\ E_x e^{ikz} & E_y e^{ikz} & 0 \end{vmatrix} \right\} \\ &= \text{Re} \left\{ e^{-i\omega t} \left(-i k e^{ikz} E_y, i k e^{ikz} E_x, 0 \right) \right\} \\ &= \text{Re} \left\{ i k e^{ikz - i\omega t} (-E_y, E_x, 0) \right\} \\ &\stackrel{(30)}{=} \text{Re} \left\{ i \omega e^{ikz - i\omega t} \underline{B}_0 \right\} \quad \forall t, \forall z \end{aligned}$$

since nothing depends on x, y

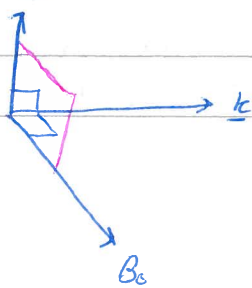
$$\begin{aligned} \Rightarrow \underline{B}_0 &= \frac{k}{\omega} (-E_y, E_x, 0) = \frac{1}{\omega} \begin{vmatrix} \hat{i} & \hat{j} & \hat{k} \\ 0 & 0 & k \\ E_x & E_y & 0 \end{vmatrix} \\ &= \frac{1}{\omega} \underline{k} \times \underline{E}_0, \quad \underline{k} = (0, 0, k) = k \underline{n}, \quad \underline{n} = (0, 0, 1) \end{aligned}$$

different k

\underline{k} is wave vector

$$\text{So } \underline{B}_0 = \frac{1}{\omega} \underline{k} \times \underline{E}_0 \quad (40)$$

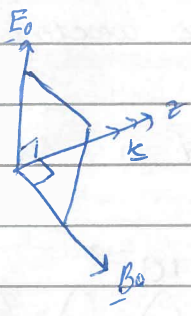
So via (40), $\underline{B}_0 \perp \underline{E}_0$ (also $\underline{B}_0 \perp \underline{k}$, $\underline{E}_0 \perp \underline{k}$)
i.e. $\underline{E}_0, \underline{B}_0, \underline{k}$ form a right-handed orthogonal 1-tuple



02/02/15

Maxwell - 4 eq^{ns}, 2 vector unknowns - eliminate one in favour of other

k wavevector along z



E_0, B_0 complex valued vectors

$$\underline{E}(x,t) = \text{Re} \{ \underline{E}_0 e^{i(kz - \omega t)} \}$$

$$\underline{B}(x,t) = \text{Re} \{ \underline{B}_0 e^{i(kz - \omega t)} \}$$

Exercise: Show that $\underline{E}_0 \perp \underline{k}$ with \underline{B}_0 via (40)

$$\underline{B}_0 = \frac{1}{\omega} \underline{k} \times \underline{E}_0 \quad (40)$$

Solve also (31), & Exam 2014 Q1

So plane EM waves may have different 'polarisations';

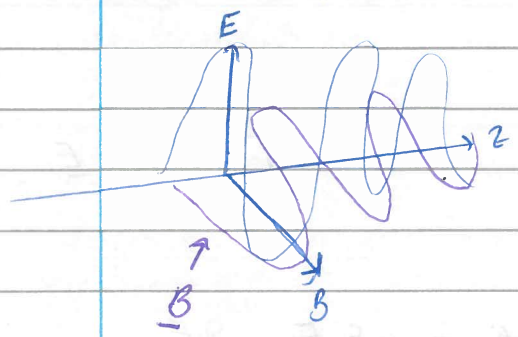
Polarisation plane is the plane containing \underline{E} and \underline{k}

Note that $\underline{E}_0, \underline{B}_0$ are generally complex, so generally $\underline{E}(x,t)$ will rotate; however \forall plane T-M EM wave can be decomposed into sum of two 'polarised' waves:

$$\underline{E}(x,t) = \text{Re} \left\{ \underline{E}_0 e^{i(kz - \omega t)} \right\} = \underbrace{(\text{Re } \underline{E}_0) \cos(\omega t - kz)}_{\underline{E}^{(1)}} + \underbrace{(\text{Im } \underline{E}_0) \sin(\omega t - kz)}_{\underline{E}^{(2)}}$$

sin odd fn. so \leftarrow comes out

where $\underline{E}^{(1)}, \underline{E}^{(2)}$ are polarised



3D picture
sinusoidal shapes
pictures move z direction w/ wave speed c

ω angular frequency

$$k = \frac{\omega}{c} \quad c = (\epsilon \mu)^{-1/2} \quad \lambda = \frac{2\pi}{k}$$

wavelength \leftarrow

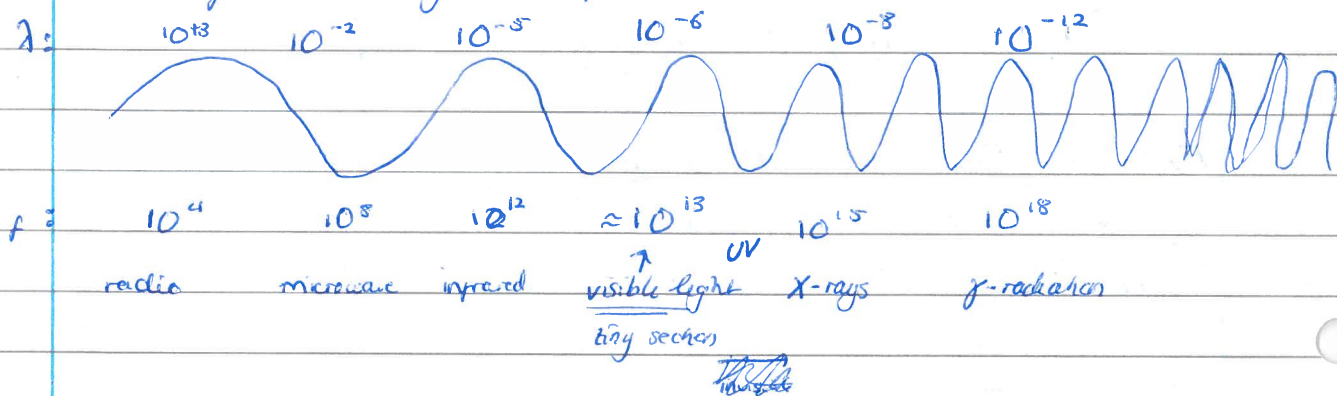
$$T = \frac{2\pi}{\omega} \quad \text{time period (s)}$$

$$f = T^{-1} = \frac{\omega}{2\pi} \quad \text{frequency (Hz = s}^{-1}\text{)}$$

$$\text{frequency} = f = \frac{\omega}{2\pi} = \frac{c}{k} = \frac{c}{2\pi\lambda} = \frac{c}{\lambda}$$

$$\Rightarrow \lambda = \frac{c}{f} \quad \text{higher frequency, shorter wavelength}$$

Depending upon f, λ EM waves are found in different parts of electromagnetic spectrum



Certain things go close to acoustic case - parallels

Electromagnetic energy, Poynting's vector

Similarly to acoustics, see (22), we can derive a conservation law for EM energy from MEs (30) & (31):

$$(30) \iff \nabla \times \underline{E} = -\frac{\partial \underline{B}}{\partial t} \quad \text{Faraday's}$$

$$(31) \iff \nabla \times \underline{B} = \mu \epsilon \frac{\partial \underline{E}}{\partial t}$$

Mult. (30) by \underline{B}

(31) by \underline{E} and subtract

Take dot product of (30) with \underline{B} and of (31) with \underline{E} and subtract \Rightarrow

$$\underline{B} \cdot (\nabla \times \underline{E}) - \underline{E} \cdot (\nabla \times \underline{B}) = -\frac{\underline{B} \cdot \partial \underline{B}}{\partial t} - \mu \epsilon \underline{E} \cdot \frac{\partial \underline{E}}{\partial t}$$

From vector calculus (check):

$$\nabla \cdot (\underline{A} \times \underline{B}) = \underline{B} \cdot (\nabla \times \underline{A}) - \underline{A} \cdot (\nabla \times \underline{B})$$

$$\nabla \cdot (\underline{E} \times \underline{B}) = -\frac{\partial}{\partial t} \left(\frac{1}{2} \underline{B} \cdot \underline{B} + \frac{1}{2} \mu \epsilon \underline{E} \cdot \underline{E} \right)$$

$$\frac{\partial}{\partial t} \left(\frac{1}{2} \epsilon |\underline{E}|^2 + \frac{1}{2\mu} |\underline{B}|^2 \right) + \nabla \cdot \left(\frac{1}{\mu} \underline{E} \times \underline{B} \right) = 0 \quad (41)$$

which is similar to (22)

Here: $w := \frac{1}{2} \epsilon |\underline{E}|^2 + \frac{1}{2\mu} |\underline{B}|^2$ EM field's energy density

$\Rightarrow \underline{S} := \frac{1}{\mu} \underline{E} \times \underline{B}$, called Poynting's vector describes EM flow (per unit time across unit area)

↑
has direction + magnitude shows where energy flows

↑
integrate (41), divergence theorem

So for plane T-H EM waves, e.g. let $E(x,t) = E_x$

$$\underline{E}(x,t) = (E_x, 0, 0) \cos(kz - \omega t) \Rightarrow$$

by (40)

$$\underline{B}(x,t) = \frac{1}{\omega} \underline{k} \times \underline{E}_0$$

$$= \frac{1}{\omega} \begin{vmatrix} i & j & k \\ 0 & 0 & k \\ E_x & 0 & 0 \end{vmatrix} \cos(\omega t - kz)$$

$$= \left(0, \frac{E_x}{c}, 0 \right) \cos(\omega t - kz)$$

$$\Rightarrow \underline{S} = \frac{1}{\mu} \underline{E} \times \underline{B} = \frac{1}{\mu c} \begin{vmatrix} i & j & k \\ E_x & 0 & 0 \\ 0 & E_x & 0 \end{vmatrix} \cos^2(\omega t - kz)$$

$$= \frac{|E_x|^2}{\mu c} (0, 0, 1) \cos^2(\omega t - kz)$$

so the EM energy flows in the z-direction
 \equiv direction of propagation, as expected.

For scattering, need obstacles

We introduce appropriate boundary conditions

1.3 Surface Boundary and Interface conditions

Both in acoustics and in EM, boundary/interface conditions are required e.g. at a solid surface or at an interface between two media (e.g. air and water)

Mathematically, our PDE boundary conditions (BCs) typically are:

- Dirichlet BC: unknown function is specified on the boundary;
- Neumann's BCs: for normal derivative $\frac{\partial p}{\partial n}$ given on $\uparrow S$ boundary

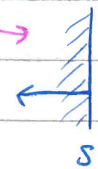
Can also do hybrid of Dirichlet + Neumann

- Mixed/Robin/Impedance BCs, with a combination of p and $\frac{\partial p}{\partial n}$ given on boundary S

1.3.1 Acoustic BCs

(i) Acoustically hard ~~BC~~ surface doesn't yield to acoustic wave \leftrightarrow is rigid \leftrightarrow

can move along surface \downarrow
but not through \leftrightarrow



$\underline{u} \cdot \underline{n} = 0$ (\Leftrightarrow no normal component in velocity \underline{u} on S); assuming irrotational fields

$$\underline{u} = \nabla \phi \Rightarrow \underline{u} \cdot \underline{n} = \nabla \phi \cdot \underline{n} = \frac{\partial \phi}{\partial n} = 0$$

directional derivative

\Rightarrow Zero Neumann BC for ϕ

(ii) Acoustically soft surface freely yields to a wave (e.g. water surface) \rightarrow Physically on S , $p = 0$, (on S , pressure = atmospheric pressure) \leftrightarrow Dirichlet - BCs

(iii) Impedance BCs are intermediate between (i) and (ii): a surface impedance, Z , measures to what extent the surface resists to the applied pressure; it generally depends on frequency (T-H case $\leftrightarrow \omega$); so for T-H acoustics, $p = \text{Re} \{ \hat{p} e^{-i\omega t} \}$, $\underline{u} = \text{Re} \{ \hat{\underline{u}} e^{-i\omega t} \}$

$\hat{u}_n := \hat{u} \cdot \hat{n}$ *higher resistance, more pressure needed to apply*

$$z := \frac{\hat{p}}{\hat{u}_n} \iff \hat{p} - z \hat{u}_n = 0 \iff \hat{p} - z \frac{\partial \hat{\phi}}{\partial n} = 0 \quad *$$

$$\left(\phi = \text{Re} \left\{ \hat{\phi}(x) e^{-i\omega t} \right\} \right)$$

$$(21) \iff \frac{\partial \phi}{\partial t} = -\frac{1}{\rho_0} p \stackrel{T-H}{\iff} -i\omega \hat{\phi} e^{-i\omega t} = -\frac{1}{\rho_0} \hat{p} e^{-i\omega t}$$

$$\implies \boxed{\hat{p} = i\omega \rho_0 \hat{\phi}} \quad (42)$$

$$\implies i\omega \rho_0 \hat{\phi} - z \frac{\partial \hat{\phi}}{\partial n} = 0 \quad (43)$$

which is the impedance BC.

Also, via (42),

$$i\omega \rho_0 \hat{p} - z(\omega) \frac{\partial \hat{p}}{\partial n} = 0 \quad (43')$$

Notice: if $z \rightarrow 0 \implies \hat{p} = \hat{\phi} = 0$ (soft BC (ii)); *

if $z \rightarrow \infty \iff \frac{1}{z} \rightarrow 0 \implies \frac{\partial \hat{\phi}}{\partial n} = \frac{\partial \hat{p}}{\partial n} = 0$ (hard BC (i))

Generally $z(\omega)$ is complex, physically,

$\text{Re } z \geq 0$, with $\text{Re } z > 0$ for energy absorbing surfaces

$\text{Re } z = 0$ for energetically neutral surfaces

(will see in §2)

1-3-2 EM interface conditions / BCs

② ϵ_2, μ_2 \uparrow \underline{n} $\underline{E}^{(2)}, \underline{B}^{(2)}$ At an interface between two EM media (ϵ_1, μ_1 in medium 1, ϵ_2, μ_2 in medium 2), the EM fields $\underline{E}^{(1)}, \underline{B}^{(1)}$ & $\underline{E}^{(2)}, \underline{B}^{(2)}$ respectively, are physically required to obey the following interface conditions:

i) Tangential components of \underline{E} are continuous (cts) \Leftrightarrow

$$\underline{E}^{(1)} \times \underline{n} = \underline{E}^{(2)} \times \underline{n} \quad \begin{array}{l} \text{projections on interface plane} \\ \text{are the same} \end{array} \quad (44)$$

\uparrow
disregards normal component

ii) Normal components of \underline{B} are cts \Leftrightarrow

$$\underline{B}^{(1)} \cdot \underline{n} = \underline{B}^{(2)} \cdot \underline{n} \quad (45)$$

Also, if there exist a surface charge distribution ρ_s and/or surface current distribution \underline{j}_s , then (iii)

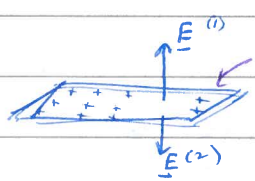
(iii) $\epsilon_2 \underline{E}^{(2)} \cdot \underline{n} - \epsilon_1 \underline{E}^{(1)} \cdot \underline{n} = \rho_s$ (46)

electric permittivity \rightarrow

(iv) $\frac{1}{\mu_2} \underline{B}^{(1)} \times \underline{n} - \frac{1}{\mu_1} \underline{B}^{(2)} \times \underline{n} = \underline{j}_s$ (47)

