

3402 Waves and Wave Scattering Notes
Based on the spring 2013 lectures by
Prof V Smyshtyayev.

I have made every effort to copy down all the content on the board during lectures. I accept no responsibility what so ever for mistakes on the notes nor changes to the syllabus for the current year. I highly recommend that the reader attends all lectures, making their own notes and to use this document as a reference only.

Chris Owen.

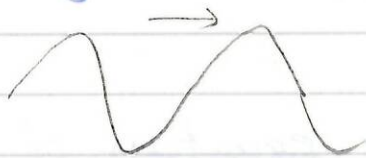
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MATH 3402 Waves and Waves Scattering
Valery S'MYSHLYAEV.

10x3 = 30 classes \rightarrow 100% Exams.

0. Introduction

The course is about maths of waves (sound/acoustic, electromagnetic), and their 'scattering', waves:



- water waves
- sound (acoustics)
- radio/light (electromagnetic)

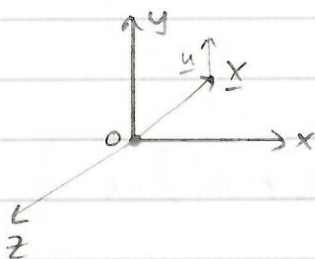
Common waves propagate (hence transfer energy / information; they scatter when hit an "obstacle"; "Disturbances" not "matter" moves!

1. Governing equations

1.1 Acoustic (sound) waves

Acoustic (sound) waves are small amplitude mechanics disturbances propagating in a "fluid", typically a compressible gas (e.g. air) or liquid (e.g. water).

The governing eqn are derived from basic conservation laws for an "inviscid fluid".



$\underline{x} = (x, y, z)$, t time; \mapsto
 $\rho = \rho(\underline{x}, t)$ fluid density
 $\underline{u} = \underline{u}(\underline{x}, t)$ fluid velocity
 $p = p(\underline{x}, t)$ pressure.

Conservation of mass: (cf MATH 2301)

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \underline{u}) = 0 \quad (1)$$

Conservation of momentum (\equiv Newton's 2nd law for continuum media, neglecting gravity, viscosity etc)

$$\frac{\partial \underline{u}}{\partial t} + \underline{u} \cdot \nabla \underline{u} = -\frac{1}{\rho} \nabla p \quad (2)$$

Additionally, physically, an equation of state is held, connecting one unknown to another, e.g. pressure p to density ρ .

Under the "adiabatic" assumption (no heat exchange), let

$$p = P(\rho) \quad (3)$$

We will be considering small disturbances (= perturbations) of a uniform state at rest:

$$p = p_0 + \tilde{p}, \quad \rho = \rho_0 + \tilde{\rho}, \quad \underline{u} = \underline{u}_0 + \tilde{\underline{u}}$$

p_0, ρ_0 are constant; $\tilde{p}, \tilde{\rho}, \tilde{\underline{u}}$ are "small".

Substituting into (1) - (3)

$$(1) \Rightarrow \frac{\partial}{\partial t} (\cancel{\rho_0} + \tilde{\rho}) + \nabla \cdot ((\rho_0 + \tilde{\rho}) \tilde{\underline{u}}) = 0. \quad (1')$$

$$(2) \Rightarrow \frac{\partial \tilde{u}}{\partial t} + \tilde{u} \cdot \nabla \tilde{u} = - \frac{1}{\rho_0 + \tilde{\rho}} \nabla (\rho_0 + p) \quad (2')$$

$$(3) \Rightarrow p_0 + \tilde{p} = P(\rho_0 + \tilde{\rho}) \quad (3')$$

Since $\tilde{p}, \tilde{\rho}, \tilde{u}$ are all small, neglecting in (1') & (2') the "higher order smallness" terms (like $\nabla \cdot (\tilde{\rho} \tilde{u})$) \Rightarrow

$$(1') \Rightarrow \frac{\partial \tilde{p}}{\partial t} + \rho_0 \nabla \cdot \tilde{u} = 0 \quad (4)$$

$$(2') \Rightarrow \frac{\partial \tilde{u}}{\partial t} = - \frac{1}{\rho_0} \nabla \tilde{p} \quad (5)$$

For (3'), using Taylor expansion

$$P(\rho_0 + \tilde{\rho}) \approx P(\rho_0) + \frac{dP}{d\rho}(\rho_0) \tilde{\rho} (+ O(\tilde{\rho}^2))$$

$$\Rightarrow \cancel{p_0} + \tilde{p} \approx \cancel{P(\rho_0)} + \frac{\partial P}{\partial \rho}(\rho_0) \tilde{\rho} \Rightarrow$$

Physically $P' = \frac{\partial P}{\partial \rho}(\rho_0) > 0$, so denote

$$P' = c^2 \quad (\Leftrightarrow c = (P')^{1/2})$$

where the physical constant $c > 0$, the "wave speed" (will later see describes the speed

of waves). $\Rightarrow \tilde{p} = c^2 \tilde{\rho} \Rightarrow$

$$\tilde{\rho} = \frac{1}{c^2} \tilde{p} \quad (6).$$

So, eliminating $\tilde{\rho}$ in (4)-(5) via (6):

$$\frac{1}{c^2} \frac{\partial \tilde{p}}{\partial t} + \rho_0 \nabla \cdot \tilde{u} = 0 \quad (7)$$

$$(5) \Leftrightarrow \frac{\partial \tilde{u}}{\partial t} = -\frac{1}{\rho_0} \nabla \tilde{p} \quad (8)$$

(7)-(8) are main equations of ("linear") acoustics.

To start analysing (7)-(8), differentiate (7) in t and use (8) \Rightarrow

$$\frac{1}{c^2} \frac{\partial^2 \tilde{p}}{\partial t^2} + \rho_0 \nabla \cdot \frac{\partial \tilde{u}}{\partial t} = 0 \Rightarrow$$

$$\frac{1}{c^2} \frac{\partial^2 \tilde{p}}{\partial t^2} + \cancel{\rho_0} \nabla \cdot \left(\cancel{\frac{1}{\rho_0}} \nabla \tilde{p} \right) = 0. \Leftrightarrow$$

$$\frac{1}{c^2} \frac{\partial^2 \tilde{p}}{\partial t^2} - \nabla^2 \tilde{p} = 0 \quad (9)$$

Here $\nabla^2 p := \nabla \cdot (\nabla \tilde{p}) = \frac{\partial^2 \tilde{p}}{\partial x^2} + \frac{\partial^2 \tilde{p}}{\partial y^2} + \frac{\partial^2 \tilde{p}}{\partial z^2}$
 $=: \Delta p$ (=Laplacian of \tilde{p}),

(9) is the wave equation (a partial differential eqn, PDE, for \tilde{p}).

For solutions of (9), $\tilde{p} = \tilde{p}(t, x, y, z)$ seek first \tilde{p} depending on t and x only (not on y & z) \Rightarrow

$$\frac{1}{c^2} \frac{\partial^2 \tilde{p}}{\partial t^2} - \frac{\partial^2 \tilde{p}}{\partial x^2} = 0 \quad (9')$$

$\Rightarrow \tilde{p} = f(x - ct) + g(x + ct)$ solves (9')
 $\forall f, g \in C^2$ (twice differentiable):

$$\frac{1}{c^2} \left((-c)^2 f'' + c^2 g'' \right) - (f'' + g'') = 0 \quad \checkmark$$

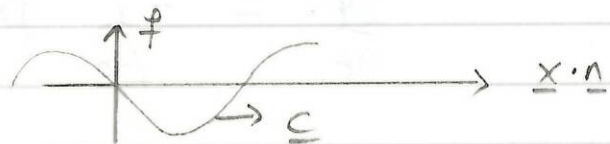
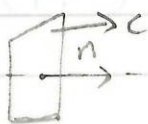
More generally, for any "direction" vector \underline{n} , $|\underline{n}|=1$, a solution of (9) is:

$$\tilde{p} = f(\underline{x} \cdot \underline{n} - ct) + g(\underline{x} \cdot \underline{n} + ct) \quad (10)$$

In (10), $f(\underline{x} \cdot \underline{n} - ct)$ describe a ^{plane} "wave" travelling in positive \underline{n} direction with a "phase speed" c ;

$g(\underline{x} \cdot \underline{n} + ct)$ describes a plane wave travelling in negative \underline{n} direction phase speed c :

$p = f(\underline{x} \cdot \underline{n} - ct)$ is constant \forall plane perpendicular (\perp) to \underline{n} , which plane moves in \underline{n} with speed c :



Returning to (7) & (8):

$$\frac{1}{c^2} \frac{\partial \tilde{p}}{\partial t} + \rho_0 \nabla \cdot \tilde{u} = 0 \quad (7)$$

$$\frac{\partial \tilde{u}}{\partial t} = -\frac{1}{\rho_0} \nabla \tilde{p} \quad (8)$$

Let at time $t=0$ $\nabla \times \underline{u} = 0$ (ie. flow is ^{initially} irrotational)
 \Rightarrow Taking curl of (8) \Rightarrow

$$\frac{\partial}{\partial t} (\nabla \times \tilde{u}) = -\frac{1}{\rho_0} \nabla \times (\nabla \tilde{p}) = 0.$$

$$\Rightarrow \frac{\partial}{\partial t} (\nabla \times \tilde{u}) \equiv 0 \Rightarrow \nabla \times \tilde{u} \equiv \text{const} = 0 \text{ (at } t=0)$$

$\Rightarrow \nabla \times \tilde{u} \equiv 0 \quad \forall t \Rightarrow \exists \phi(x, t)$ a (scalar) velocity potential, such that $\tilde{u} = \nabla \phi$. Then,
(8) $\Leftrightarrow \nabla \left(\frac{\partial \phi}{\partial t} \right) = \nabla \left(-\frac{1}{\rho_0} \tilde{p} \right) \Rightarrow$

$$\nabla \left(\frac{\partial \phi}{\partial t} + \frac{1}{\rho_0} \tilde{p} \right) = 0 \Rightarrow \frac{\partial \phi}{\partial t} + \frac{1}{\rho_0} \tilde{p} = A(t).$$

$A(t)$ constant of integration, must be taken zero for finiteness of perturbation \tilde{p} , $\tilde{u} = \nabla \phi \Rightarrow$

$$\frac{\partial \phi}{\partial t} = -\frac{1}{\rho_0} \tilde{p} \quad (11)$$

Differentiating (11) w.r.t t , via (7):

$$\frac{\partial^2 \phi}{\partial t^2} = -\frac{1}{\rho_0} \left(\frac{\partial \tilde{p}}{\partial t} \right) = +\frac{1}{\rho_0} \left(+c^2 \cancel{\rho_0} \nabla \cdot (\nabla \phi) \right) \Rightarrow$$

$$\frac{1}{c^2} \frac{\partial^2 \phi}{\partial t^2} - \nabla^2 \phi = 0 \quad (12)$$

i.e. ϕ solves the same wave eqn as \tilde{p} cf (9).

Spherically symmetric waves.

Seek for solutions of wave eqn (9)/(12) in spherical coordinates $\underline{x} = (r, \theta, \phi) \Rightarrow$

$$\nabla^2 \tilde{p} = \Delta \tilde{p} = \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial \tilde{p}}{\partial r} \right) +$$

$$+ \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left(\sin \theta \frac{\partial \tilde{p}}{\partial \theta} \right)$$

$$+ \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2 \tilde{p}}{\partial \phi^2}$$

Seek for spherically symmetric solns of (9)
i.e. $\tilde{p} = \tilde{p}(t, r, \cancel{\theta}, \cancel{\phi}) = \tilde{p}(t, r) \Rightarrow$ (9) reads:

$$\frac{1}{c^2} \frac{\partial^2 \tilde{p}}{\partial t^2} - \frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial \tilde{p}}{\partial r} \right) = 0.$$

$$\text{Try } \tilde{p} = \frac{q}{r} \Rightarrow \frac{\partial \tilde{p}}{\partial r} = \frac{1}{r} \frac{\partial q}{\partial r} - \frac{1}{r^2} q \Rightarrow$$

$$r^2 \frac{\partial \tilde{p}}{\partial r} = r \frac{\partial q}{\partial r} - q \Rightarrow$$

$$\frac{\partial}{\partial r} \left(r^2 \frac{\partial \tilde{p}}{\partial r} \right) = r \frac{\partial^2 q}{\partial r^2} + \cancel{\frac{\partial q}{\partial r}} - \cancel{\frac{\partial q}{\partial r}} \Rightarrow$$

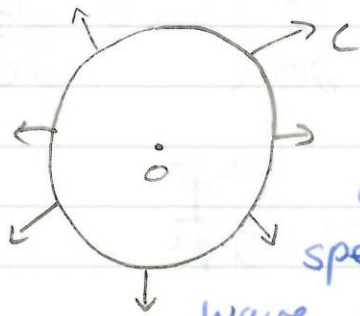
$$\frac{1}{c^2} \cancel{\frac{1}{r}} \frac{\partial^2 q}{\partial r^2} - \cancel{\frac{1}{r}} \frac{\partial^2 q}{\partial r^2} = 0$$

\Rightarrow (1-dim wave eqn for q , as before).

$$q = f(r - ct) + g(r + ct) \Rightarrow$$

$$\tilde{p} = \frac{f(r - ct)}{r} + \frac{g(r + ct)}{r} \quad (13)$$

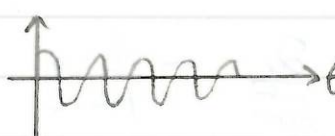
(13) is general form for a spherically symmetric soln of wave equation (9).



f - part of (13) describes a spherical wave travelling away from centre/origin O , with speed c ; g - part describes a wave traveling towards O (f "outgoing", " g " incoming). $\forall f, g \in C^2$

Time-harmonic waves.

Seek for solutions of (9) with following time-dependence: $\tilde{p}(\underline{x}, t) = |\hat{p}(\underline{x})| \cos(\omega t - \psi)$

 i.e. "pure tone" solution with angular frequency ω ,

$\Rightarrow T = \frac{2\pi}{\omega}$ time period; phase ψ , amplitude $|\hat{p}(\underline{x})|$

$$\Leftrightarrow \bar{p}(x, t) = \operatorname{Re}\left\{\hat{p}(x)e^{-i\omega t}\right\}, \quad (14)$$

where $\hat{p}(x) := |\hat{p}(x)|e^{i\psi}$ is called complex amplitude
i.e. incorporating the phase.

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Two hours next Monday.