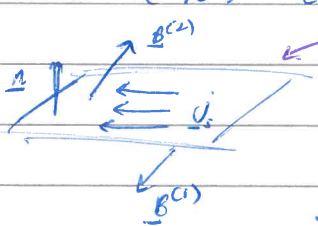
magnetic permeability $\rightarrow \mu$

(In (46), (47) the unit normal \underline{n} points in to medium 2)

Check:  rvely charged so \underline{E} goes away from plane

$$\rho_s > 0 \Rightarrow \underline{E}^{(2)} \cdot \underline{n} > 0, \quad \underline{E}^{(1)} \cdot \underline{n} < 0$$

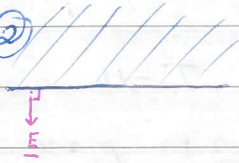
\Rightarrow (46) OK

 \leftarrow current going right to left

"right hand rule" for $\underline{B}^{(1)}, \underline{B}^{(2)}$, due to $\underline{j}_s \Rightarrow$ signs in (47) OK

Perfectly conducting BCs

②



If medium 2 is a "perfect conductor"

$$\Rightarrow \vec{E}^{(2)} = 0$$

cannot support electric field

\Rightarrow from (44) $\vec{E}^{(1)} \times \hat{n} = 0$ i.e. \vec{E} must have a zero tangential component. (End of chapter 1)

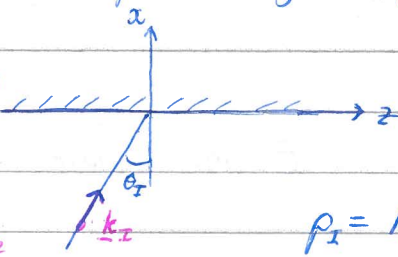
2. Canonical cases

2. 'Canonical' Cases (Reflection & Refraction)

2.1 Reflection by a plane (acoustics)

y is perp.
to blackboard

propagating
towards surface



Consider a plane T-H acoustic

wave incident upon a plane boundary

$$S' = \{(x, y, z) : x = 0\}$$

$$p_I = A_I e^{-i\omega t + i\mathbf{k}_I \cdot \mathbf{x}} \quad (\text{dropping "Re"})$$

Assume it has 'angle of incidence' θ_I

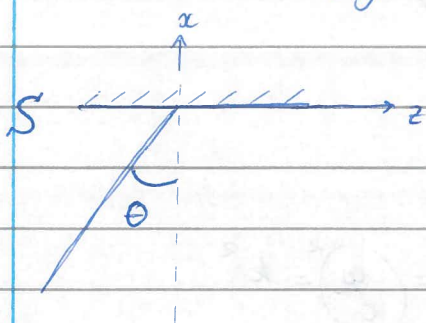
$\mathbf{k}_I = k(\cos\theta_I, 0, \sin\theta_I)$ wavevector;

$k := |\mathbf{k}_I| = \frac{\omega}{c}$ the wavenumber

$0 \leq \theta_I < \frac{\pi}{2}$ (moving 'towards' S')

09/02/15

Canonical - standard, generic



normally write p_R , but have dropped it

$$p_I = A_I e^{i k_I \cdot x - i \omega t}$$

$$k_I = k (\cos \theta_I, 0, \sin \theta_I)$$

expect surface bounces wave back

The boundary S will "reflect" the incident wave p_I

What is the reflected wave p_R ? Answer depends on BC

This depends on the boundary conditions on S

Consider acoustically soft BC: $p=0$ on S (§1.3.1)

$$\Leftrightarrow p = p_I + p_R = 0 \text{ on } S$$

Seek p_R also as a plane T-H wave:

$$p_R = A_R e^{i k_R \cdot x - i \omega_R t}$$

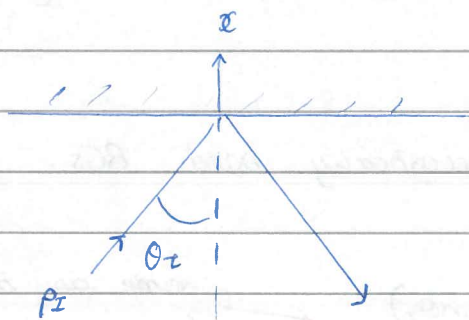
will see if T-H can only be

$$k_R = (k_x, k_y, k_z)$$

with same frequency but let's assume different

$$|k_R| =: k_R = \frac{\omega_R}{c} \rightarrow \text{general rule for T-H waves}$$

\leftarrow wave speed, c .



$k_x < 0$ (moving 'away' from S)

We have incident wave moving towards, reflected wave moving away

$$\text{BCs: } x \in S \Leftrightarrow x=0, \underline{x} = (0, y, z)$$

$$p_I + p_R = 0 \Leftrightarrow p_R = -p_I \text{ for } x=0 \Leftrightarrow$$

$$A_R e^{i k_y y + i k_z z - i \omega_R t} = -A_I e^{i k_x \sin \theta_I z - i \omega t}$$

$\forall t, k_y, k_z$

\rightarrow We assume $A_I \neq 0$

- e Varying $t \implies \omega_R = \omega$
- o " $y \implies k_y = 0 \rightarrow$ since y on LHS, not on RHS
- o " $z \implies k_z = k \sin \theta_I$

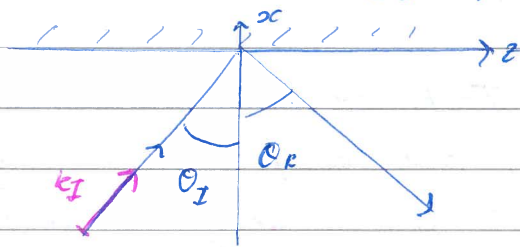
Also $A_R = -A_I$

Now $k_R^2 = k_x^2 + k_y^2 + k_z^2 = \left(\frac{\omega_R}{c}\right)^2 = \left(\frac{\omega}{c}\right)^2 = k^2$

$k_x = \pm (k^2 - k_z^2)^{1/2} = \pm \sqrt{k^2 - k^2 \sin^2 \theta_I}$

~~$k_x = \pm \dots$~~ $= -k \cos \theta_I$ (since $k_x < 0$)

So $k_R = k (-\cos \theta_I, 0, \sin \theta_I)$



Interpretation: $\theta_R = \theta_I$ ('Specular reflection law')

$A_R = -A_I \iff$ Phase shift for TE upon reflection ($-1 = e^{i\pi}$)

Exercises: 1) Show that for acoustically hard BCs

$\left(\frac{\partial \phi}{\partial n} = 0 \iff \frac{\partial p}{\partial n} = 0 \right)$

$A_R = A_I e^{-i\omega t + ik((-x \cos \theta_I) + z \sin \theta_I)}$

← same as before but no minus sign

Exam 2012 Q2C6)

2) Find the reflected wave for a general impedance

BC (1.42) ← ch. 1, formula 4.2

~~$e^{-i\omega t}$~~

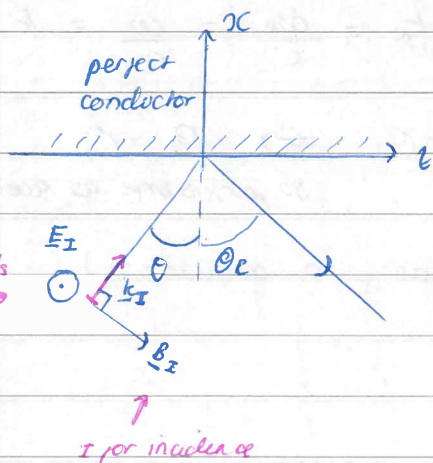
$$i\omega\epsilon_0\rho - z\frac{\partial\rho}{\partial z} = 0$$

Conclude that $|A_R| \leq |A_I|$
 $\iff \operatorname{Re} z > 0$

This is a natural assumption
 Reflected signal could only weaken

(Exam 2012 Q2(c))

2.2 Reflection of plane T-H EM waves by perfectly conducting planes



Consider a plane T-H EM wave propagating in an "insulator" medium (no currents) occupying half-space $x < 0$; with angle of incidence θ , unknown angle of reflection θ_R ; ϵ, μ given
 $\implies c = (\epsilon\mu)^{-1/2}$ wave speed

Let the incidence wave be polarised so that \underline{E}_I is parallel to y-axis, pointing out of the plane:

$$\underline{E}_I = (0, E_I, 0) e^{-i\omega t + ik(\cos\theta z + z\sin\theta)} \quad (1)$$

$$\underline{B}_I \stackrel{(1.40)}{=} \frac{1}{\omega} \underline{k}_I \times \underline{E}_I = \frac{E_I}{c} (-\sin\theta, 0, \cos\theta) e^{-i\omega t + ik(\cos\theta z + z\sin\theta)} \quad (2)$$

$$\underline{k}_I = k(\cos\theta, 0, \sin\theta)$$

The reflected wave is also sought as a plane wave

$$\underline{E}_R = (E_x, E_y, E_z) e^{-i\omega_R t + ik_R(-x\cos\theta_R + z\sin\theta_R)} \quad (3)$$

$$\underline{B}_R = (B_x, B_y, B_z) e^{-i\omega_R t + ik_R(-x\cos\theta_R + z\sin\theta_R)} \quad (4)$$

$$k_e = \omega_e/c ; E_x, E_y, E_z, B_x, B_y, B_z, \omega_e, \theta_e \text{ to be found}$$

From perfectly conducting BCs (1.48) on

$$S = \{x=0\}, \quad \underline{E} \times \underline{n} = 0 \quad \underline{n} = (1, 0, 0)$$

$$\underline{E} = \underline{E}_I + \underline{E}_e = \text{total electric field}$$

\underline{E} has zero y and z components on S ($x=0$)

y component:

$$E_y e^{-i\omega_e t + ik_e z \sin \theta_e} = -E_I e^{-i\omega_e t + ik_e z \sin \theta}$$

$$\Rightarrow \omega_e = \omega \text{ (varying } t); \Rightarrow k_e = \frac{\omega_e}{c} = \frac{\omega}{c} = k$$

$$\Rightarrow \text{(varying } z) \quad \overset{\leftarrow \text{since } k_e = k}{k_e \sin \theta_e} = k \sin \theta \Rightarrow \theta_e = \theta$$

so far, same as acoustics

(i.e. specular reflection law, same as in acoustics)

$$\text{Also } E_y = -E_I$$

$$\underline{z}\text{-components: } \Rightarrow E_z = 0;$$

Also, since $(\underline{E}_e, \underline{B}_e)$ is a plane EM wave

$$\Rightarrow \underline{E}_e \perp \underline{k}_e \iff \underline{E}_e \cdot \underline{k}_e = 0 \iff$$

$$k(E_{0x} \cos \theta_e + \cancel{E_z} \sin \theta_e) = 0 \Rightarrow \underline{E_x} = 0$$

So $\underline{E}_e = (0, -E_I, 0) e^{-i\omega t + ik_e z}$ and so points into the plane;

Finally, from (1.40)

$$\underline{B}_e = \frac{1}{\omega} \underline{k}_e \times \underline{E}_e = \frac{k}{\omega} \begin{vmatrix} \hat{i} & \hat{j} & \hat{k} \\ -\cos \theta & 0 & \sin \theta \\ 0 & -E_I & 0 \end{vmatrix} e^{-i\omega t + ik_e z} \quad \theta_e = \theta$$

$$\frac{E_I}{c} (\sin\theta, 0, \cos\theta) e^{-i\omega t + ikz \cdot x}$$

□

What else can we say about this problem

PC charges can accumulate on boundary
 nothing above since perfect conductor (PC)

Finally, let's find ^{induced} electric current on S' :

Notice that $\underline{E} \equiv 0$ in $x > 0$ (in the perfect conductor)
 $\Rightarrow \underline{B} \equiv 0$ in $x > 0$ from ME (1.26)

$$\frac{\partial B}{\partial t} = 0 \Rightarrow \underline{B} = 0 \quad (\text{constants ruled out since in T-H case})$$

Now from interface condition (1.47)

$$\uparrow \quad \underline{B}^{(2)} \equiv 0$$

$$\frac{1}{\mu} \underline{B}^{(1)} \times \underline{n} - \frac{1}{\mu_2} \underline{B}^{(2)} \times \underline{n} = \underline{j}_s$$

$$\Rightarrow \underline{j}_s = \frac{1}{\mu} (\underline{B}_I + \underline{B}_R) \times (1, 0, 0):$$

$$\underline{B}_I \times \underline{n} = \frac{E_I}{c} e^{-i\omega t + ikz \cdot x} \begin{vmatrix} \hat{i} & \hat{j} & \hat{k} \\ -\sin\theta & 0 & \cos\theta \\ 1 & 0 & 0 \end{vmatrix}$$

$$(\underline{B}_I + \underline{B}_R) \times \underline{n} = \frac{E_I}{c} e^{-i\omega t + ikz \cdot x} \begin{vmatrix} \hat{i} & \hat{j} & \hat{k} \\ 0 & 0 & 2\cos\theta \\ 1 & 0 & 0 \end{vmatrix}$$

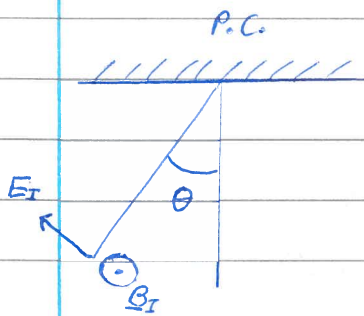
$$= \frac{E_I}{c} 2\cos\theta (0, 1, 0) e^{-i\omega t + ikz \cdot x}$$

$$\underline{j}_s = \frac{2\cos\theta E_I}{\mu c} e^{-i\omega t + ikz \cdot x} (0, 1, 0)$$

is the induced current

□

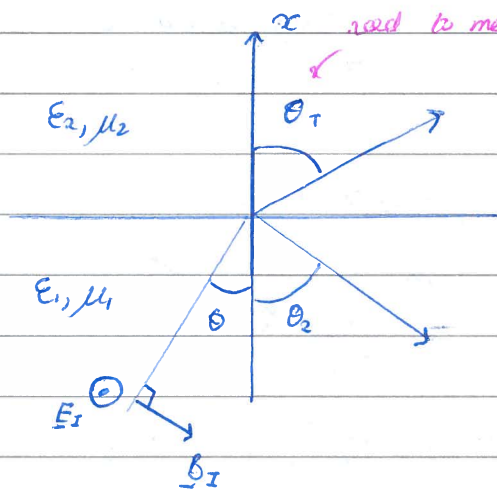
Exercise



Consider a differently polarized incident wave: \underline{E}_I in-plane
 \underline{B}_I out-of-plane

(Exam 2011 Q2)

2.3 Reflection and Refraction by a plane interface between two insulator media



x used to measure everything relative to normal

$\theta_T \leftarrow$ transmission

Consider two insulator EM media with interface

$$S = \{x=0\};$$

medium 1: ϵ_1, μ_1

— " — 2: ϵ_2, μ_2

A plane EM wave incident from medium 1 same as in §2.2 i.e. described by ① and ②

$$\text{①} \Rightarrow \underline{E}_I = E_I (0, 1, 0) e^{-i\omega t + ik(x\cos\theta + z\sin\theta)}$$

$$\text{②} \Rightarrow \underline{B}_I = \frac{E_I}{c} (-\sin\theta, 0, \cos\theta) e^{-i\omega t + ik(x\cos\theta + z\sin\theta)}$$

Seek the reflected wave also as in ③ & ④

$$\text{③} \Rightarrow \underline{E}_R = (E_x, E_y, E_z) e^{-i\omega_R t + ik_R(-x\cos\theta_R + z\sin\theta_R)}$$

$$\text{④} \Rightarrow \underline{B}_R = (B_x, B_y, B_z) e^{-i\omega_R t + ik_R(-x\cos\theta_R + z\sin\theta_R)}$$

with different unknowns $E_x, E_y, E_z, \dots, B_z$

Additionally a 'transmitted' wave is sought in $x > 0$ (medium 2):

$$\underline{E}_T = (\hat{E}_x, \hat{E}_y, \hat{E}_z) e^{-i\omega_T t + ik_T(x \cos \theta_T + z \sin \theta_T)} \quad (5)$$

$$\underline{B}_T = (\hat{B}_x, \hat{B}_y, \hat{B}_z) e^{-i\omega_T t + ik_T(x \cos \theta_T + z \sin \theta_T)} \quad (6)$$

$$\left(k_R = \frac{\omega_R}{c_1}, \quad k_T = \frac{\omega_T}{c_2}, \quad \epsilon_1 = (\epsilon_1 \mu_1)^{-1/2}, \quad \epsilon_2 = (\epsilon_2 \mu_2)^{-1/2} \right)$$

From (1.44) $\Leftrightarrow \underline{E}^{(1)} \times \underline{n} = \underline{E}^{(2)} \times \underline{n} \Leftrightarrow$
 continuity of y and z components of \underline{E} ($x=0$)

\Rightarrow

y-components:

$$E_y e^{-i\omega t + ikz \sin \theta} + E_y e^{-i\omega_T t + ik_T z \sin \theta_T} = \hat{E}_y e^{-i\omega_T t + ik_T z \sin \theta_T} \quad (\forall t, z)$$

\Rightarrow (q § 2.2)

$$\omega_R = \omega_T = \omega \Rightarrow k_R = k$$

$$\Rightarrow (\text{varying } z) \quad k \sin \theta = k_T \sin \theta_T = k_T \sin \theta_T$$

$$\theta_R = \theta$$

(as before: specular reflection) and additionally

$$k_T \sin \theta_T = k \sin \theta \Leftrightarrow \frac{\omega}{c_2} \sin \theta_T = \frac{\omega}{c_1} \sin \theta$$

$\frac{\sin \theta_T}{c_2} = \frac{\sin \theta}{c_1}$	<i>will write in equivalent form using refractive index</i>	(7)
---	---	-----

Let $n := \frac{c_0}{c}$, c_0 speed of light (in vacuum)

called refractive index ($\Rightarrow n > 1, n=1$ in vacuum)

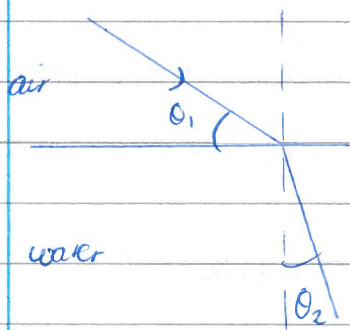
\Rightarrow

$$n_2 \sin \theta_2 = n_1 \sin \theta_1 \quad (n \sin \theta = \text{constant})$$

"
 θ_T

$$(8)$$

(7) or (8) is Snell's law of refraction



wave speed
smaller in water
← chosen this angle so it can go faster

Further (from equating the y-components)

$$E_z + E_y = \hat{E}_y \quad (9)$$

z-components:

$$E_z = \hat{E}_z \quad (10)$$

$$\text{Also, } \underline{E}_z \cdot \underline{k}_R = 0, \quad \underline{E}_T \cdot \underline{k}_T = 0$$

$$\Rightarrow -E_x \cos \theta + E_z \sin \theta = 0 \quad (11)$$

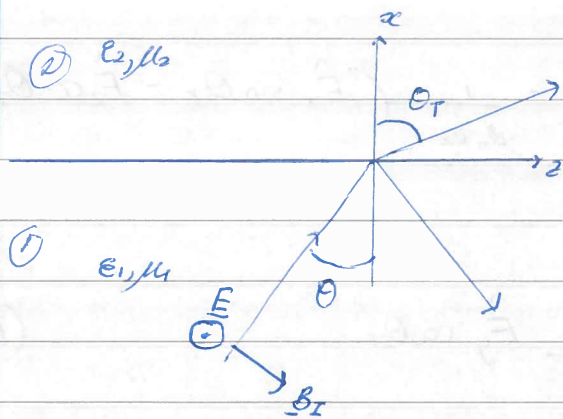
$$\hat{E}_x \cos \theta_T + \hat{E}_z \sin \theta_T = 0 \quad (12)$$

Further from (1.47) and with no surface currents (since insulators - cannot support currents)

$$\frac{1}{\mu_1} (\underline{B}_I + \underline{B}_R) \times \underline{n} = \frac{1}{\mu_2} \underline{B}_T \times \underline{n} \quad \Leftrightarrow$$

Continuity of y & z components of $\frac{B}{\mu}$

23/02/15



Could have boundary w/
appropriate bdy conditions or
could have interface

Snell's Law of Refraction: $\frac{\sin \theta}{c_1} = \frac{\sin \theta_T}{c_2}$ (7)

$$\underline{E}_i = (E_x, E_y, E_z) e^{i(kx - \omega t)}$$

$$\underline{E}_T = (\hat{E}_x, \hat{E}_y, \hat{E}_z) e^{i(k_T x - \omega t)}$$

$$\underline{E}_R = (0, E_R, 0) e^{i(kx - \omega t)}$$

$$E_z + E_R = \hat{E}_y \quad (9)$$

(10, 11, 12)

y, z components of $\frac{1}{\mu} \underline{B}$ continuous (*)

$$(1.40) \Rightarrow \underline{B}_i = \frac{1}{\omega} \underline{k}_i \times \underline{E}_i$$

distinguishing between k_s

$$= \frac{k}{\omega} \begin{vmatrix} \hat{i} & \hat{j} & \hat{k} \\ -\cos \theta & 0 & \sin \theta \\ E_x & E_y & E_z \end{vmatrix} e^{-i\omega t + ikx}$$

$$= \frac{1}{\omega} (-E_y \sin \theta, E_z \cos \theta + E_x \sin \theta, -E_x \cos \theta) e^{i(kx - \omega t)}$$

$$\underline{B}_T = \frac{1}{\omega} \underline{k}_T \times \underline{E}_T = \frac{k_T}{\omega} \begin{vmatrix} \hat{i} & \hat{j} & \hat{k} \\ \cos \theta_T & 0 & \sin \theta_T \\ \hat{E}_x & \hat{E}_y & \hat{E}_z \end{vmatrix} e^{-i\omega t + ik_T x}$$

$$= \frac{1}{\omega} (-\hat{E}_y \sin \theta_T, \hat{E}_x \sin \theta_T - \hat{E}_z \cos \theta_T, \hat{E}_x \cos \theta_T) e^{i(k_T x - \omega t)}$$

y component of (*):

$$\frac{1}{\mu_1 c_1} (E_z \cos \theta + E_x \sin \theta) = \frac{1}{\mu_2 c_2} (\hat{E}_x \sin \theta_T - \hat{E}_z \cos \theta_T) \quad (13)$$

z-comp of (*):

$$\frac{1}{\mu_1 c_1} \left(\frac{E_z - E_y}{\cos \theta} \right) = \frac{1}{\mu_2 c_2} \hat{E}_y \cos \theta_T \quad (14)$$

So (9)-(14) are six equations for six unknowns
 $E_x, E_y, E_z, \hat{E}_y, \hat{E}_z, \hat{E}_x$

Notice (10)-(13) equations for
 $E_{oc}, E_z, \hat{E}_x, \hat{E}_z$ only;

$$\begin{aligned} (13) \Rightarrow & \frac{1}{\mu_1 c_1} \left(E_z \cos \theta + \overbrace{E_z \sin \theta}^{E_x \text{ by (11)}} \sin \theta \right) \\ & = \frac{1}{\mu_2 c_2} \left(- \underbrace{\hat{E}_z \sin \theta_T}_{\hat{E}_{oc} \text{ by (12)}} \sin \theta_T - \underbrace{E_z \cos \theta_T}_{\hat{E}_z \text{ by (10)}} \right) \end{aligned}$$

$$\frac{E_z}{\mu_1 c_1 \cos \theta} = - \frac{E_z}{\mu_2 c_2 \cos \theta_T} \Rightarrow E_z = 0$$

$$\left(\frac{1}{\mu_1 c_1 \cos \theta} + \frac{1}{\mu_2 c_2 \cos \theta_T} > 0 \right)$$

$$\Rightarrow \boxed{0 = E_z \stackrel{(10)}{=} \hat{E}_z \stackrel{(12)}{=} \hat{E}_x \stackrel{(11)}{=} E_{oc}}$$

So we have left (9), (14) for E_y, \hat{E}_y

Remark: So like the incident field \underline{E}_I , both reflected and transmitted electric fields have only y-component \iff "transverse electric" \iff TE wave

$$(9) \rightarrow E_I + E_y = \hat{E}_y$$

$$(14) \rightarrow \frac{1}{\mu_1 c_1} (E_I - E_y) \cos \theta = \frac{1}{\mu_2 c_2} \hat{E}_y \cos \theta_T$$

$$(9) \rightarrow (14) : \underbrace{\frac{\cos \theta}{\mu_1 c_1}}_{=: m_1} (E_I - E_y) = \underbrace{\frac{\cos \theta_T}{\mu_2 c_2}}_{=: m_2} (E_I + E_y)$$

$$m_1 E_I - m_1 E_y = m_2 E_I + m_2 E_y \quad \Rightarrow$$

$E_y = \frac{m_1 - m_2}{m_1 + m_2} E_I$
$\hat{E}_y = E_y + E_I = \frac{2m_1}{m_1 + m_2} E_I$

$$m_1 := \frac{\cos \theta}{\mu_1 c_1}, \quad m_2 := \frac{\cos \theta_T}{\mu_2 c_2}$$

If both media 1 and 2 are "non-magnetic", i.e.
 $\mu_1 = \mu_2 = \mu$ (as often the case physically) but $\epsilon_1 \neq \epsilon_2$
 (so $c_1 \neq c_2$, $c = (\epsilon \mu)^{-1/2}$)

$$\Rightarrow n_1 = \frac{c_0}{c_1}, \quad n_2 = \frac{c_0}{c_2}, \quad \text{refractive indices}$$

$$E_y = \frac{n_1 \cos \theta - n_2 \cos \theta_T}{n_1 \cos \theta + n_2 \cos \theta_T} E_I \quad (15)$$

$$\hat{E}_y = \frac{2n_1 \cos \theta}{n_1 \cos \theta + n_2 \cos \theta_T} E_I \quad (16)$$

Now via Snell's law (8), $n_1 \sin \theta = n_2 \sin \theta_T = D$

$$E_y = \frac{\frac{D}{\sin \theta} \cos \theta - \frac{D}{\sin \theta_T} \cos \theta_T}{\frac{D}{\sin \theta} \cos \theta + \frac{D}{\sin \theta_T} \cos \theta_T} E_I$$

$$= \frac{\sin \theta_T \cos \theta - \sin \theta \cos \theta_T}{\sin \theta_T \cos \theta + \sin \theta \cos \theta_T} E_T$$

$$E_y = \frac{\sin(\theta_T - \theta)}{\sin(\theta + \theta_T)} E_T \quad (17)$$

$$\hat{E}_y = \frac{2 \sin \theta_T \cos \theta}{\sin(\theta + \theta_T)} E_T \quad (18)$$

similarly

After $\underline{E}_R, \underline{E}_T$ found \Rightarrow by (1.40)

$$\underline{B}_R = \frac{1}{\omega} \underline{k}_R \times \underline{E}_R, \quad \underline{B}_T = \frac{1}{\omega} \underline{k}_T \times \underline{E}_T$$



Total internal reflection

$$\text{Snell's law (7)} \iff \frac{\sin \theta_T}{c_2} = \frac{\sin \theta}{c_1} \implies$$

$$\sin \theta_T = \frac{c_2}{c_1} \sin \theta = \frac{n_1}{n_2} \sin \theta$$

\leftarrow notice reverse $\frac{c_2}{c_1} \rightarrow \frac{n_1}{n_2}$

Let $c_2 > c_1$ and $\frac{c_2}{c_1} \sin \theta > 1 \iff \theta > \sin^{-1}\left(\frac{c_1}{c_2}\right) =: \theta_c$
 = "critical angle"

$\theta > \theta_c$ is when total internal reflection occurs:

All the above derived formulae still hold, however θ_T is now complex:

$$\sin \theta_T = \frac{c_2}{c_1} \sin \theta =: \alpha > 1 \implies$$

$$\cos \theta_T = \left(1 - \sin^2 \theta_T\right)^{1/2} = \left(1 - \alpha^2\right)^{1/2} = i\sqrt{\alpha^2 - 1} =: i\beta$$

$\beta > 0$

\implies from (15)

$$E_y = \frac{n_1 \cos \theta - n_2 \cos \theta_T}{n_1 \cos \theta + n_2 \cos \theta_T} E_T = \frac{n_1 \cos \theta - i\beta n_2}{n_1 \cos \theta + i\beta n_2} E_T$$

\uparrow
complex conjugates!

so same modulus



$$|E_y| = |E_z| \quad (\text{since } n_1 \cos \theta \pm i\beta n_2 \text{ complex conjugates})$$

↔ reflected & incident amplitudes are the same

⇒ all energy reflects (although phases may change upon reflection)

[For $\theta < \theta_c$ (15) ⇒ $|E_y| < |E_z|$ ↔ reflected energy less than incident, with the difference going to transmitted energy]

above interface

For $\theta > \theta_c$, for $x > 0$ (in medium 2):

$$\underline{k}_T = k_T (\cos \theta_T, 0, \sin \theta_T) \\ = k_T (i\beta, 0, \alpha)$$

$$\underline{E}_T = \text{Re} \left\{ (0, \hat{E}_y, 0) e^{i k_T \cdot \underline{x} - i\omega t} \right\}$$

$$= \text{Re} \left\{ (0, \hat{E}_y, 0) e^{-k_T \beta x + i k_T \alpha z - i\omega t} \right\}$$

$$= (0, |\hat{E}_y| e^{i\psi}, 0) e^{-k_T \beta x} \cos(\omega t - k_T \alpha z - \psi), 0$$

exp. decay in x ϕ

so \underline{E}_T exponentially decays as $x \rightarrow +\infty$;

$$(1.40) \Rightarrow \underline{B}_T = \frac{1}{\omega} \underline{k}_T \times \underline{E}_T$$

$$= \text{Re} \left\{ \frac{k_T}{\omega} \begin{vmatrix} \hat{i} & \hat{j} & \hat{k} \\ i\beta & 0 & \alpha \\ 0 & \hat{E}_y & 0 \end{vmatrix} e^{-k_T \beta x + i k_T \alpha z - i\omega t} \right\}$$

$$= \frac{e^{-k_T \beta x}}{\omega} \left(-\alpha |\hat{E}_y| \cos(\omega t - k_T \alpha z - \psi), 0, \beta |\hat{E}_y| \sin \phi \right)$$

ϕ

Now for Poynting vector \underline{S}_T (Eq 1-2) for transmitted wave

$$\underline{S}_T = \frac{1}{\mu} \underline{E}_T \times \underline{B}_T = \frac{|\hat{E}_y|^2 e^{-2k_T \beta z}}{\mu c_2} \begin{vmatrix} \hat{i} & \hat{j} & \hat{k} \\ 0 & \cos \phi & 0 \\ -\alpha \cos \phi & 0 & \beta \sin \phi \end{vmatrix}$$