14/1/13

$u, p, \rho,$

$$\frac{1}{c^2} \frac{\partial^2 \tilde{p}}{\partial t^2} - \nabla^2 \tilde{p} = 0 \quad (9)$$

I-H waves

$$\begin{aligned} \tilde{p}(x, t) &= |\hat{p}| \cos(\omega t - \varphi) \\ &= \operatorname{Re} \left\{ \hat{p}(x) e^{-i\omega t} \right\} \quad (14) \end{aligned}$$

$$\hat{p}(x) = |\hat{p}| e^{i\varphi}$$

Plugging (14) into (9):

$$\operatorname{Re} \left\{ \frac{1}{c^2} (-i\omega)^2 \hat{p} e^{-i\omega t} - (\nabla^2 \hat{p}) e^{-i\omega t} \right\} = 0.$$

$$\Leftrightarrow -\operatorname{Re} \left\{ \underbrace{\left(\nabla^2 \hat{p} + \left(\frac{\omega}{c} \right)^2 \hat{p} \right)}_{S(x)} e^{-i\omega t} \right\} = 0, \forall t, x.$$

$$\Leftrightarrow (\operatorname{Re} S) \cos \omega t + (\operatorname{Im} S) \sin \omega t \equiv 0 \quad \forall t, \forall x$$

$$\Rightarrow \operatorname{Re} S \equiv 0, \operatorname{Im} S \equiv 0 \quad (\text{varying } t \text{ for fixed } x).$$

$$\Leftrightarrow S \equiv 0. \Rightarrow \boxed{\nabla^2 \hat{p} + k^2 \hat{p} = 0}, \quad k := \frac{\omega}{c} \quad (15)$$

(15) is called Helmholtz eqn ("reduced wave eqn" : for T-H waves); $k := \omega/c$ is wave number.

For velocity $\underline{\tilde{u}}(\underline{x}, t)$; (dropping " \sim " hence forth), for T-H case, also seek

$$\underline{u}(\underline{x}, t) = \text{Re}(\underline{\hat{u}}(\underline{x}) e^{-i\omega t})$$

$$\Rightarrow \text{via (8)} \Leftrightarrow \frac{\partial \underline{u}}{\partial t} = -\frac{1}{\rho_0} \nabla p$$

$$\Rightarrow \text{Re}(-i\omega \underline{\hat{u}} e^{-i\omega t}) = -\frac{1}{\rho_0} \text{Re}(\nabla \hat{p} e^{-i\omega t})$$

$$\Leftrightarrow \cancel{\text{Re}} \left\{ (-i\omega \underline{\hat{u}} + \frac{1}{\rho_0} \nabla \hat{p}) \cancel{e^{-i\omega t}} \right\} = 0$$

$$\Leftrightarrow \underline{\hat{u}} = -\frac{i}{\rho_0 \omega} \nabla \hat{p} \quad (16)$$

Plane time-harmonic (T-H) waves:

For a plane wave travelling in positive \underline{n} direction which is also T-H:

$$p(\underline{x}, t) = f(\underline{x} \cdot \underline{n} - ct) = \text{Re} \left\{ \hat{p}(\underline{x}) e^{-i\omega t} \right\}$$

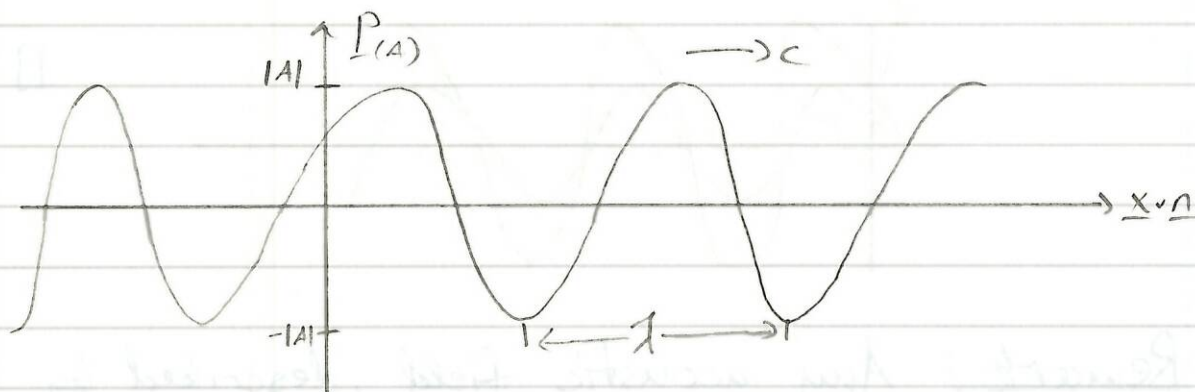
$$= \text{Re} \left(e^{i\frac{\omega}{c}(-ct + \underline{x} \cdot \underline{n})} A \right), \quad A \in \mathbb{C} \text{ (complex constant)}$$

$$p = \operatorname{Re} \left(\underbrace{A e^{i \underline{k} \cdot \underline{x}}}_{\text{"}\hat{p}\text{"}} e^{-i \omega t} \right), \Leftrightarrow \hat{p} A e^{i \underline{k} \cdot \underline{x}} \quad (17)$$

where $\underline{k} := k_{\underline{n}}$ is called wave vector;
 $A = |A| e^{i\varphi}$ complex amplitude

So, (1.7) \Leftrightarrow

$$p(\underline{x}, t) = |A| \cos(\omega t - k_{\underline{n}} \cdot \underline{x} + \varphi)$$



$$\lambda = \frac{2\pi}{k} \text{ is } \underline{\text{wave length}} \text{ (= spatial period)}$$

$$\left(T = \frac{2\pi}{\omega}, \quad f = \frac{1}{T} = \frac{\omega}{2\pi} \right)$$

Also, plugging (17) to (16):

$$\hat{u} = -\frac{c}{\rho_0 \omega} \nabla \hat{p} = -\frac{c}{\rho_0 \omega} \nabla \left(A e^{i \underline{k} \cdot \underline{x}} \right)$$

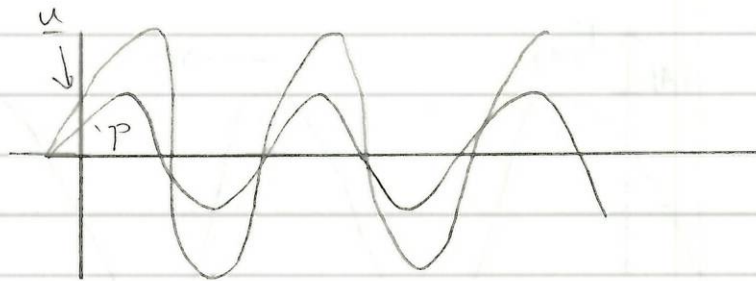
$$= -\frac{c}{\rho_0 \omega} A i \underline{k} e^{i \underline{k} \cdot \underline{x}} = \frac{A}{\rho_0 \omega} \underline{k}_{\underline{n}} e^{i \underline{k} \cdot \underline{x}}$$

$\frac{1}{\rho_0 \omega} = \frac{1}{c}$

$$= \frac{A}{\rho_0 c} \underline{n} e^{i\mathbf{k} \cdot \underline{x}}$$

$$\Rightarrow \underline{u} = \operatorname{Re} \left\{ \frac{A \underline{n}}{\rho_0 c} e^{i\mathbf{k} \cdot \underline{x} - i\omega t} \right\} = \frac{p}{\rho_0 c} \underline{n} \quad (18)$$

i.e. for plane T-H waves, velocity \underline{u} and pressure p are "in phase" and move in the \underline{n} -direction.



Remark: Any acoustic field, described by wave eqn (9) can be expressed as a superposition of T-H waves via Fourier Transform (FT) in time.

Assuming sufficient decay of $p(\underline{x}, t)$ for $t \rightarrow \pm\infty$, seek $p(\underline{x}, t)$ as FT of $p^*(\underline{x}, \omega)$:

$$p(\underline{x}, t) = (2\pi)^{-1/2} \int_{-\infty}^{\infty} p^*(\underline{x}, \omega) e^{-i\omega t} d\omega \quad (*)$$

\Rightarrow the inverse F.T. gives:

$$p^*(\underline{x}, \omega) = (2\pi)^{-1/2} \int_{-\infty}^{\infty} p(\underline{x}, t) e^{i\omega t} dt$$

Then (9) \Rightarrow

$$0 = \frac{1}{c^2} \frac{\partial^2 p}{\partial t^2} - \nabla^2 p = (2\pi)^{1/2} \int_{-\infty}^{\infty} \left[\frac{1}{c^2} (-i\omega)^2 p^* - \nabla^2 p^* \right] e^{-i\omega t} d\omega$$

$$\stackrel{\substack{\uparrow \\ (\frac{\omega}{c} = k)}}{=} (2\pi)^{1/2} \int_{-\infty}^{\infty} (\nabla^2 p^* + k^2 p^*) e^{-i\omega t} d\omega = 0, \quad \forall t$$

\Rightarrow (by FT inversion of zero)

$$\nabla^2 p^* + k^2 p^* = 0.$$

i.e. p^* solves Helmholtz eqn (1.5). So (*) is indeed a superposition of $p^* e^{-i\omega t} \leftrightarrow$ T-H waves.

□

Causality and the Sommerfeld Radiation condition

Causality: the effect/consequence can only follow the cause/source. In the context of acoustic waves, solutions of wave eqn (9)/(12), this means their solutions are required to be zero for $t \leq t_0$ for some time to the "source" switches on.

Consider spherically symmetric waves (13):

$$p(x, t) = \frac{f(r-ct)}{r} + \frac{g(r+ct)}{r}, \quad f, g \in C^2$$

We require: $\exists t_0$ such that, outside a "source region" ($r \leq r_0$), i.e. $\forall r \geq r_0, \forall t < t_0$ holds $p \equiv 0$
 \Rightarrow (Exercise): $g(\vec{r}) = \text{constant}$, so could be "added" to $f \Rightarrow$

$$p(x, t) = \frac{f(r - ct)}{r}$$

is a general "causal" solution of (9), i.e. keeping only outgoing (and removing "incoming") wave, consider now radially-symmetric T-H wave:

$$(13) \Rightarrow p(r, t) = \text{Re} \left\{ A \frac{e^{\frac{i\omega}{c}(-ct+r)}}}{r} + B \frac{e^{-\frac{i\omega}{c}(ct+r)}}}{r} \right\}$$

$$= \text{Re} \left[\frac{A}{r} e^{ikr - i\omega t} + \frac{B}{r} e^{-ikr - i\omega t} \right] \quad A, B \in \mathbb{C}$$

We expect for physical solns not to contain the incoming part, i.e. to make sure $B = 0$.

This is achieved by the Sommerfeld's radiation condition:

$$r \frac{\partial p}{\partial r} - ikp \rightarrow 0, \text{ as } r \rightarrow \infty \quad (19)$$

uniformly as $r \rightarrow \infty$. Check

$$r \frac{\partial p}{\partial r} = \left(\cancel{ikA} e^{ikr} - ikB e^{-ikr} - \frac{Ae^{ikr}}{r} - \frac{Be^{-ikr}}{r} \right) e^{-i\omega t}$$

$$\ominus \quad ikrp = (\cancel{ikAe^{ikr}} + ikBe^{-ikr}) e^{-i\omega t}$$

$$r \frac{\partial p}{\partial r} - ikrp = \left(\underbrace{-2ikBe^{-ikr}}_{\substack{\downarrow \\ 0 \text{ unless } B=0}} - \underbrace{\frac{Ae^{ikr} + Be^{-ikr}}{r}}_{\substack{\downarrow \\ 0, r \rightarrow \infty}} \right) e^{-i\omega t}$$

$\Rightarrow B=0$, as desired. \square

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$$r \frac{\partial p}{\partial r} - ikr p \rightarrow 0, \quad r \rightarrow \infty \quad (19)$$

Hence $p = \frac{A}{r} e^{ikr - i\omega t}$

Remark: Sommerfeld radiation condition (19) is required to hold for any T-H waves, solns of (15). It ensures physically T-H waves are outgoing; mathematically it ensures uniqueness of solutions to boundary-value problems for (15).

One can see that FTs of causal solution to (9) satisfy (19). \square

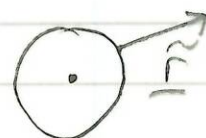
Back to symmetric T-H outgoing waves; under (19)

$$p = \frac{A}{r} e^{ikr - i\omega t} = \hat{p} e^{-i\omega t}, \quad \hat{p} = \frac{A}{r} e^{ikr}$$

$$(16) \equiv \underline{\hat{u}} = \frac{-i}{\rho_0 \omega} \nabla \hat{p} = \frac{-i}{\rho_0 \omega} A \nabla \left(\frac{e^{ikr}}{r} \right)$$

$$= \frac{-iA}{\rho_0 \omega} \left(\frac{ik}{r} - \frac{1}{r^2} \right) e^{ikr} \underline{\hat{r}}, \quad \underline{\hat{r}} := \frac{\underline{x}}{r} = \frac{\underline{x}}{|\underline{x}|}$$

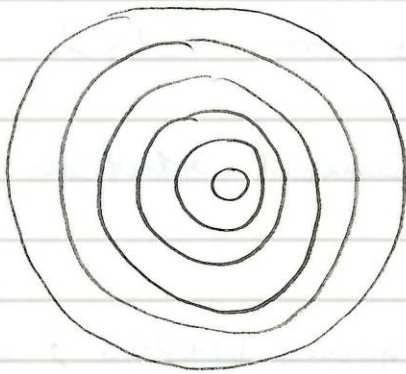
- unit vector in the radial direction.



$$\Rightarrow \underline{\hat{u}} = \hat{u}_r \underline{\hat{r}} \quad , \quad \hat{u}_r = \frac{\hat{p}}{\rho_0 c} \left(1 + \frac{i}{kr} \right), \quad (20)$$

$$\left(\frac{\omega}{k} = c \right)$$

Notice: The 1st term in (20) is same as for plane T-H waves, see (18), and is "in-plane" with pressure p ; 2nd term out of phase with p by $\pi/2$ ($i = e^{i\pi/2}$);



If kr "large", $kr \gg 1$, the 2nd term is negligible $\Leftrightarrow k = 2\pi/\lambda, \Rightarrow r/\lambda \gg 1 \Leftrightarrow r \gg \lambda$
 \Leftrightarrow distances of "many" wavelengths.

Acoustic energy, Intensity (Kirchhoff 1876)
 Going back to main acoustics eqns (7) & (8)

$$(7) \Rightarrow \frac{1}{c^2} \frac{\partial p}{\partial t} + \rho_0 \nabla \cdot \underline{u} = 0.$$

$$(8) \Rightarrow \underline{u} \cdot \rho_0 \frac{\partial \underline{u}}{\partial t} = -\nabla p$$

Take dot product of (8) with $\underline{u} \Rightarrow$

$$\underline{u} \cdot \rho_0 \frac{\partial \underline{u}}{\partial t} = -\underline{u} \cdot \nabla p = -\nabla \cdot (p \underline{u}) + p \nabla \cdot \underline{u}$$