($\phi := \omega t - k_T x - \psi$)

$$= \frac{|\hat{E}_y|^2 e^{-2k_T \beta z}}{\mu c_2} (\beta \sin \phi \cos \phi, 0, \alpha \cos^2 \phi)$$

For time averaged $\langle \underline{S}_T \rangle = \frac{1}{T} \int_0^T \underline{S}_T(\underline{x}, t) dt$

different π to S_T \rightarrow

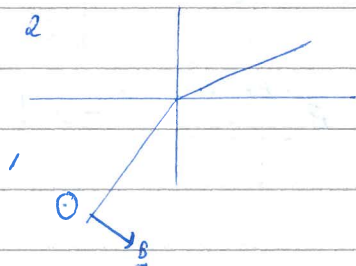
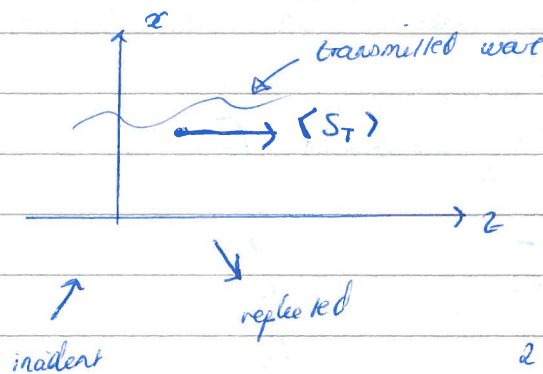
$$\langle \underline{S}_T \rangle = \frac{|\hat{E}_y|^2 e^{-2k_T \beta z}}{\mu c_2} (0, 0, \frac{\alpha}{2})$$

$$= \frac{\alpha |\hat{E}_y|^2 e^{-2k_T \beta z}}{2 \mu c_2} (0, 0, 1)$$

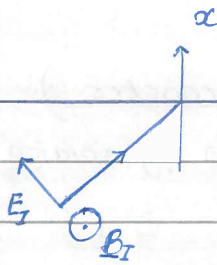
unit vector in z direction

(Since $\langle \sin \phi \cos \phi \rangle = 0$, $\langle \cos^2 \phi \rangle = \frac{1}{2}$)

So the (averaged) energy flows z-direction i.e. along the boundary (no energy away from boundary)



Exercise: Similar analysis can be done for incident plane wave polarised in (x, z) -plane



→ z ← "transverse magnetic" ↔ TM
(cf TE case above)

All similar conclusions e.g. $\theta_R = \theta$,
Snell's law (7), total internal reflection
for $\theta > \theta_c$ if $c_2 > c_1$

Difference: (17) - (18) replaced by

$$B_y = \frac{\sin 2\theta - \sin 2\theta_T}{\sin 2\theta + \sin 2\theta_T} B_I$$

$$\hat{B}_y = \frac{2 \sin 2\theta}{\sin 2\theta + \sin 2\theta_T} B_I$$

For $\theta = \theta_B$ (the "Brewster angle") $B_y = 0 \iff$
 $\sin 2\theta = \sin 2\theta_T \implies (\theta_T \neq \theta) \implies 2\theta_T = \pi - 2\theta$
 $\iff \theta_T = \frac{\pi}{2} - \theta \implies$ (using (7))
 $\theta_B = \tan^{-1} \left(\frac{c_1}{c_2} \right)$

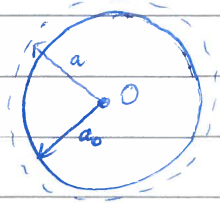
So for $\theta = \theta_B$ - No reflected wave present (all is transmitted); Exam 2010 Q2. \square

~~2010 Q2~~

2.4 Waves due to spherical sources (acoustics)

Here some fundamental examples of sound waves generated by "spherical sources" are considered

2.4.1 A radially pulsating sphere



Consider a sphere centered at origin O which "pulsates": its radius changes with time

$$a = a(t)$$

We assume $|a(t) - a_0| \ll a_0$, a_0 "original"

radius i.e. the perturbations in the acoustic medium surrounding the sphere are small

Let the medium have wave speed c , density ρ_0

We expect an outgoing spherically-symmetric acoustic wave generated as a result outside the sphere, i.e.

$r \geq a_0$ ($r := |\underline{x}|$) r is distance to origin

radially symmetric
w/ "q" wave eqⁿ

$$\underline{u} = \nabla \Phi$$

↑
velocity

$$\Phi = f\left(\frac{r-ct}{r}\right)$$

$$\text{Seek } f(r-ct) = -ca_0 \psi\left(t - \frac{r}{c} + \frac{a_0}{c}\right)$$

looking ahead,
when $r=a_0$ last 2
terms cancel

(ψ is still an arbitrary function to be found)

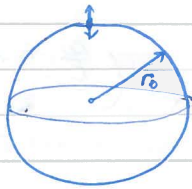
Boundary conditions (BCs): $r = a_0 \Rightarrow$

$$\underline{u} = a'(t) \underline{\tilde{r}} =: v_s(t) \underline{\tilde{r}} \quad \text{where } v_s(t) = a'(t) \text{ is}$$

"sphere's radial velocity"

have previously
used this →
notation

$$\underline{\tilde{r}} := \frac{\underline{x}}{|\underline{x}|} = \text{unit radial vector}$$

 $r(t)$

$$\phi = \frac{f(r-ct)}{r} \quad \text{or} \quad -ca_0 \psi\left(t - \frac{r}{c} + \frac{a_0}{c}\right)$$

$$\underline{u} = a^2(t) \underline{\hat{r}} = v_s(t) \underline{\hat{r}}$$

↑
unit radial
direction

 $r = a_0$

$$\phi = -\frac{ca_0}{r} \psi\left(t - \frac{r}{c} + \frac{a_0}{c}\right) \Rightarrow$$

$$\underline{u} = \nabla \phi = \frac{\partial \phi}{\partial r} \underline{\hat{r}} = u_r \underline{\hat{r}}$$

$$u_r = \frac{\partial \phi}{\partial r} = \frac{ca_0}{r^2} \psi\left(t - \frac{r}{c} + \frac{a_0}{c}\right) + \frac{a_0}{r} \psi'\left(t - \frac{r}{c} + \frac{a_0}{c}\right)$$

$$\text{BCs: at } r = a_0, \quad u_r = v_s(t)$$

$$\Rightarrow \frac{c}{a_0} \psi(t) + \psi'(t) = v_s(t) \quad (19)$$

which is an ODE for $\psi(t)$

I.F: $e^{\frac{c}{a_0}t}$

Solve (19) using integrating factor $e^{\frac{c}{a_0}t}$

$$\frac{d}{dt} \left(e^{\frac{c}{a_0}t} \psi(t) \right) = e^{\frac{c}{a_0}t} v_s(t)$$

For causal solutions, $v_s(t) \equiv \psi(t) \equiv 0$ for $t \leq t_0$

$$\Rightarrow e^{\frac{c}{a_0}t} \psi(t) = \int_{t_0}^t e^{\frac{c}{a_0}\tau} v_s(\tau) d\tau, \quad t \geq t_0$$

$$\Leftrightarrow \psi(t) = \int_{t_0}^t e^{-\frac{c}{a_0}(t-\tau)} v_s(\tau) d\tau$$

So ψ is found $\rightarrow \phi \rightarrow \underline{u} = \nabla \phi$ (velocity found);

$$(1.11) \leftrightarrow \frac{\partial \phi}{\partial t} = -\frac{1}{\rho_0} p \Rightarrow p = -\rho_0 \frac{\partial \phi}{\partial t} \quad \square$$

For T-H case, $v_s(t) = \text{Re} \{ \hat{v}_s e^{-i\omega t} \}$, $\hat{v}_s \in \mathbb{C}$

Seek $\psi(t) = \text{Re} \{ B e^{-i\omega t} \}$, $B \in \mathbb{C}$
to be found

Plugging in to (19):

$$\text{Re} \left\{ \frac{c}{a_0} B e^{-i\omega t} - i\omega B e^{-i\omega t} \right\} = \text{Re} \left\{ \hat{v}_s e^{-i\omega t} \right\} \quad \forall t$$

Holds for ψ so erase Re parts

$$\Rightarrow B = \frac{d_0 \hat{V}_s}{c - i\omega d_0} = \frac{d_0 \hat{V}_s}{c(1 - ikd_0)} \quad \left(\frac{\omega}{c} = k\right)$$

$$\begin{aligned} \Rightarrow \phi &= -\frac{cd_0}{r} \psi\left(t - \frac{r}{c} + \frac{d_0}{c}\right) \\ &= -\frac{cd_0}{r} \operatorname{Re} \left\{ B e^{-i\omega\left(t - \frac{r}{c} + \frac{d_0}{c}\right)} \right\} \\ &= \operatorname{Re} \left\{ \frac{-cd_0}{r} \frac{d_0 \hat{V}_s}{c(1 - ikd_0)} e^{-i\omega t + ikr - ikd_0} \right\} \\ &= \operatorname{Re} \left\{ \frac{d_0^2 \hat{V}_s e^{-ikd_0}}{1 - ikd_0} \cdot \frac{e^{ikr - i\omega t}}{r} \right\} \\ &= \hat{\phi} \end{aligned}$$

From (1.42) $\longleftrightarrow \hat{p} = i\omega\mu_0 \hat{\phi} \implies$

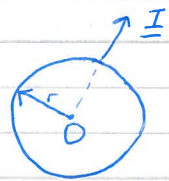
$$p = \operatorname{Re} \left\{ \hat{p} e^{-i\omega t} \right\} = \operatorname{Re} \left\{ \underbrace{\frac{-i\omega\mu_0 d_0^2 \hat{V}_s e^{-ikd_0}}{1 - ikd_0}}_{=A} \cdot \frac{e^{ikr - i\omega t}}{r} \right\} \quad (20)$$

$$= \operatorname{Re} \left\{ \frac{A e^{ikr - i\omega t}}{r} \right\} \quad A := \frac{-i\omega\mu_0 d_0^2 \hat{V}_s}{1 - ikd_0}$$

So for time-averaged intensity $\langle \underline{I} \rangle$,

$$\begin{aligned} \langle \underline{I} \rangle &= \langle I_r \rangle \underline{\hat{r}}; \quad (1.23) \implies \langle I_r \rangle = \frac{|A|^2}{2\mu_0 c r^2} \\ &= \frac{\omega^2 \mu_0^2 d_0^4 |\hat{V}_s|^2 c}{2(1 + (kd_0)^2) \mu_0 c^2 r^2} \quad \leftarrow \text{multiplied by } \frac{c}{c} \end{aligned}$$

$$= \frac{(kd_0)^2 d_0^2 |\hat{V}_s|^2 c \mu_0}{2(1 + k^2 d_0^2) r^2}$$



\implies The average radiated power:

$$\begin{aligned} P_{\text{av}} &= \langle I_r \rangle \times \underbrace{4\pi r^2}_{\text{Area of } S_r = \text{sphere of radius } r} \quad \leftarrow \text{since } \langle I_r \rangle \text{ constant} \\ &= \frac{2\pi (kd_0)^2 d_0^2 |\hat{V}_s|^2 c \mu_0}{1 + (kd_0)^2} \end{aligned}$$

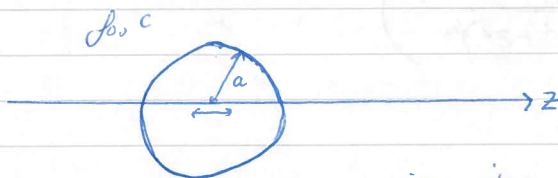
In 'low-frequency' regime $\longleftrightarrow kd_0 \ll 1 \quad (\iff \frac{\omega}{c} d_0 \ll 1 \iff \omega \ll \frac{c}{d_0}, \text{ i.e. } \omega \text{ small})$

$$P_{\text{av}} \approx 2\pi (ka_0)^2 a_0^2 |\hat{v}_s|^2 c f_0 \quad (21)$$

i.e. $P_{\text{av}} \sim \omega^2$, ~~where~~ $k = \frac{\omega}{c}$ i.e. of "order" ω^2 in frequency

□

2.4.2 Transversely oscillating rigid sphere



Let sphere of radius a be rigid (i.e. $a = \text{constant}$) and oscillate along z -axis i.e. its centre (and hence any other

point) has velocity $\underline{v}_c(t) = v_c(t) \hat{\underline{z}}$ ($\hat{\underline{z}} = \underline{e}_z = \text{unit vector in } z \text{ direction}$)

Assume the oscillations small.

In the surrounding acoustic medium ($r \gg a$), seek acoustic velocity $\underline{u} = \nabla \phi$, with $\phi = F(r)$ (see later)

$$\phi = \frac{\partial}{\partial z} \left[\frac{1}{r} \psi \left(t - \frac{r}{c} + \frac{a}{c} \right) \right] \quad (22)$$

[Check: $\frac{1}{r} \psi \left(t - \frac{r}{c} + \frac{a}{c} \right)$ solves wave equation (1.12)]

$$\Rightarrow \frac{1}{c^2} \frac{\partial^2 w}{\partial t^2} - \nabla^2 w = 0$$

$$\Rightarrow \frac{\partial}{\partial z} \left(\frac{\partial^2 w}{\partial t^2} - \nabla^2 w \right) = 0$$

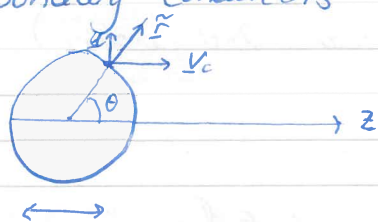
should there be $\frac{1}{c^2}$ there

$$\Leftrightarrow \frac{\partial^2}{\partial t^2} \left(\frac{\partial w}{\partial z} \right) - \nabla^2 \left(\frac{\partial w}{\partial z} \right) = 0$$

$$\Rightarrow \frac{\partial w}{\partial z} \text{ solves (1.12)}$$

So (22) solves (1.12) \forall smooth $\psi(\xi)$

Boundary conditions (BCs):



At $r = a$ (on the sphere's surface) acoustic medium velocity \underline{u} relative to the sphere must be parallel to the surface S

$$\Leftrightarrow (\underline{u} - \underline{v}_c) \cdot \hat{\underline{r}} = 0$$

$$\Leftrightarrow \underline{u} \cdot \hat{\underline{r}} = \underline{v}_c \cdot \hat{\underline{r}}$$

$$\Leftrightarrow \frac{\partial \phi}{\partial r} = \nabla \phi \cdot \underline{r} = \underline{u} \cdot \underline{\hat{r}} = v_c \cdot \underline{\hat{r}} = v_c \cos \theta$$

θ angle between \underline{x} and z -axis \Rightarrow

$$\frac{\partial \phi}{\partial r} = v_c(t) \cos \theta \quad r=a \quad (23)$$

Notice that for 'radially-symmetric' $F = F(r)$

$$\frac{\partial F}{\partial z} = \frac{\partial r}{\partial z} \frac{\partial F}{\partial r}, \quad r = (x^2 + y^2 + z^2)^{1/2}$$

$$\Rightarrow \frac{\partial r}{\partial z} = \frac{z}{r} \left(= \frac{z}{(x^2 + y^2 + z^2)^{1/2}} \right)$$

look at diagr \rightarrow = $\cos \theta$
am to see

this

$$\Rightarrow \frac{\partial F(r)}{\partial z} = \cos \theta \frac{\partial F}{\partial r} \quad \text{So } (22) \Rightarrow$$

$$\Rightarrow \phi = \cos \theta \frac{\partial}{\partial r} \left[\frac{1}{r} \psi \left(t - \frac{r}{c} + \frac{a}{c} \right) \right] \quad (22')$$