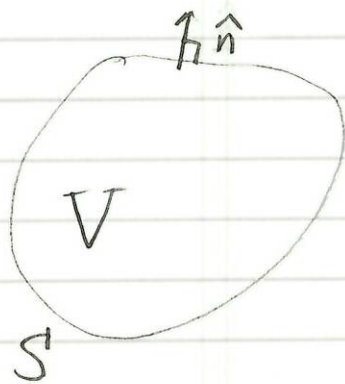
$$= -\nabla \cdot (p \underline{u}) - \frac{p}{\rho_0 c^2} \frac{\partial p}{\partial t}$$

$$\Rightarrow \rho_0 \underline{u} \cdot \frac{\partial \underline{u}}{\partial t} + \frac{1}{\rho_0 c^2} p \frac{\partial p}{\partial t} + \nabla \cdot (p \underline{u}) = 0 \Leftrightarrow$$

$$\frac{\partial}{\partial t} \left(\underbrace{\frac{1}{2} \rho_0 \underline{u} \cdot \underline{u}}_W + \underbrace{\frac{1}{2} \frac{p^2}{\rho_0 c^2}}_I \right) + \nabla \cdot \underbrace{(p \underline{u})}_I = 0 \quad (21)$$

$$\Leftrightarrow \frac{\partial W}{\partial t} + \nabla \cdot \underline{I} = 0.$$

This "conservation law" is interpreted as follows:
Integrate over a volume V with boundary S :



$$0 = \int_V \frac{\partial W}{\partial t} dV + \int_V \nabla \cdot \underline{I} dV$$

$$= \frac{\partial}{\partial t} \int_V W dV + \int_S \underline{I} \cdot \underline{n} dS = 0.$$

(by Gauss / Divergence Theorem)

Interpretation:

$$W = \underbrace{\frac{1}{2} \rho_0 |\underline{u}|^2}_{\text{kinetic energy density}} + \underbrace{\frac{p^2}{2 \rho_0 c^2}}_{\text{acoustic (cf "elastic") potential energy density}}$$

Hence $\underline{I} = p \underline{u}$ = acoustic energy flux or
acoustic intensity (vector):

\underline{I} represent energy transported per unit
area at unit time.

2/1/13

$$\frac{\partial w}{\partial t} + \nabla \cdot \underline{I} = 0, \quad \underline{I} = p \underline{u}.$$

For plane T-H waves:



$$(18) \Rightarrow \underline{u} = \underline{n} \frac{p}{\rho_0 c} \Rightarrow I = p \underline{u} = \frac{p^2}{\rho_0 c} \underline{n}$$

i.e. the energy is transported in the direction of propagation \underline{n} (= phase velocity).

For spherically symmetric T-H waves:

$$p = \text{Re} \left(\frac{A}{r} e^{i(kr - \omega t)} \right) = \frac{|A|}{r} \cos(\omega t - kr - \psi),$$

$$A = |A| e^{i\psi} \text{ (complex amplitude)}$$

$$(20) \Rightarrow \underline{u} = \text{Re}(\hat{\underline{u}} e^{-i\omega t}) = \text{Re}(\hat{u}_r \hat{\Gamma} e^{-i\omega t}) \\ = u_r \hat{\Gamma}$$

$$u_r = \text{Re} \left[\frac{\hat{p}}{\rho_0 c} \left(1 + \frac{i}{kr} \right) e^{-i\omega t} \right]$$

$$= \text{Re} \left[\frac{A}{r \rho_0 c} e^{i kr} \left(1 + \frac{i}{kr} \right) e^{-i\omega t} \right]$$

$$= \frac{|A|}{r \rho_0 c} \cos(\omega t - kr - \psi) + \frac{|A|}{kr \rho_0 c} \sin(\omega t - kr - \psi);$$

$$\Rightarrow \underline{I} = p \underline{u} = p u_r \underline{\tilde{r}} = I_r \underline{\tilde{r}}, \text{ where.}$$

$$I_r = p u_r = \frac{|A|^2 \cos^2(\omega t - kr - \psi)}{\rho_0 c r^2}$$

$$+ \frac{|A|^2}{k r^3 \rho_0 c} \cos(\omega t - kr - \psi) \sin(\omega t - kr - \psi)$$

For time-period average $\langle I_r \rangle := \frac{1}{T} \int_0^T I_r dt$.

($T = 2\pi/\omega$ time period), since:

$$\int_0^T \cos^2(\omega t - kr - \psi) dt = \frac{T}{2} \text{ (exercise)}$$

$$\int_0^T \cos(\omega t - kr - \psi) \sin(\omega t - kr - \psi) dt = 0. \text{ (check)}$$

$$\Rightarrow \langle I_r \rangle = \frac{1}{T} \frac{|A|^2}{\rho_0 c r^2} \frac{T}{2} = \frac{|A|^2}{2 \rho_0 c r^2}$$

$$\text{So } \langle I_r \rangle = \frac{|A|^2}{2 \rho_0 c r^2} = \frac{\langle p^2 \rangle}{\rho_0 c} \quad (22)$$

$$\left(\langle p^2 \rangle = \frac{1}{T} \int_0^T \frac{|A|^2}{r^2} \cos^2(\omega t - kr - \psi) dt = \frac{|A|^2}{2 r^2} \right)$$

So the 2nd term in u_r doesn't contribute.

1.2 Electromagnetic (EM) waves

The EM fields described by the Maxwell's equations (MEs). For electric field $\underline{E}(\underline{x}, t)$, magnetic field $\underline{B}(\underline{x}, t)$, given electric charge density ρ and electric current density \underline{j} in a uniform medium, the MEs are

$$\nabla \cdot \underline{E} = \frac{\rho}{\epsilon} \quad (23), \quad \nabla \cdot \underline{B} = 0 \quad (24)$$

$$\nabla \times \underline{E} = -\frac{\partial \underline{B}}{\partial t} \quad (25), \quad \nabla \times \underline{B} = \mu\epsilon \frac{\partial \underline{E}}{\partial t} + \mu \underline{j} \quad (26)$$

where ϵ, μ are electric permittivity and magnetic permeability of the medium.

[Remark: ME's are generalisations, for continuum EM media, of Coulomb's law = (23), Faraday's law of EM induction = (25), Ampere's law = (26), (24) $\equiv \oint \underline{B}$ froms closed loop.]

In an EM medium without "sources" ($\rho \equiv 0, \underline{j} \equiv 0$), MEs are:

$$\nabla \cdot \underline{E} = 0 \quad (27), \quad \nabla \cdot \underline{B} = 0 \quad (28)$$

$$\nabla \times \underline{E} = -\frac{\partial \underline{B}}{\partial t} \quad (29), \quad \nabla \times \underline{B} = \mu\epsilon \frac{\partial \underline{E}}{\partial t} \quad (30)$$

To eliminate \underline{B} , apply curl to (29), and use (30):

$$\nabla \times \nabla \times \underline{E} = -\frac{\partial}{\partial t} \nabla \times \underline{B} = -\mu \epsilon \frac{\partial^2 \underline{E}}{\partial t^2}$$

Also (from vector calculus)

$$\nabla \times \nabla \times \underline{A} = \nabla(\nabla \cdot \underline{A}) - \nabla \cdot (\nabla \underline{A})$$

$$\nabla \times \nabla \times \underline{E} = \nabla(\nabla \cdot \underline{E}) - \nabla^2 \underline{E}$$

|| by (27)

$$-\nabla^2 \underline{E} = -\mu \epsilon \frac{\partial^2 \underline{E}}{\partial t^2}$$

$$\Leftrightarrow \frac{1}{c^2} \frac{\partial^2 \underline{E}}{\partial t^2} - \nabla^2 \underline{E} = 0 \quad (31)$$

$$c := (\epsilon \mu)^{-1/2} \text{ wave speed.}$$

Similarly (Exercise) $\nabla \times (30) + (29) \Rightarrow$

$$\frac{1}{c^2} \frac{\partial^2 \underline{B}}{\partial t^2} - \nabla^2 \underline{B} = 0 \quad (32)$$

So both \underline{E} and \underline{B} satisfy (vector) wave equations (31), (32), cf (9), (12) in acoustics. For vacuum, $c = c_0 = (\epsilon_0 \mu_0)^{-1/2} \approx 3 \times 10^8 \text{ ms}^{-1}$; if any other media it can only be smaller, physically.