\Rightarrow from (23)

$$\cos \theta \frac{\partial^2}{\partial r^2} \left[\frac{1}{r} \psi \left(t - \frac{r}{c} + \frac{a}{c} \right) \right]_{r=a} = v_c(t) \cos \theta \quad (23')$$

$$\text{Now } \frac{\partial}{\partial r} \left[\frac{1}{r} \psi \left(t - \frac{r}{c} + \frac{a}{c} \right) \right] = -\frac{1}{r^2} \psi(\dots) - \frac{1}{cr} \psi' \left(t - \frac{r}{c} + \frac{a}{c} \right)$$

$$\frac{\partial^2}{\partial r^2} \left[\frac{1}{r} \psi(\dots) \right] = \frac{2}{r^3} \psi \left(t - \frac{r}{c} + \frac{a}{c} \right) + \frac{2}{cr^2} \psi'(\dots) + \frac{1}{c^2 r} \psi''(\dots)$$

So (23') \Rightarrow set $r=a$ since this is BC

$$\frac{2}{a^3} \psi(t) + \frac{2}{ca^2} \psi'(t) + \frac{1}{c^2 a} \psi''(t) = v_c(t)$$

$$2c^2 \psi + 2ca \psi' + a^2 \psi'' = c^2 a^3 v_c(t) \quad (24)$$

which is an ODE for ψ (linear 2nd order)

Restricting to T-H case, let

$$v_c(t) = \text{Re} \left\{ \hat{v}_c e^{-i\omega t} \right\} \quad \hat{v}_c \in \mathbb{C}$$

used β previously, changed to A so it cancels later

and seek $\psi(t) = \text{Re} [A e^{-i\omega t + ika}] \quad A \in \mathbb{C}$

$$(24) \Rightarrow (2c^2 A - 2i\omega c a A - \omega^2 a^2 A) e^{-i\omega t + ika} = c^2 a^3 \hat{v}_c e^{-i\omega t}$$

$$A = \frac{c^2 a^3 \hat{v}_c e^{-ika}}{2c^2 - 2i\omega c a - \omega^2 a^2}$$

$$A = \frac{a^3 \hat{v}_c e^{-ika}}{2 - 2ika - k^2 a^2} \quad \left(\frac{\omega}{c} = k \right) \quad (25)$$

$\Rightarrow \psi(t) = \text{Re} [A e^{-i\omega t + ika}] \Rightarrow$ via (22')

$$\Phi = \cos\theta \frac{\partial}{\partial r} \left[\frac{1}{r} \text{Re} \left\{ A e^{-i\omega \left(t - \frac{r}{c} + \frac{\pi}{2} \right) + ika} \right\} \right]$$

↑
replacing t by phase argument

$$= \text{Re} \left\{ \cos\theta \frac{\partial}{\partial r} \left[\frac{A}{r} e^{-i\omega t + ikr - ika + ika} \right] \right\}$$

$$\Phi = \text{Re} \left\{ \underbrace{\cos\theta \frac{\partial}{\partial r} \left(\frac{A}{r} e^{ikr} \right)}_{\hat{\Phi}} e^{-i\omega t} \right\} \quad (26)$$

(1.42)

$$\hat{\rho} = i\omega p_0 \hat{\Phi} = i\omega p_0 \cos\theta \frac{A}{r} e^{ikr} \left(-\frac{1}{r} + ik \right)$$

$$\hat{\rho} = -k\omega p_0 \cos\theta \frac{A}{r} e^{ikr} \left(1 + \frac{i}{kr} \right) \quad (27)$$

In the "far field" ($\Leftrightarrow kr \gg 1 \Leftrightarrow r \gg \frac{1}{k}$) ← 'many wavelengths'

$$\hat{\rho} \approx -k^2 c p_0 \cos\theta \frac{A}{r} e^{ikr} \quad (28)$$

For radial component of velocity u_r

$$u_r = \underline{u} \cdot \underline{\hat{r}} = \frac{\partial \Phi}{\partial r} = \text{Re} \left\{ \cos\theta \frac{\partial^2}{\partial r^2} \left(\frac{A}{r} e^{ikr} \right) e^{-i\omega t} \right\}$$

In the far field

In the far field:

using Leibniz rule for differentiation

$$u_r = \text{Re} \left\{ -k^2 \frac{A}{r} e^{ikr} \left(1 + \frac{2i}{kr} - \frac{2}{(kr)^2} \right) \cos\theta e^{-i\omega t} \right\}$$

$$u_r \approx \frac{-k^2}{r} \cos\theta \text{Re} \left\{ A e^{ikr - i\omega t} \right\}$$

complex phase

$$A = |A| e^{i\psi}$$

For time-averaged radial acoustic intensity

$$\langle I_r \rangle = \langle p u_r \rangle \approx \frac{k^4 c f_0 \cos^2\theta |A|^2}{r^2} \langle \cos^2(\omega t - kr - \psi) \rangle$$

from A

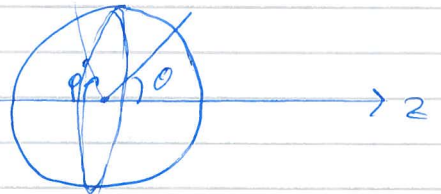
we can see that p will have this form since \hat{p} has e^{ikr}
($kr \gg 1$)

$$\langle I_r \rangle = \frac{k^4 c f_0 |A|^2 \cos^2\theta}{2r^2}$$

not a constant (unlike light)
so need to be careful

The average radiated power: take S_r , sphere of a large radius $r \Rightarrow$

$$P_{\text{av}} = \iint_{S_r} \langle I_r \rangle dS_r$$



$$= \frac{k^4 c f_0 |A|^2}{2r^2} \int_0^{2\pi} \int_0^{\pi} \cos^2\theta \underbrace{r^2 \sin\theta}_{dS} d\varphi d\theta$$

$$= \frac{1}{2} k^4 c f_0 |A|^2 2\pi \int_0^{\pi} \cos^2\theta \sin\theta d\theta$$

$$= -\pi k^4 c f_0 |A|^2 \left[\frac{1}{3} \cos^3\theta \right]_0^{\pi}$$

$$= \frac{2}{3} \pi k^4 c f_0 |A|^2$$

$$(25) \Leftrightarrow A = \frac{a^3 \hat{v}_c e^{-ika}}{2 - 2ika - (ka)^2}$$

since complex we do not know it so take modulus

$$\Rightarrow |A|^2 = \frac{a^6 |\hat{v}_c|^2}{(2 - k^2 a^2)^2 + 4k^2 a^2}$$

oops! wrong again below.

$$= \frac{a^6 |\hat{v}_c|^2}{(2 - k^2 a^2)^2 + 4k^2 a^2}$$

$$= \frac{a^6 |\hat{v}_c|^2}{4 + (ka)^4} \implies$$

$$P_{\text{av}} = \frac{2\pi k^4 c f_0}{3} \frac{a^6 |\hat{v}_c|^2}{4 + (ka)^4}$$

$$P_{\text{av}} = \frac{2\pi}{3} \frac{(ka)^4 c f_0 a^2 |\hat{v}_c|^2}{4 + (ka)^4} \quad (29)$$

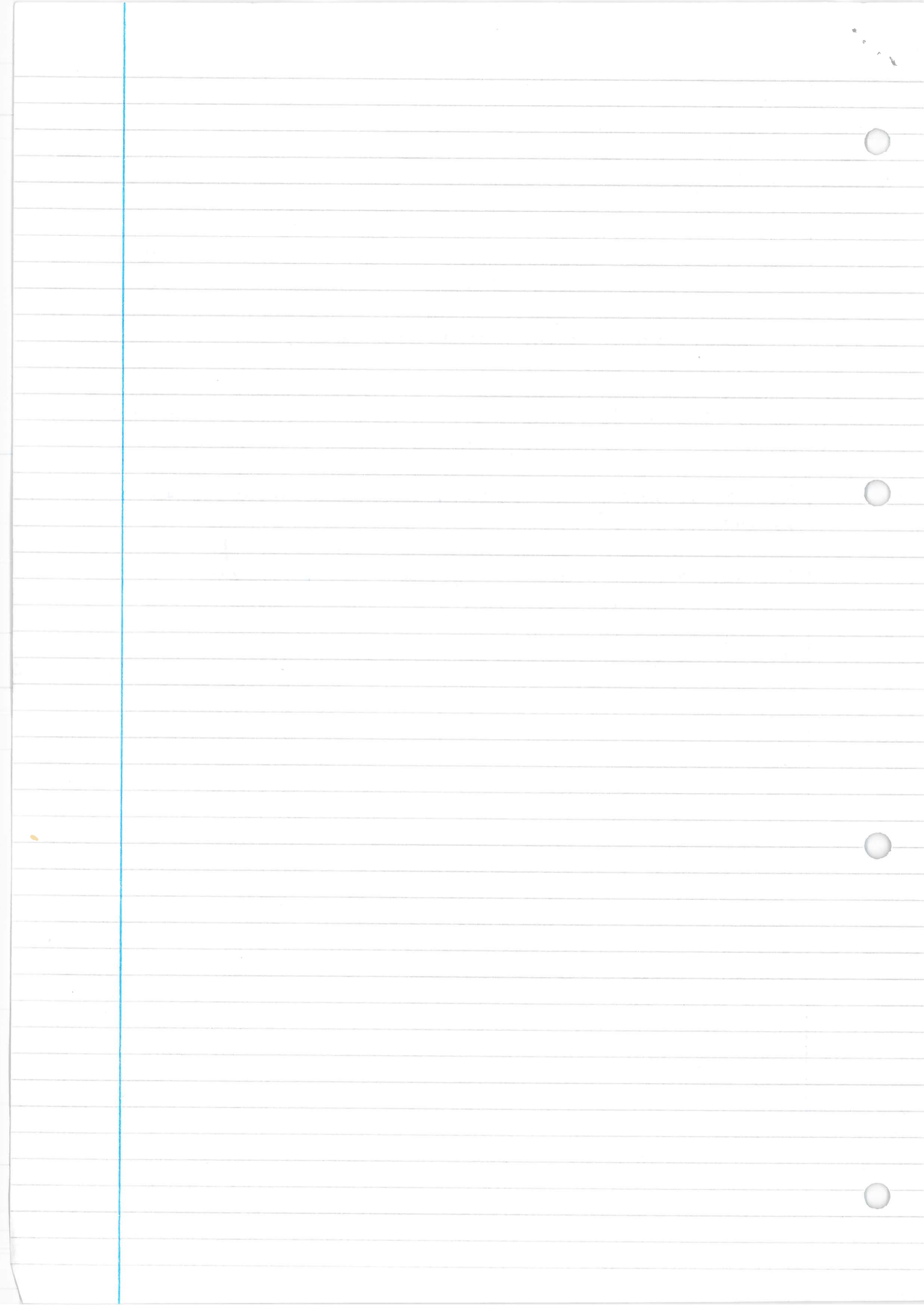
In low frequency regime $\iff ka \ll 1$

$$P_{\text{av}} \approx \frac{\pi}{6} (ka)^4 c f_0 a^2 |\hat{v}_c|^2 \quad (30)$$

Intrinsically transversally osc. sphere produces less energy than pulsating

So the radiated power is "of order 4" with respect to frequency $(ka)^4 = (\omega a/c)^4$

End of chapter 2!



3. Green's functions; Multiple expansions

09/08/15

3.1 Monopole & free-space Green's function



radially symmetric T-H
 Consider spherically pulsating sphere (radius a_0 ; $v_s(t)$)
 $v_s(t) = \text{Re} \{ \hat{v}_s e^{-i\omega t} \}$
 see § 2.4 = 1

velocity
 ↓

$\Rightarrow p = \text{Re} \{ \hat{p} e^{-i\omega t} \}$, where; see (2.20),

defined in whole space w/ singularity at origin \rightarrow

$\hat{p} = \hat{S} \frac{e^{ikr}}{r}$ ← if have outgoing spherically symm. wave this must be the case (1)

$r > a_0$, and $\hat{S} = \frac{-i\omega \rho_0 a_0^2 \hat{v}_s}{1 - ika_0} e^{-ika_0}$

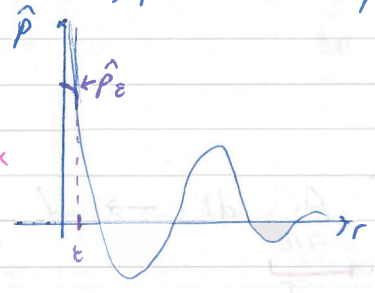
Let $a_0 \rightarrow 0$ and \hat{v}_s increases, so that \hat{S} remains constant
 In the "limit" we have \hat{p} defined by (1) everywhere except $r=0$, with a singularity at $r=0$

Such \hat{p} is called to be due to a "point source" (in reality, a source can be regarded as a point source if:

- (i) $a_0 \ll \lambda = \frac{2\pi}{k}$ (= wavelength) $\Leftrightarrow a_0 t \ll 1$
- (ii) Observation point, $r \gg a_0$

So \hat{p} given by (1) solves Helmholtz equation
 $(\nabla^2 + k^2) \hat{p} = 0$ for $r \neq 0$

should have looked at Re & Im parts separately since \hat{p} is complex valued



For a small $\epsilon > 0$, let $\hat{p}_\epsilon(r) := \hat{p}(r)$ for $r > \epsilon$, but \hat{p}_ϵ is "smoothed out" for $r < \epsilon$

$(\nabla^2 + k^2) \hat{p}_\epsilon = -\hat{S} \Delta_\epsilon(r)$ (2)

in contrast to \hat{p} is H eqⁿ

where RHS $\Delta_\epsilon(r) \neq 0$ only for $r < \epsilon$
 (\Leftrightarrow has 'support' in ball B_ϵ of radius ϵ centred at $r=0$).

Let $\epsilon \rightarrow 0$, fix $R > 0$ and integrate (2) over the ball B_R :

$\int_{B_R} \nabla^2 \hat{p}_\epsilon dV + \int_{B_R} k^2 \hat{p}_\epsilon dV = -\hat{S} \int_{B_R} \Delta_\epsilon(r) dV$ ← delta epsilon

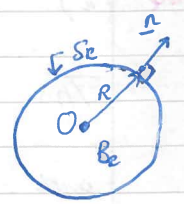
$\Leftrightarrow I_1 + I_2 = I_3$

$I_1 = \int_{B_R} \nabla \cdot (\nabla \hat{p}_\epsilon) dV = \int_{S_R} \nabla \hat{p}_\epsilon \cdot \underline{n} dS$

divergence thm

$= \int_{S_R} \frac{\partial \hat{p}_\epsilon}{\partial n} dS = \int_{S_R} \frac{\partial}{\partial r} \left(\hat{S} \frac{e^{ikr}}{r} \right) dS$

sphere is surface bounding ball



$$= \hat{S} \int_{S_R} e^{ikr} \left(\frac{ik}{r} - \frac{1}{r^2} \right) dS$$

$$= \hat{S} e^{ikR} \left(\frac{ik}{R} - \frac{1}{R^2} \right) \times 4\pi R^2$$

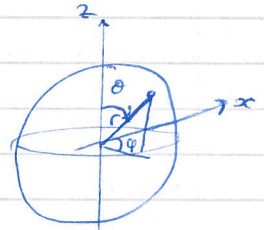
$$= \underline{4\pi \hat{S} e^{ikR} (ikR - 1)}$$

exactly
want to take limit $\epsilon \rightarrow 0$ so can attempt to take it here

$$I_2 = \int_{B_\epsilon} k^2 \hat{p}_\epsilon dV \longrightarrow \int_{B_\epsilon} k^2 \hat{p} dV$$

$$= \hat{S} k^2 \int_0^{2\pi} \int_0^\pi \int_0^R \frac{e^{ikr}}{r} r^2 \sin\theta dr d\theta d\phi$$

dV in sph. coords



$$\int_0^{2\pi} \int_0^\pi \sin\theta d\theta d\phi = 4\pi$$

(r, θ, ϕ) spherical coordinates

$$I_2 = 4\pi k^2 \hat{S} \int_0^R e^{ikr} r dr = -4\pi ik \hat{S} \int_0^R \frac{d}{dr} (e^{ikr}) r dr$$

$$= -4\pi ik \hat{S} e^{ikR} R + 4\pi ik \hat{S} \int_0^R e^{ikr} dr$$

$$= \underline{-4\pi \hat{S} e^{ikR} ikR + 4\pi \hat{S} (e^{ikR} - 1)}$$

since one integral evaluated exactly & other one was limit

$$\Rightarrow I_1 + I_2 \xrightarrow{\epsilon \rightarrow 0} 4\pi \hat{S} e^{ik\epsilon} (ikR - 1 - ikR + 1) - 4\pi \hat{S}$$

$$= -4\pi \hat{S}$$

$$\Rightarrow I_3 \rightarrow -4\pi \hat{S} \text{ as } \epsilon \rightarrow 0$$

$$-\hat{S} \int_{B_\epsilon} \Delta_\epsilon dV \rightarrow -4\pi \hat{S} \iff \int_{B_\epsilon} \underbrace{\frac{\Delta_\epsilon}{4\pi}}_{\delta_\epsilon} dV \rightarrow 1$$

Hence denoting $\delta_\epsilon := \frac{1}{4\pi} \Delta_\epsilon$ we observe:

$$(i) \int_{\mathbb{R}^3} \delta_\epsilon dV \rightarrow 0 \text{ as } \epsilon \rightarrow 0$$

nothing special about radius

R - can integ. over whole space

$$(ii) \delta_\epsilon(\underline{x}) = \delta_\epsilon(|\underline{x}|) = \delta_\epsilon(r) \neq 0 \text{ only within } r < \epsilon$$

This implies $\lim_{\epsilon \rightarrow 0} \delta_\epsilon(\underline{x}) =: \delta(\underline{x})$ called 'Dirac delta-function'

with the "limit" understood appropriately

So, $\delta(\underline{x})$ can be viewed as a function such that

$$(i) \delta(\underline{x}) = \begin{cases} 0 & \underline{x} \neq \underline{0} \\ +\infty & \underline{x} = \underline{0} \end{cases}$$

($\mathbb{R}^3 = \text{whole space}$)

$$ii) \int_{\mathbb{R}^3} \delta(\underline{x}) dV = 1$$

A rigorous definition requires more advanced analysis \longleftrightarrow "distribution theory"

One important property of δ -functions often called "sifting property":
 \forall 'reasonable' (e.g. continuous) functions f

$$\int_{\mathbb{R}^3} \delta(\underline{x} - \underline{x}') f(\underline{x}') dV(\underline{x}') = \int_{\mathbb{R}^3} \delta(\underline{x} - \underline{x}') f(\underline{x}') dV(\underline{x}') \stackrel{\text{integrate w/ } \underline{x}'}{\downarrow} = f(\underline{x}) \quad (3)$$

(Follows from 'definition' of $\delta(\underline{x})$: $\delta(\underline{x} - \underline{x}') \neq 0 \iff \underline{x}' = \underline{x}$
 \rightarrow only $f(\underline{x})$ matters \longleftrightarrow "sifts in")

Hence per \hat{p} given by (i), as $\epsilon \rightarrow 0$,
 $(\nabla^2 + k^2) \hat{p} = -\hat{S} 4\pi \delta(\underline{x})$;

More generally, for a point source at \underline{x}_s ,
 $\hat{p}(\underline{x}) = \hat{S} \frac{e^{ikR}}{R}$, $R := |\underline{x} - \underline{x}_s|$ (1')
 \leftarrow new origin

$$\implies (\nabla^2 + k^2) \hat{p} = -4\pi \hat{S} \delta(\underline{x} - \underline{x}_s) \quad (4)$$

Parameter $\hat{S} \in \mathbb{C}$ referred to as a "monopole amplitude"

Green's functions: The solution of (4) with
RHS = $\delta(\underline{x} - \underline{x}_s)$ i.e. $\hat{S} = -\frac{1}{4\pi}$, is called the
pro-dash instead of S

free-space Green's function $G_f(\underline{x}, \underline{x}')$:
green's free space

$$(\nabla^2 + k^2) G_f(\underline{x}, \underline{x}') = \delta(\underline{x} - \underline{x}') \quad (5)$$

Hence, from (1'), $\hat{S} = -\frac{1}{4\pi}$

$$G_f(\underline{x}, \underline{x}') = -\frac{e^{ik|\underline{x} - \underline{x}'|}}{4\pi|\underline{x} - \underline{x}'|} \quad (6)$$

\leftarrow replacing R by $|\underline{x} - \underline{x}'|$
to be more explicit

(Free-space Green's functions for general, partial differential equations (PDEs) with constant linear coefficients are given also called "fundamental solutions")

3.2 Green's functions

Green's fns play a fundamental role mathematically
Consider a field due to an arbitrary distributed source $f(\underline{x})$:

$$\mathcal{L} \hat{p} := (\nabla^2 + k^2) \hat{p} = f(\underline{x}) \quad (7)$$

where \mathcal{L} is Helmholtz 'differential operator'

$\mathcal{L} : \hat{p} \rightarrow f$; more generally \mathcal{L} could be 'wave eqn, Maxwell's systems related operator etc.

Normally (7) is supplemented by boundary/initial etc. (e.g. Sommerfeld radiation condition) to form a boundary-valued problem (BVP), to determine \hat{p} from f uniquely, symbolically $\hat{p} = \mathcal{L}^{-1} f$

Assume \mathcal{L} invertible. See \mathcal{L}^{-1} as an 'integral' operator

$$\hat{p}(\underline{x}) = (\mathcal{L}^{-1} f)(\underline{x}) = \int_{\Omega} k(\underline{x}, \underline{x}') f(\underline{x}') dV(\underline{x}')$$

Ω is domain of \hat{p}, f ;
 $k(\underline{x}, \underline{x}')$ "kernel" to be found

$$\text{Now, } f(\underline{x}) = (\mathcal{L} \hat{p})(\underline{x}) = \mathcal{L}(\mathcal{L}^{-1} f)(\underline{x})$$

$$= \mathcal{L}_{\underline{x}} \int_{\Omega} k(\underline{x}, \underline{x}') f(\underline{x}') dV(\underline{x}')$$

$$= \int_{\Omega} \mathcal{L}_{\underline{x}} k(\underline{x}, \underline{x}') f(\underline{x}') dV(\underline{x}')$$

we assume we can move $\mathcal{L}_{\underline{x}}$ inside

The Green's function serves the purpose

$$\text{Let } k(\underline{x}, \underline{x}') = G(\underline{x}, \underline{x}') \Rightarrow$$

$$\mathcal{L}_{\underline{x}} k(\underline{x}, \underline{x}') = \mathcal{L} G(\underline{x}, \underline{x}') = \delta(\underline{x} - \underline{x}') \quad \text{by (5)}$$

$$\int_{\Omega} \mathcal{L}_{\underline{x}} k(\underline{x}, \underline{x}') f(\underline{x}') dV(\underline{x}') = \int_{\Omega} \delta(\underline{x} - \underline{x}') f(\underline{x}') dV(\underline{x}') \\ \stackrel{\text{by (3)}}{=} f(\underline{x}) \quad \text{as required}$$

$$(\mathcal{L}^{-1} f)(\underline{x}) = \int_{\Omega} G(\underline{x}, \underline{x}') f(\underline{x}') dV(\underline{x}') \quad (8)$$

For Helmholtz equations on domains with boundaries the Green's function is generally in the form:

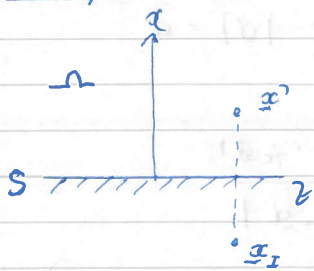
$$G(\underline{x}, \underline{x}') = G_f(\underline{x}, \underline{x}') + G_B(\underline{x}, \underline{x}') \quad \text{with}$$

G_f free-space Green's function (6) and G_B solving:

$$(\nabla^2 + k^2)u_B = 0 \text{ in } \Omega;$$

$u_T + u_B$ satisfy the boundary conditions (BCs)

Example Ω half space - $\{(x, y, z) : x \geq 0\}$



with acoustically soft boundary
 $S = \{x = 0\}$

Let $x' = (x', y', z')$, $x' > 0$, a "source point".
 Consider an "image" $x_I = (-x', y', z') \Rightarrow$
 $x_I \notin \Omega$; and take $u_B(x, x') = -u_T(x, x_I)$

method of images \rightarrow

$$(i) (\nabla^2 + k^2)u_B = -(\nabla^2 + k^2)u_T(x, x_I) \stackrel{(5)}{=} -\delta(x - x_I) = 0$$

in Ω ($\delta(x - x_I) = 0 \forall x \neq x_I \notin \Omega$)

$$(ii) \text{BC: } x \in S \iff x = (0, y, z) \Rightarrow$$

$$u(x, x') = u_T(x, x') - u_T(x, x_I)$$

dist. to all other matters

$$\stackrel{(6)}{\rightarrow} = u_T(|x - x'|) - u_T(|x - x_I|) = 0$$

$$\text{since } |x - x'| = |x - x_I| \iff u(x, x') = 0 \forall x \in S \iff$$

Acoustically soft (Dirichlet) BC □

Neumann rather than Dirichlet

Exercise: For acoustically hard BCs, $\frac{\partial \hat{p}}{\partial n} = n \cdot \nabla \hat{p} = 0$
 on S . Show that

$$u(x, x') = u_T(x, x') + u_T(x, x_I)$$

Exam 2012 Q4(b) □

3.3 Dipoles and QuadrupolesDipolesposition vector
of midpoint

Take two 'monopole' (= point) sources a small distance d apart, with amplitudes \hat{S} and $-\hat{S}$ i.e. oscillating "in anti phase"
 let $\pm \hat{S}$ be at $\underline{x}_s \pm \frac{1}{2} \underline{d}$ where $|\underline{d}| = d$

Resulting acoustic field is:

$$\hat{p}(\underline{x}) \stackrel{(1)}{=} \hat{S} \frac{e^{ik|\underline{x} - \underline{x}_s - \frac{1}{2}\underline{d}|}}{|\underline{x} - \underline{x}_s - \frac{1}{2}\underline{d}|} - \hat{S} \frac{e^{ik|\underline{x} - \underline{x}_s + \frac{1}{2}\underline{d}|}}{|\underline{x} - \underline{x}_s + \frac{1}{2}\underline{d}|}$$

$$\stackrel{(6)}{=} -4\pi \hat{S} G_+(x, x_s + \frac{1}{2}d) + 4\pi \hat{S} G_+ \hat{S} G_+(x, x_s - \frac{1}{2}d) \quad (9)$$

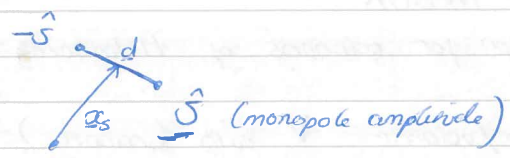
Let $|\underline{d}| = d$ be small $\leftrightarrow kd \ll 1 \Rightarrow$

Using Taylor series,

$$\begin{aligned} \hat{p}(\underline{x}) &= -4\pi \hat{S} G_+(\underline{x}, \underline{x}_s) - 4\pi \hat{S} \nabla_{\underline{s}} G_+(\underline{x}, \underline{x}_s) \cdot \frac{1}{2} \underline{d} \\ &+ 4\pi \hat{S} G_+(\underline{x}, \underline{x}_s) - 4\pi \hat{S} \nabla_{\underline{s}} G_+(\underline{x}, \underline{x}_s) \cdot \frac{1}{2} \underline{d} + O(d^2) \\ &= -4\pi \hat{S} \underline{d} \cdot \nabla_{\underline{s}} G_+(\underline{x}, \underline{x}_s) + O(d^2) \end{aligned} \quad (9)$$

↑
re write \underline{x} here

16/03/15



$$\hat{p}(x) = -4\pi \hat{S} \cdot d \cdot \nabla_s G_f(x, x_s) + O(d^2)$$

To main order, $\hat{p}(x) = -4\pi \hat{D} \cdot \nabla_s G_f(x, x_s)$ (9)

where $\hat{D} := \hat{S} d$ is the dipole amplitude vector

Since $G_f(x, x_s) = G_f(x - x_s)$, see (6)

$$\Rightarrow \nabla_s G_f = -\nabla G \Rightarrow \hat{p}(x) = 4\pi \hat{D} \cdot \nabla G_f(x, x_s) \stackrel{(6)}{=} 4\pi \hat{D} \cdot \left(-\frac{1}{4\pi} \frac{e^{ikR}}{R} \right)$$

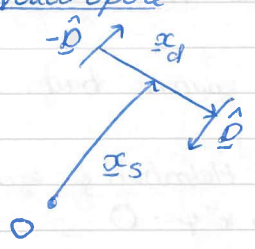
(6) \Rightarrow
 $G_f(x, x') = \frac{-e^{ik|x-x'|}}{4\pi|x-x'|}$

$$\hat{p}(x) = -\hat{D} \cdot \nabla \left(\frac{e^{ikR}}{R} \right) \quad R := |x - x_s| \quad (9')$$

$$\hat{p} = -\frac{d}{dR} \left(\frac{e^{ikR}}{R} \right) \hat{r} \cdot \hat{D} \quad \hat{r} := \frac{x - x_s}{|x - x_s|} \quad (9'')$$

(9)/(9')/(9'') are 'dipole fields'

Quadrupole



Consider now two dipoles a small distance $d = |d|$ apart, with opposite dipole vectors $\pm \hat{D} \Rightarrow$ (for small $d \Leftrightarrow kd \ll 1$)

$$\hat{p}(x) \stackrel{(9)}{=} -4\pi \hat{D} \cdot \nabla_s G_f(x, x_s + \frac{1}{2}d) + 4\pi \hat{D} \cdot \nabla_s G_f(x, x_s - \frac{1}{2}d)$$

Taylor $\approx -4\pi (d \cdot \nabla_s) (\hat{D} \cdot \nabla_s) G_f(x, x_s)$

$$\stackrel{\nabla_s = -\nabla}{=} +4\pi (d \cdot \nabla) (\hat{D} \cdot \nabla) \left(\frac{1}{4\pi} \frac{e^{ikR}}{R} \right) = \sum_{\alpha, \beta=1}^3 Q_{\alpha\beta} \frac{\partial^2}{\partial x_\alpha \partial x_\beta} \left(\frac{e^{ikR}}{R} \right) \quad (10)$$

where $Q_{\alpha\beta} = d_\alpha d_\beta$ $R = |x - x_s|$ $x = (x_1, x_2, x_3)$

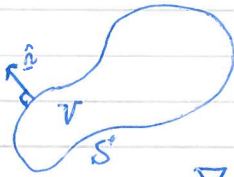
(10) is a 'quadrupole' field \leftrightarrow combination of 2nd derivatives of a 'monopole' field $\frac{e^{ikR}}{R}$

3.4 Kirchhoff - Helmholtz Integral Theorem

This is an important integral representation for solutions of Helmholtz equations
Based on Green's formula / identity:

Let f, g be two functions in domain / volume V with (smooth) boundary S ; Apply divergence thm to

$$\underline{F} = f \nabla g - g \nabla f \quad \Rightarrow$$



$$\int_V \nabla \cdot \underline{F} \, dV = \int_S \underline{F} \cdot \underline{n} \, dS$$

$$\nabla \cdot \underline{F} = \nabla f \cdot \nabla g + f \nabla^2 g - \nabla g \cdot \nabla f - g \nabla^2 f$$

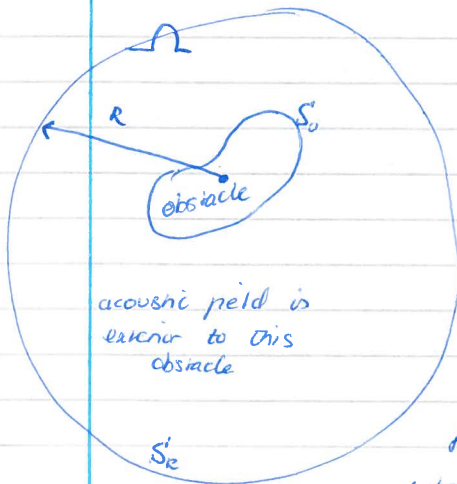
$$\underline{F} \cdot \underline{n} = f \nabla g \cdot \underline{n} - g \nabla f \cdot \underline{n} = f \frac{\partial g}{\partial n} - g \frac{\partial f}{\partial n} \quad \Rightarrow$$

directional derivative

$$\int_V (f \nabla^2 g - g \nabla^2 f) \, dV = \int_S \left(f \frac{\partial g}{\partial n} - g \frac{\partial f}{\partial n} \right) \, dS \quad (11)$$

Important!