Similarly to acoustics, (31)-(32) admit time-harmonic (TH) solutions:

$$\underline{E}(\underline{x}, t) = \text{Re}(\underline{\hat{E}}(\underline{x}) e^{-i\omega t}) \quad (33)$$

$$\underline{B}(\underline{x}, t) = \text{Re}(\underline{\hat{B}}(\underline{x}) e^{-i\omega t}) \quad (34)$$

where $\underline{\hat{E}}(\underline{x})$, $\underline{\hat{B}}(\underline{x})$ are complex-valued vector fields. As in acoustics (33) $\xrightarrow{\text{sub}} (31)$, (34) $\rightarrow (32)$ yields Helmholtz eqn for $\underline{\hat{E}}$, $\underline{\hat{B}}$, cf (15).

$$\nabla^2 \underline{\hat{E}} + k^2 \underline{\hat{E}} = 0 \quad (35), \quad \nabla^2 \underline{\hat{B}} + k^2 \underline{\hat{B}} = 0 \quad (36)$$

where $k := \omega/c$ the wavenumber.

Plane T-H EM waves.

Similarly to acoustics, the solns of MEs admit plane TH (EM) waves, propagating in any direction \underline{n} . Choosing \underline{n} to coincide with z -direction, they have the form: no generality loss.

$$\underline{E}(\underline{x}, t) = \text{Re}(\underline{E}_0 e^{ikz - i\omega t}), \quad (37)$$

$$\underline{B}(\underline{x}, t) = \text{Re}(\underline{B}_0 e^{ikz - i\omega t}) \quad (38)$$

where \underline{E}_0 , \underline{B}_0 are complex vector constants.

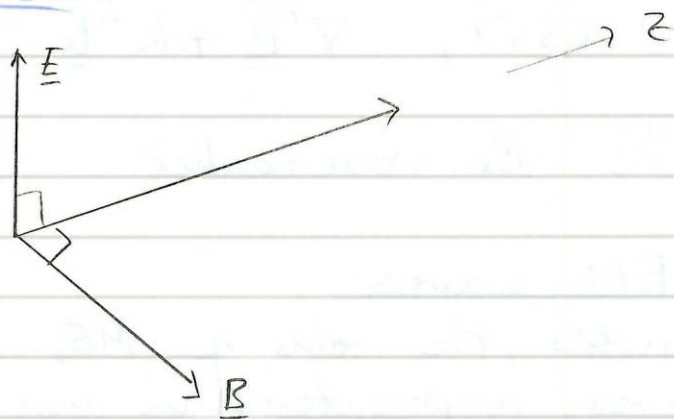
By above construction, $\forall \underline{E}_0, \underline{B}_0$ (31) & (32) are satisfied

$$\text{Now (27)} \equiv \nabla \cdot \underline{E} = 0 = \text{Re} \left((\underline{E})_z i k e^{i k z - i \omega t} \right) = 0.$$

$$\Rightarrow (\underline{E}_0)_z = 0 \iff \underline{E}_0 = (E_x, E_y, E_z), \underline{E}_z = 0$$

$$(28) \equiv \nabla \cdot \underline{B} = 0 \Rightarrow \underline{B}_0 = (B_x, B_y, B_z), \underline{B}_z = 0$$

So both \underline{E}_0 and \underline{B}_0 must be perpendicular (\perp) to propagation direction (z), i.e. EM waves are transverse.



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Recall: $\nabla \cdot \underline{E} = 0$ (27), $\nabla \cdot \underline{B} = 0$ (28)

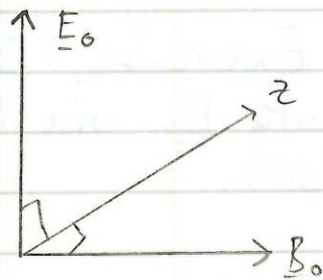
$$\nabla \times \underline{E} = -\frac{\partial \underline{B}}{\partial t} \quad (29), \quad \nabla \times \underline{B} = \mu \epsilon \frac{\partial \underline{E}}{\partial t} \quad (30)$$

$$\underline{E} = \text{Re}(\underline{E}_0 e^{ikz - i\omega t}) \quad (37)$$

$$\underline{E}_0 = (E_x, E_y, 0)$$

$$\underline{B} = \text{Re}(\underline{B}_0 e^{ikz - i\omega t}) \quad (38)$$

$$\underline{B}_0 = (B_x, B_y, 0)$$



Now plug (37) - (38) to (29):

$$\nabla \times \underline{E} = \text{Re} \left[e^{-i\omega t} \begin{vmatrix} \underline{i} & \underline{j} & \underline{k} \\ \frac{\partial}{\partial x} & \frac{\partial}{\partial y} & \frac{\partial}{\partial z} \\ E_x e^{ikz} & E_y e^{ikz} & 0 \end{vmatrix} \right]$$

$$= \text{Re} \left[e^{-i\omega t} (-ikE_y e^{ikz}, ikE_x e^{ikz}, 0) \right]$$

$$= \text{Re} \left[\cancel{ik} e^{ikz - i\omega t} (-E_y, E_x, 0) \right]$$

$$= \text{Re} \left[\cancel{i\omega} e^{ikz - i\omega t} (B_x, B_y, 0) \right]$$

due to (29)

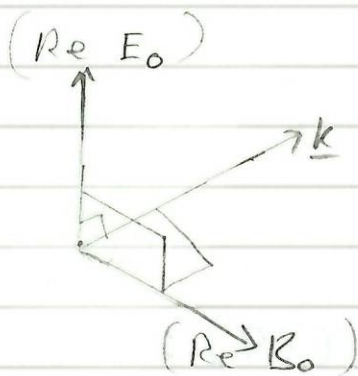
$$\Rightarrow \underline{B}_0 = \frac{k}{\omega} (-E_y, E_x, 0) = \frac{k}{\omega} \begin{vmatrix} \underline{i} & \underline{j} & \underline{k} \\ 0 & 0 & 1 \\ E_x & E_y & E_z \end{vmatrix}$$

$$= \frac{\underline{k}}{\omega} \underline{n} \times \underline{E}_0, \quad \underline{n} = (0, 0, 1) \text{ unit vector in propagation direction, } \Leftrightarrow$$

$$\underline{B}_0 = \frac{1}{\omega} \underline{k} \times \underline{E}_0 \quad (39)$$

where $\underline{k} := k\underline{n}$ the wave vector.

So, via (39), $\underline{E}_0, \underline{B}_0, \underline{k}$ form a right-handed orthogonal triple.



Exercise: Show that $\underline{E}_0 \perp \underline{k}$ with \underline{B}_0 via (39) solve also (30).

So Plane T-H EM waves may have different "polarisation". Polarisation plane is the plane containing \underline{E} and \underline{k} .