(11) is the Green's identity / formula



acoustic field is exterior to this obstacle

Consider a T-H, outgoing, acoustic field outside an obstacle with a boundary S_0 in a domain Ω exterior to

Let V_R be a part of Ω within ball B_R of a large radius R

Let ψ be a solution of Helmholtz equation (1.9) or (1.12) $\nabla^2 \psi + k^2 \psi = 0$

Apply (11) for: $V = V_R$ (part exterior to obstacle)

$$f(\underline{x}) = \psi(\underline{x}), \quad g(\underline{x}) = G(\underline{x}, \underline{x}')$$

where $G(\underline{x}, \underline{x}')$ is the Green's function, $\underline{x}' \in V_R$

$$\Rightarrow S = S_0 \cup S_R \Rightarrow$$

with \underline{x}' in parameter

$$\int_{V_R} \left(\psi(\underline{x}) \nabla^2 G(\underline{x}, \underline{x}') - G(\underline{x}, \underline{x}') \nabla^2 \psi(\underline{x}) \right) \, dV(\underline{x})$$

$$= \int_{S_0 \cup S_R} \left[\psi(\underline{x}) \frac{\partial G(\underline{x}, \underline{x}')}{\partial n(\underline{x})} - G(\underline{x}, \underline{x}') \frac{\partial \psi(\underline{x})}{\partial n} \right] \, dS(\underline{x})$$

$$(5) \Leftrightarrow \nabla^2 a + k^2 a = \delta(\underline{x} - \underline{x}')$$

$$\text{LHS} = \int_{V_R} \psi(\underline{x}) (-k^2 G + \delta(\underline{x} - \underline{x}')) + G(k^2 \psi) \, dV$$

(5), (1.9)

$$= \int_{V_R} \psi(\underline{x}) \delta(\underline{x} - \underline{x}') \, dV(\underline{x}) = \psi(\underline{x}');$$

RHS \Rightarrow : Show that, as $R \rightarrow \infty$, $\int_{S_R} \dots \rightarrow 0$, provided ψ

satisfies also Sommerfeld Radiation Condition (1.20)

$$\Leftrightarrow r \left(\frac{\partial \psi}{\partial r} - ik \psi \right) \rightarrow 0 \quad \text{as } r = |\underline{x}| \rightarrow \infty$$

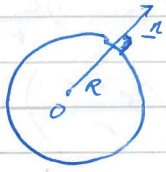
$$\Leftrightarrow \frac{\partial \psi}{\partial r} = ik\psi + o\left(\frac{1}{r}\right) \quad \left(o\left(\frac{1}{r}\right), \text{'o-small'} \Leftrightarrow o\left(\frac{1}{r}\right) \times r \rightarrow 0 \quad r \rightarrow \infty \right)$$

Notice that G , given by (6) also satisfies (1.20):

$$(6) \Leftrightarrow G(x, x') = -\frac{e^{ik|x-x'|}}{4\pi|x-x'|} \underset{|x-x'|=r \rightarrow \infty}{\sim} -\frac{e^{ikr}}{4\pi r} \quad (r \rightarrow \infty)$$

Also $G = O\left(\frac{1}{r}\right)$ as $r \rightarrow \infty$

Also $\psi = O\left(\frac{1}{r}\right)$ as $r \rightarrow \infty \Leftrightarrow (1.20') \Rightarrow$

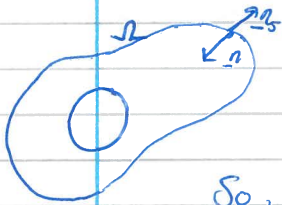


Using also $\frac{\partial}{\partial n} = \frac{\partial}{\partial r} \cdot n = S_n$,

$$\begin{aligned} \int_{S_R} \psi \frac{\partial G}{\partial n} - G \frac{\partial \psi}{\partial n} &= \int_{S_R} \psi \frac{\partial G}{\partial r} - G \frac{\partial \psi}{\partial r} \\ &= \int_{S_R} \psi (ikG + o(r^{-1})) - G (ik\psi + o(r^{-1})) dS \\ &= \int_{S_R} \underbrace{O(r^{-1}) \times O(r^{-1})}_{o(R^{-2})} dS = \int_{S_R} o(R^2) dS = o(R^{-2}) 4\pi R^2 \\ &\rightarrow 0 \text{ as } R \rightarrow \infty \end{aligned}$$

So as $R \rightarrow \infty$, (11) reduces to:

$$\psi(x') = \int_{\partial \Omega} \left[\psi(x) \frac{\partial G(x, x')}{\partial n(x)} - G(x, x') \frac{\partial \psi(x)}{\partial n} \right] dS(x) \quad (12)$$



Notice in (12), \underline{n} is exterior to $\Omega \Leftrightarrow$ interior to obstacle O

So, choosing ψ , a TH pressure \hat{p} , swapping $x \Leftrightarrow x' = x_s$
 $\underline{n} \rightarrow \underline{n}_s = -\underline{n}$, exterior to $O \Rightarrow$

$$\hat{p}(x) = \int_{\partial \Omega} \left[G(x, x_s) \frac{\partial \hat{p}(x_s)}{\partial n_s} - \hat{p}(x_s) \frac{\partial G(x, x_s)}{\partial n_s(x_s)} \right] dS(x_s) \quad (13)$$

(12)/(13) are called Kirchhoff-Helmholtz integral theorem

Notice also that (1.42) $\Leftrightarrow \hat{p} = i\omega p_0 \hat{\phi} \Rightarrow$

$$\frac{\partial \hat{p}}{\partial n_s} = i\omega p_0 \frac{\partial \hat{\phi}}{\partial n_s} = i\omega p_0 \underbrace{\underline{n}_s \cdot \nabla \hat{\phi}}_{\substack{\downarrow \\ \text{normal velocity } \hat{v} \cdot \underline{n}_s}} = i\omega p_0 \hat{v}_n$$

$$G = -\frac{e^{ikR}}{4\pi R}; \quad \frac{\partial G}{\partial n_s}(x, x_s) = \underline{n}_s \cdot \nabla_s G(x, x_s) \\ \stackrel{\underline{n}_s = -\underline{n}}{=} -\underline{n}_s \cdot \nabla G(x, x_s)$$

$$\Rightarrow \hat{p}(x) = -\frac{1}{4\pi} \int_{\partial \Omega} \left[i\omega p_0 \hat{v}_n(x_s) \frac{e^{ikR}}{R} + \hat{p}(x_s) (\underline{n}_s \cdot \nabla) \left(\frac{e^{ikR}}{R} \right) \right] dS(x_s) \quad (13')$$

observation point
far away

Let $k|\alpha_s| \ll 1 \iff$ low frequency; $k|\alpha| \gg k|\alpha_s| \implies$
Taylor series about $\alpha_s = 0$: ($R = |\underline{x} - \alpha_s|$, $r = |\alpha|$)

$$\frac{e^{ikR}}{R} = \frac{e^{ikr}}{r} - (\alpha_s \cdot \nabla) \frac{e^{ikr}}{r} + \frac{1}{2} (\alpha_s \cdot \nabla)^2 \left(\frac{e^{ikr}}{r} \right) + \dots$$

$$\implies \hat{p}(\underline{x}) = \frac{-1}{4\pi} \int_{S_0} [i\omega p_0 \hat{v}_n \left(1 - (\alpha_s \cdot \nabla) + \frac{1}{2} (\alpha_s \cdot \nabla)^2 + \dots \right) \frac{e^{ikr}}{r} + \hat{p}(\alpha_s) (\underline{n}_s \cdot \nabla) \left(1 - (\alpha_s \cdot \nabla) + \frac{1}{2} (\alpha_s \cdot \nabla)^2 + \dots \right) \frac{e^{ikr}}{r}] dS$$

$$\hat{p}(\underline{x}) = \underbrace{\hat{S}}_{\text{monopole term}} \frac{e^{ikr}}{r} - \underbrace{\hat{D} \cdot \nabla \left(\frac{e^{ikr}}{r} \right)}_{\text{dipole term}} + \underbrace{\sum_{\alpha, \beta} Q_{\alpha\beta} \frac{\partial^2}{\partial x_\alpha \partial x_\beta} \left(\frac{e^{ikr}}{r} \right)}_{\text{quadrupole term}} \quad (14)$$

where

monopole term

dipole term

quadrupole term

etc.

$$\hat{S} = \frac{-i\omega p_0}{4\pi} \int_{S_0} \hat{v}_n(\alpha_s) dS(\alpha_s)$$

$$\hat{D} = \frac{-1}{4\pi} \int_{S_0} [i\omega p_0 \hat{v}_n \alpha_s + \hat{p}(\alpha_s) \underline{n}_s] dS(\alpha_s)$$

$$Q_{\alpha\beta} = \dots \text{ (exercise) }$$

(14) is a multipole expansion for $\hat{p}(\underline{x})$

3.5 Scattering example: by air bubbles in water



Let air bubble of initial radius a_0 in water be subjected to an incident acoustic field; a_0 'small':
 $a_0 \ll \lambda = \frac{2\pi}{k}$ = the wavelength

Then for a plane wave incident T-H wave in x -direction
 $p_i = A e^{i(kx - \omega t)} \approx A e^{-i\omega t}$ ($|x| \leq a_0 \Rightarrow kx \ll 1$)

Then (1.16) $\iff \hat{u} = \frac{-i}{\rho_0 \omega} \nabla \hat{p} \Rightarrow u_i$ is small;

For scattered field, seek p_s in a spherically-symmetric form:

$$p_s = B \frac{e^{i(kr - \omega t)}}{r}, \quad B \in \mathbb{C} \quad \text{to be found} \quad (A \in \mathbb{C}, \text{ known}) \quad (15)$$

$$(1.16) \Rightarrow u_s = \frac{1}{i\rho_0 \omega} \frac{\partial p_s}{\partial r} \hat{r} = v_r \hat{r} \quad \text{'radial velocity'}$$

$$\Rightarrow v_r = \frac{B}{i\rho_0 \omega r^2} (i kr - 1) e^{i(kr - \omega t)} \quad (16)$$

Assume the bubble is made of an ideal gas responding 'adiabatically'
 $\iff p V^\gamma = \text{constant}$, $\gamma > 1$ 'adiabatic constant'

Differentiating in t : $\frac{dp}{dt} V^\gamma + \gamma p V^{\gamma-1} \frac{dV}{dt} = 0$

$$\Rightarrow \frac{dp}{dt} = -\frac{\gamma p}{V} \frac{dV}{dt}; \quad \text{Now } p = P_0 + p'$$

$$V = \frac{4\pi}{3} (a_0 + a')^3, \quad p' \ll P_0 \quad a' \ll a_0 \quad (\text{perturbations small})$$

$$\Rightarrow \frac{dp'}{dt} \approx -\frac{\gamma P_0}{\frac{4\pi a_0^3}{3}} \times \frac{4\pi \cdot 3(a_0 + a')^2}{3} \frac{da'}{dt}$$

neglecting perturbation

$$\approx -\frac{3\gamma P_0}{a_0} V_b$$

can neglect to cancel the denominator V_b velocity of bubble

Now p' (the perturbed pressure) assumed T-H

$$\iff p' = \hat{p} e^{-i\omega t} \Rightarrow \frac{dp'}{dt} = -i\omega p'$$

$$\Rightarrow p' = \frac{3\gamma P_0 V_b}{i\omega a_0} \quad (17)$$

$$V_b := \frac{da'}{dt} = \text{radial velocity of bubble}$$

vel. + press. must be same inside + out at surface

On the bubble's surface, must hold:

$$v_b = v_i + v_s \approx v_s = (16); \text{ and}$$

$$p' = \cancel{p_a} p_i + p_s \Big|_{r=a_0}$$

↑ ↑
incident scattered

23/03/15



bubble, radius a_0
acoustic wave incident upon bubble, which acts as a scatterer
 $ka_0 \ll 1$ dimensionless parameter

$$p_i \approx A e^{-i\omega t}$$

$$p_s = \frac{B e^{ikr - i\omega t}}{r} \quad (15)$$

$$v_r = \frac{B}{i\omega r^2} (ikr - 1) e^{ikr - i\omega t} \quad (16)$$

(17)

← the p_s 's called this v_s , I think they are the same

$$P' = \frac{3\gamma P_0}{i\omega a_0} v_b, \quad P' = \underbrace{p_i + p_s}_{r=a_0}, \quad \text{also } v_b \approx v_r \text{ (incident negligible)}$$

$$\downarrow (17), (16)$$

$$\frac{3\gamma P_0}{i\omega a_0} \times \frac{B(ika_0 - 1) e^{ika_0 - i\omega t}}{i\omega a_0^2} = A e^{-i\omega t} + \frac{B e^{ika_0 - i\omega t}}{a_0}$$

replacing v_b with v_r , and $r = a_0$

$$P' \times v_r = p_i + p_s$$

$$B e^{ika_0} \left[-1 + \frac{3\gamma P_0}{\omega^2 \rho a^2} (1 - ika_0) \right] = a_0 A$$

↑
mul. by i

Since $ka_0 \ll 1$, $e^{ika_0} = 1 + ika_0 + O((ka_0)^2)$

$$\Rightarrow B (1 + ika_0 + O((ka_0)^2)) \left[-1 + \frac{\omega_0^2}{\omega^2} (1 - ika_0) \right] = a_0 A$$