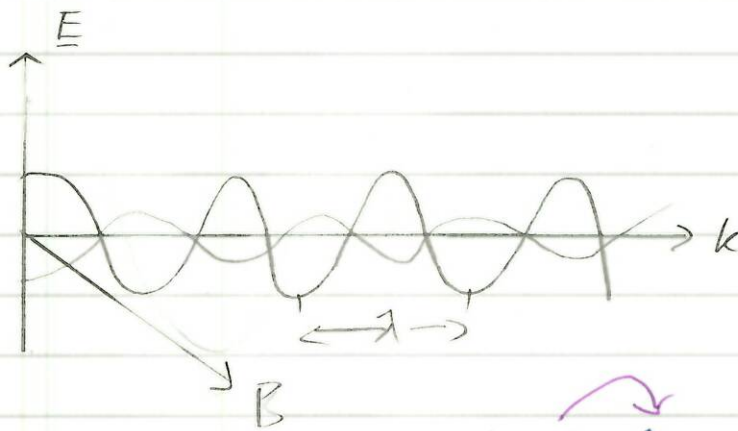
Remember $\underline{E}_0, \underline{B}_0$ are generally complex. If \underline{E}_0 is proportional to a real vector (i.e. $\underline{E}_0 = \underline{\tilde{E}}_0 e^{i\phi}$, $\underline{\tilde{E}}_0$ real component vector) then polarisation plane is unchanged; more generally, polarisation will "rotate", however any plane T-H EM wave can be decomposed in sum of two polarised waves:

$$\underline{E}(x,t) = \text{Re}(\underline{E}_0 e^{i\mathbf{k}\cdot\mathbf{x} - i\omega t})$$

$$= (\text{Re } \underline{E}_0) \cos(\omega t - \mathbf{k}\cdot\mathbf{x})$$

$$+ (\text{Im } \underline{E}_0) \sin(\omega t - \mathbf{k}\cdot\mathbf{x}) =: \underline{E}^{(1)} + \underline{E}^{(2)},$$

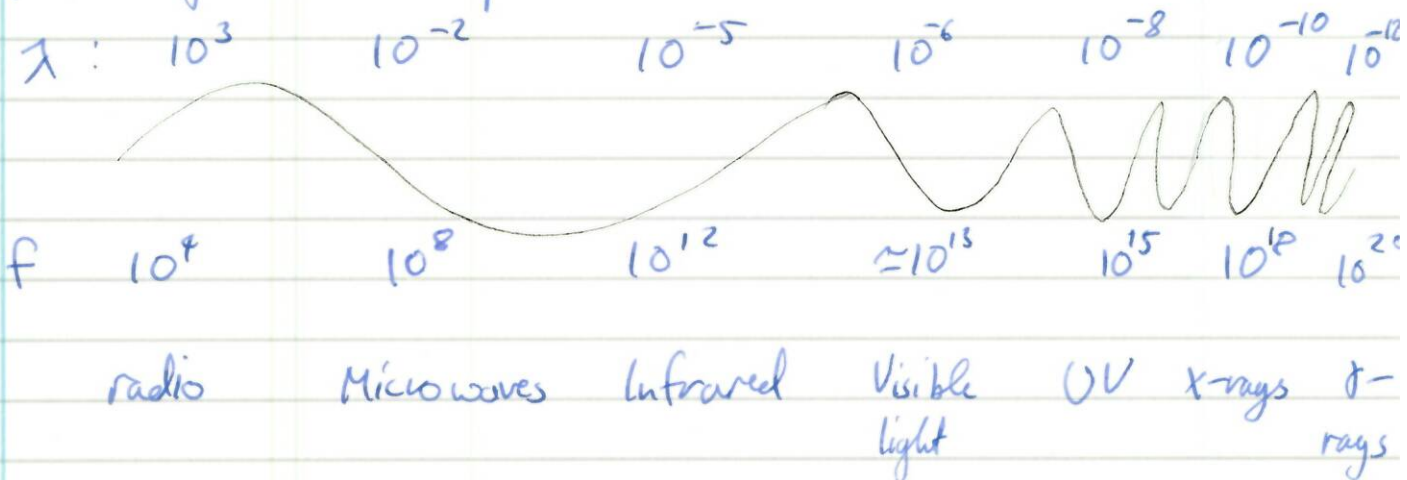
where $\underline{E}^{(1)}$ & $\underline{E}^{(2)}$ are polarised.



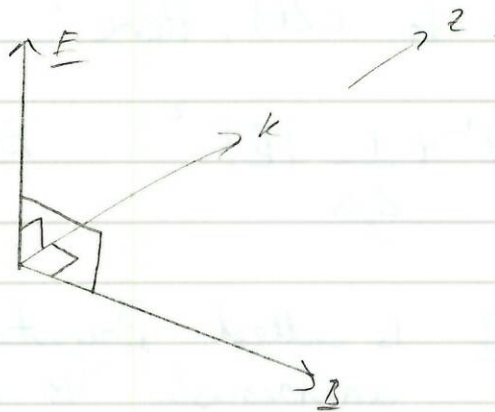
wavelength; $\lambda = \frac{2\pi}{k} = \frac{2\pi c}{\omega} = \frac{c}{f}$

$f = \frac{\omega}{2\pi}$ frequency

Depending on f, λ . EM waves found in different parts of the "spectrum".



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Poynting's vector. Similarly to acoustics of (21), using ME's (29), (30), we can derive a conservation law for EM energy:

$$(29) \Leftrightarrow \underline{B} \cdot \nabla \times \underline{E} = -\frac{\partial \underline{B}}{\partial t}$$

$$(30) \Leftrightarrow \underline{E} \cdot \nabla \times \underline{B} = \mu \epsilon \frac{\partial \underline{E}}{\partial t}$$

Take dot product of (29) with \underline{B} , (30) with \underline{E} and subtract:

$$\underline{B} \cdot (\nabla \times \underline{E}) - \underline{E} \cdot (\nabla \times \underline{B}) = -\underline{B} \cdot \frac{\partial \underline{B}}{\partial t} - \mu \epsilon \underline{E} \cdot \frac{\partial \underline{E}}{\partial t}$$

From vector calculus (check):

$$\nabla \cdot (\underline{A} \times \underline{B}) = (\nabla \times \underline{A}) \cdot \underline{B} - (\nabla \times \underline{B}) \cdot \underline{A}$$

$$\Rightarrow \nabla \cdot (\underline{E} \times \underline{B}) = -\frac{\partial}{\partial t} \left(\frac{1}{2} \underline{B} \cdot \underline{B} + \frac{1}{2} \mu \epsilon \underline{E} \cdot \underline{E} \right)$$

$$\Leftrightarrow \frac{\partial}{\partial t} \left(\frac{1}{2} \epsilon |\underline{E}|^2 + \frac{1}{2\mu} |\underline{B}|^2 \right) + \nabla \cdot \left(\frac{\underline{E} \times \underline{B}}{\mu} \right) = 0 \quad (40)$$

which is similar to (21). Here:

$$w := \frac{1}{2} \epsilon |\underline{E}|^2 + \frac{1}{2\mu} |\underline{B}|^2 \quad \text{EM field's energy density.}$$

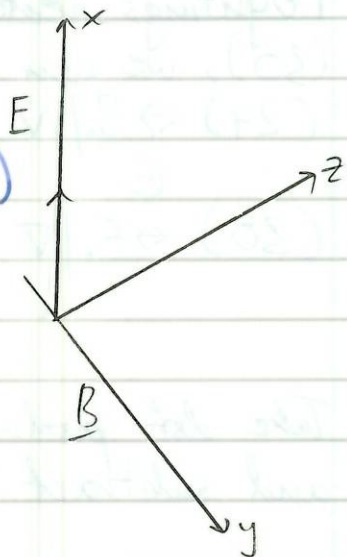
$\underline{S} := \frac{1}{\mu} \underline{E} \times \underline{B}$ is called Poynting's vector which analogously to acoustic intensity \underline{I} describes energy transfer.