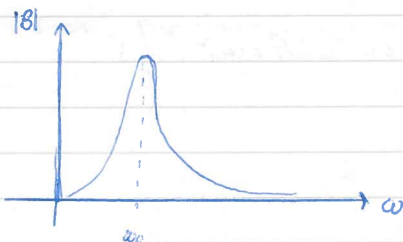
where $\omega_0 := \left(\frac{3\gamma P_0}{\rho a^2} \right)^{1/2}$

$$B \left[-1 - ika_0 + \frac{\omega_0^2}{\omega^2} (1 - ika_0)(1 + ika_0) + O((ka_0)^2) \right] = a_0 A$$

$$B \left[-1 - ika_0 + \frac{\omega_0^2}{\omega^2} + O((ka_0)^2) \right] = a_0 A$$

$$B \approx \frac{a_0 A}{\left(\frac{\omega_0}{\omega} \right)^2 - 1 - ika_0} \quad \Rightarrow \quad |B|^2 = \frac{a_0^2 A^2}{\left[\left(\frac{\omega_0}{\omega} \right)^2 - 1 \right]^2 + (ka_0)^2}$$

from Re \rightarrow $\left[\left(\frac{\omega_0}{\omega} \right)^2 - 1 \right]^2$ \leftarrow from Im $(ka_0)^2$



$|B|$ has a sharp peak at $\omega \approx \omega_0$

Moral: ω_0 is a resonance frequency of the bubble: the scattered wave's amplitude surges for $\omega \approx \omega_0$

Exercise: Exam 2013 Qn 4

4. High frequency waves, WKB method, waveguides

4.1 High frequency (HF) waves, WKB method

The WKB (Wentzel, Kramers, Brillouin) method is a method for constructing highly-oscillatory solutions of differential equations containing a 'large parameter' (e.g. ω , frequency \Rightarrow 'high frequency' regime)

Consider an ODE: $\frac{d^2\psi}{dx^2} + q(x)\psi = 0$, where $|q(x)|$ is 'large';
in particular, if $q(x) > 0$

$$q(x) = k^2(x) \Rightarrow \frac{d^2\psi}{dx^2} + k^2(x)\psi = 0 \quad (1)$$

If $k(x) \equiv k = \text{constant}$, (1) describes plane T-H waves in x -direction (1-D Helmholtz equation) with solution $\psi = Ae^{ikx} + Be^{-ikx}$; $A, B \in \mathbb{C}$
e.g. $\psi_i = \text{Re}(Ae^{ikx - i\omega t}) = |A| \cos(\omega t - kx - \psi)$



when k large ($k \gg 1$)

$$\lambda = 2\pi/k \ll 1 \text{ small}$$

\rightarrow rapid oscillations

$$k = \frac{\omega}{c}, \quad k \gg 1 \iff \omega \gg 1 \iff \text{high frequency}$$

Mathematically, $k(x)$ 'large', can be expressed via $\mu(x)$ ^{parameter}
 $k(x) = \omega\mu(x)$ where $\omega \gg 1$ is a large constant (e.g. $\omega = \text{frequency}$)
and $\mu = O(1)$ is fixed. Then

$$\frac{d^2\psi}{dx^2} + \omega^2\mu^2(x)\psi = 0 \quad (2)$$

$\left[\text{For waves } k(x) = \frac{\omega}{c(x)}, \text{ corresponding to inhomogeneous media with varying wavespeed } c(x) \Rightarrow \mu(x) = c^{-1}(x) \right]$
 $\leftarrow \omega \text{ could not depend on } x, \text{ but } c \text{ could in principle}$

We seek the 'WKB approximation' to a solution of (2) in the form:

$$\psi(x, \omega) = Ae^{i\omega\tau(x, \omega)} \quad (3)$$

with $A \in \mathbb{C}$, $\omega \gg 1$. Then (3) \rightarrow (2):

$$\frac{d\psi}{dx} = A i\omega \frac{d\tau}{dx} e^{i\omega\tau} = i\omega\tau' A e^{i\omega\tau}$$

\uparrow
just derivative since
 ω just parameter

$$\frac{d^2\psi}{dx^2} = [i\omega\tau'' A + i\omega\tau' A i\omega\tau'] e^{i\omega\tau}$$

$$\frac{d^2\psi}{dx^2} = [i\omega\tau'' - \omega^2(\tau')^2] \underbrace{A e^{i\omega\tau}}_{\psi}$$

$$(2) \Rightarrow i\omega\tau'' - \omega^2(\tau')^2 + \omega^2\mu^2(x) = 0 \quad (4)$$

Now seek $\tau(x, \omega)$ in a 'regular perturbation' form wrt small ω^{-1} :

$$\tau(x, \omega) \sim \tau_0(x) + \omega^{-1}\tau_1(x) + \omega^{-2}\tau_2(x) + \dots \quad (5)$$

Plugging (5) in (4), to main order in ω , $O(\omega^2)$:

$$-\omega^2(\tau_0')^2 + \omega^2\mu^2(x) = 0 \implies \tau_0 = \pm \mu(x)$$

$$\implies \tau_0(x) = \pm \int_{x_0}^x \mu(x') dx' + C_{\pm} \quad (6)$$

Next, $O(\omega)$ in (4):

$$i\omega \tau_0'' - \omega^2 2\tau_0' \omega^{-1} \tau_1' = 0$$

$$\implies \tau_1' = \frac{i}{2} \frac{\tau_0''}{\tau_0'} = \frac{i}{2} \frac{d}{dx} \ln|\tau_0'| \implies$$

$$\tau_1(x) = \frac{i}{2} \ln|\tau_0'| + \tilde{C}_{\pm}$$

$$= \frac{i}{2} \ln \mu(x) + \tilde{C}_{\pm}$$

don't have to worry about C_{\pm} in (6) since taking modulus

Hence, from (3) the WKB approximation is

$$\psi(x, \omega) \approx A e^{i\omega(\tau_0(x) + \omega^{-1}\tau_1(x))}$$

$$= A e^{i\omega\tau_0 + i\tau_1} = A \exp \left\{ \pm i\omega \int_{x_0}^x \mu(x') dx' - \frac{1}{2} \ln \mu(x) + C_{\pm} \right\}$$

taking linear combinations of solutions

$$= \tilde{A}_{\pm} \mu^{-1/2}(x) \exp \left\{ \pm i\omega \int_{x_0}^x \mu(x') dx' \right\} \quad \leftarrow C_{\pm} \text{ comes down as constant}$$

or,

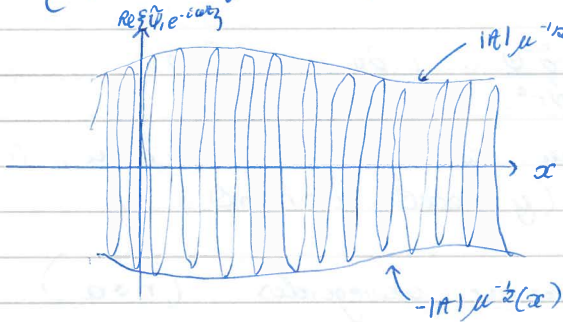
$$\psi(x, \omega) \approx \mu^{-1/2}(x) \left[A \exp \left\{ i\omega \int_{x_0}^x \mu(x') dx' \right\} + B \exp \left\{ -i\omega \int_{x_0}^x \mu(x') dx' \right\} \right] := \tilde{\psi}(x, \omega) \quad (7)$$

where $A, B \in \mathbb{C}$ arbitrary

The WKB approx (7) contains two HF 'waves' travelling to the right & left respectively: $\tilde{\psi} = \tilde{\psi}_1 + \tilde{\psi}_2$

e.g. for $\tilde{\psi}_1$,

$$\text{Re} \left\{ \tilde{\psi}_1 e^{-i\omega t} \right\} = |A| \mu^{-1/2}(x) \cos \left(\omega t - \omega \int_{x_0}^x \mu(x') dx' - \varphi \right)$$



Remark: if we substitute (7) in to (2), then (exercise: check!)

if had exact solution would get 0 on RHS \rightarrow

$$\frac{d^2 \tilde{\psi}}{dx^2} + \omega^2 \mu^2(x) \tilde{\psi}(x) = \left[\frac{3(\mu')^2}{4\mu^2} - \frac{\mu''}{2\mu} \right] \tilde{\psi}$$

This suggests (7) is accurate, provided

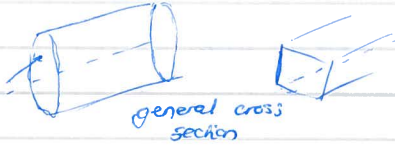
$$\frac{3(\mu')^2}{4\mu^2} - \frac{\mu''}{2\mu} \ll \omega^2 \mu^2 \iff$$

$$\frac{1}{\omega^2} \left[\frac{3(\mu')^2}{4\mu^4} - \frac{\mu''}{2\mu^3} \right] \ll 1 \quad (8)$$

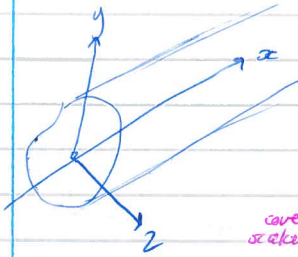
(8) holds provided $\omega \gg 1$ and $\mu(x) \gg \mu_0 > 0$ ($\mu(x)$ not too close to zero) and μ', μ'' 'not too large' ($\mu(x)$ varying slowly enough).

Exercises: j. EM waves in layered media ($\epsilon = \epsilon(x)$ and/or $\mu = \mu(x)$ but not depending on y or z) \Rightarrow MEs reduce to an ODE (2) or alike: Exam 2011 & 2014 Qn 5

4.2 Waveguides



A waveguide is generally a tube/pipe e.g. a cylinder of a rather general cross section, along which waves can propagate (The wave 'bounces' / is reflected from the walls, and so propagates along the axis.)



Choose x along the waveguide's axis; seek for T-H solutions. So, both in acoustics & EM, we seek solutions of Helmholtz equation (1.15) or (1.35)-(1.36),

per $\hat{\psi}$: $(\nabla^2 + k^2)\hat{\psi} = 0$

covering T-H & scalar & vector case

$$\Leftrightarrow \frac{\partial^2 \hat{\psi}}{\partial x^2} + \frac{\partial^2 \hat{\psi}}{\partial y^2} + \frac{\partial^2 \hat{\psi}}{\partial z^2} + k^2 \hat{\psi} = 0$$

$\underbrace{\hspace{10em}}_{=: \nabla_{\perp}^2}$

$k = \frac{\omega}{c}$. In both cases $\hat{\psi}$ is sought in the form

$$\hat{\psi}(x, y, z) = \psi(y, z) e^{i\beta x} \quad (9)$$

where β is unknown 'propagation constant', so (9) looks like a plane wave along x with 'amplitude' ψ depending on y, z

$\Rightarrow \psi$ solves

$$\nabla_{\perp}^2 \psi + (k^2 - \beta^2) \psi = 0 \quad (10)$$

where $\nabla_{\perp}^2 := \frac{\partial^2 \psi}{\partial y^2} + \frac{\partial^2 \psi}{\partial z^2} = \frac{\partial^2 \psi}{\partial r^2} + \frac{1}{r} \frac{\partial \psi}{\partial r} + \frac{1}{r^2} \frac{\partial^2 \psi}{\partial \theta^2}$



is the cross-sectional Laplacian on (y, z) -plane; (r, θ) polar coordinates ($y = r \cos \theta$, $z = r \sin \theta$)

Now specialise to acoustics, and circular waveguides ($r \leq a$)

Seek solns of (10) via separation of variables:

$$\psi(r, \theta) = R(r) \Theta(\theta)$$

$$R'' \Theta + \frac{1}{r} R' \Theta + \frac{1}{r^2} R \Theta'' + (k^2 - \beta^2) R \Theta = 0$$

$\times \frac{r^2}{R \Theta}$

$$\frac{r^2 R''}{R} + \frac{r R'}{R} + (k^2 - \beta^2) r^2 = - \frac{\Theta''}{\Theta} =: m^2$$

(m^2 separation constant) \Rightarrow

$$\Theta'' + m^2 \Theta = 0 \Rightarrow \Theta = e^{\pm i m \theta}$$

So since $\Theta(\theta)$ must be 2π -periodic $\Rightarrow m \in \mathbb{Z}$ (integer); WLOG $m \geq 0$

(waveguide)

$$r^2 R'' + r R' + \left[\underbrace{(k^2 - \beta^2)r^2}_{=: \tilde{r}^2} - m^2 \right] R = 0 \quad (11)$$

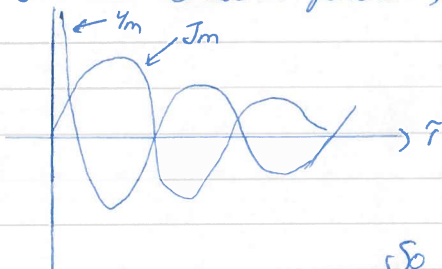
via change of variables, $\tilde{r} = (k^2 - \beta^2)^{1/2} r$

$$\tilde{r}^2 \frac{d^2 R}{d\tilde{r}^2} + \tilde{r} \frac{dR}{d\tilde{r}} + (\tilde{r}^2 - m^2) R = 0 \quad (11')$$

which is Bessel's differential equation of order m , with general solution

$$R = A J_m(\tilde{r}) + B Y_m(\tilde{r}) \quad (12)$$

J_m the Bessel's function, Y_m (so-called) Neumann function



The solution has to be continuous as $r \rightarrow 0$, which J_m is but Y_m is not
 \Rightarrow must take $B=0$

So the solution is of the form

$$\hat{\psi}(x, y, z) = A_m J_m \left(r (k^2 - \beta^2)^{1/2} \right) e^{\pm i m \theta + i \beta z} \quad (13)$$

$m \in \mathbb{Z}, m \geq 0, A_m \in \mathbb{C}$

Let the waveguide's boundary $r=a$ be e.g. acoustically hard i.e.

$$\frac{\partial \hat{\psi}}{\partial n} = \frac{\partial \hat{\psi}}{\partial r} = 0 \quad \text{as } r=a \Rightarrow \text{from (13)}$$



$$J_m' \left(a (k^2 - \beta^2)^{1/2} \right) = 0, \text{ so } \{ \alpha_{mn} \}_{n=1}^{\infty}$$

are roots of $J_m'(\alpha_{mn}) = 0$

$$\Rightarrow a (k^2 - \beta^2)^{1/2} = \alpha_{mn}$$

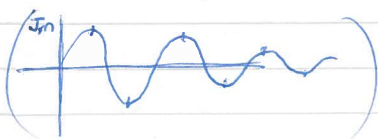
$$\Rightarrow \beta = \pm \left(k^2 - \frac{\alpha_{mn}^2}{a^2} \right)^{1/2} =: \beta_{mn}$$

So \exists infinitely many 'modal solutions'

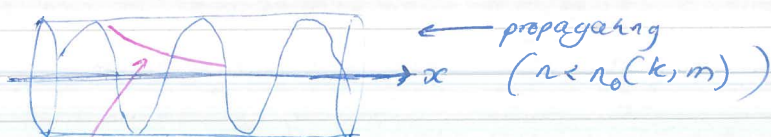
$$\hat{\psi}_{mn} = A_{mn} J_m \left(\frac{r}{a} \alpha_{mn} \right) e^{\pm i m \theta + i \beta_{mn} z}$$

$$m = 0, 1, 2, \dots; n \geq 1$$

Notice that, for fixed k and m , $\alpha_{mn} \rightarrow +\infty$ as $n \rightarrow \infty$



so β_{mn} becomes complex for $n > n_0(k, m)$:
 the corresponding modes are exponentially
 decaying: $e^{i \beta_{mn} z} = e^{-|\beta_{mn}| z}$



exp. decaying ($n > n_0(k, m)$)