For Plane T-H EM waves.

e.g. $\underline{E}(x,t) = (E_x, 0, 0) \cos(kz - \omega t)$

$$\Rightarrow \underline{B}(x,t) = \frac{1}{\omega} (\underline{k} \times \underline{E}_0)$$

$$= \left(0, \frac{E_x}{c}, 0 \right) \cos(\omega t - kz)$$



$$\Rightarrow \underline{S} = \frac{1}{\mu} \underline{E} \times \underline{B} = \frac{1}{\mu c} \begin{vmatrix} \hat{i} & \hat{j} & \hat{k} \\ E_x & 0 & 0 \\ 0 & E_x & 0 \end{vmatrix} \cos^2(\omega t - kz)$$

$$= \frac{|E_x|^2}{\mu c} (0, 0, 1) \cos^2(\omega t - kz)$$

so EM flows in the z -direction \equiv direction of propagation, as expected.

1.3 (Surface) Boundary and Interface conditions

Both in acoustics and EM, boundary/interface conditions are required e.g. at a solid surface or interface between two media (e.g. air and water). Mathematically, a PDEs boundary conditions ^(BC) typically are:

- Dirichlet BC, when the unknown function's value is specified on the boundary S .
- Neumann's BCs, for normal derivative ^{$\frac{\partial p}{\partial n}$} of the function p given on S .
- Mixed / Robin / Impedance BCs, with a combination of p and $\frac{\partial p}{\partial n}$ on S .

13.1 Acoustic BCs

i) An acoustically hard surface doesn't yield to the acoustic wave \leftrightarrow is rigid \leftrightarrow \underline{u} velocity, $\underline{u} \cdot \underline{n} = 0$.
(no motion of S)

(Assuming irrotational motion)
For ϕ velocity potential $\underline{u} = \nabla \phi$
 $\Rightarrow \nabla \phi \cdot \underline{n} = 0 \Leftrightarrow \frac{\partial \phi}{\partial n} = 0 \leftrightarrow$ Neumann's BC for ϕ .

ii) Acoustically soft surface freely yields to the wave (e.g. water surface) \rightarrow Physically on S , $p = 0$, i.e. Dirichlet BC for p .

iii) Impedance BCs are intermediate between (i) and (ii): a surface specific impedance, Z , measures to what extent the surface motion is "impeded" by the resisting pressure.

Physically, Z should depend on frequency, $Z = Z(\omega)$, so for T-H acoustics $p = \text{Re}(\hat{p}e^{-i\omega t})$, $\underline{u} = \text{Re}(\hat{\underline{u}}e^{-i\omega t})$

$$Z := \frac{\hat{p}}{\hat{u}_n}, \quad \hat{u}_n := \hat{\underline{u}} \cdot \underline{n} \quad (\text{normal velocity}).$$

$$\Leftrightarrow \hat{p} - Z\hat{u}_n = 0 \Leftrightarrow \hat{p} - Z \frac{\partial \hat{\phi}}{\partial n} = 0.$$

$$\left(\phi = \text{Re}(\hat{\phi}e^{-i\omega t}) \right)$$

$$\text{From (11)} \Leftrightarrow \frac{\partial \phi}{\partial t} = -\frac{1}{\rho_0} p \stackrel{\text{TH}}{\Leftrightarrow}$$

$$i\omega \hat{\phi} e^{-i\omega t} = -\frac{1}{\rho_0} \hat{p} e^{-i\omega t} \Rightarrow \hat{p} = i\omega \rho_0 \hat{\phi} \quad (41)$$

Hence

$$i\omega \rho_0 \hat{\phi} - Z(\omega) \frac{\partial \hat{\phi}}{\partial n} = 0 \quad (42)$$

which is the impedance / mixed BC for $\hat{\phi}$. (Same (42) $\hat{\phi} \rightarrow \hat{p}$)

$$i\omega \rho_0 \hat{p} - Z(\omega) \frac{\partial \hat{p}}{\partial n} = 0 \quad (42')$$

Notice: if $Z \rightarrow 0 \Rightarrow \hat{p} = \hat{\phi} = 0$ (soft BCs (ii)); $Z \rightarrow \infty (\Leftrightarrow 1/Z \rightarrow 0) \Rightarrow \partial \hat{\phi} / \partial n = 0$ (hard BCs (i)).

Generally, Z is complex; physically $\text{Re } Z \geq 0$, with $\text{Re } Z > 0$ for energy absorbing surfaces, $\text{Re } Z = 0$ for neutral, c.f.)

Exercise - 1 Qn 6(ii) to § 2.1.

1.3.2 EM BCs / Interface condns:

② $\uparrow \underline{n}$ ϵ_2, μ_2 $\frac{\underline{E}^{(2)}}{B^{(2)}}$ At an interface between two EM media
 ① ϵ_1, μ_1 $\frac{\underline{E}^{(1)}}{B^{(1)}}$ (ϵ_1, μ_1 medium 1; ϵ_2, μ_2 medium 2)
 the physics of EM implies the following interface conditions:

(i) Tangential components of \underline{E} are continuous (cts): \Leftrightarrow

$$\underline{E}^{(1)} \times \underline{n} = \underline{E}^{(2)} \times \underline{n}. \quad (43)$$

(ii) Normal components of \underline{B} are cts \Leftrightarrow

$$\underline{B}^{(1)} \cdot \underline{n} = \underline{B}^{(2)} \cdot \underline{n}. \quad (44)$$

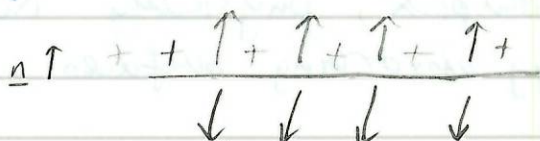
Also, if there exist a surface charge distribution ρ_s and/or surface current distn \underline{j}_s , then

$$(iii) \epsilon_2 \underline{E}^{(2)} \cdot \underline{n} - \epsilon_1 \underline{E}^{(1)} \cdot \underline{n} = \rho_s \quad (45)$$

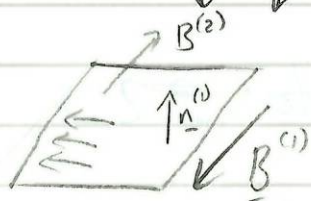
$$(iv) \frac{1}{\mu_1} \underline{B}^{(1)} \times \underline{n} - \frac{1}{\mu_2} \underline{B}^{(2)} \times \underline{n} = \underline{j}_s \quad (46)$$

(In (45), (46) \underline{n} points into medium 2.)

[Check:



$$\rho_s > 0 \Rightarrow \underline{E}^{(2)} \cdot \underline{n} > 0, \\ \underline{E}^{(1)} \cdot \underline{n} < 0$$



+ "right hand rule" from Physics \rightarrow OK]

Perfectly conducting BCs: If Medium 2 is a "perfect conductor" $\Rightarrow \underline{E}^{(2)} = 0 \Rightarrow$
from (43), $\underline{E}^{(1)} \times \underline{n} = 0$ (47)



①

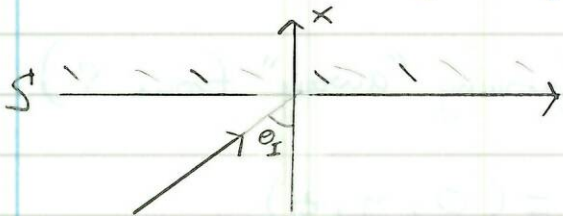
i.e. zero normal component of \underline{E} along a perfectly conducting surface.

(End of Chapter 1)

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2. Canonical Cases:

2.1 Reflection by a plane (acoustics)



Consider a T-H acoustic plane wave incident upon a plane boundary:

$$S = \{(x, y, z) \mid x=0\}$$

$$p_I = A_I e^{-i\omega t - i\mathbf{k}_I \cdot \underline{x}}$$
$$= e^{-i\omega t} \hat{p}_I,$$

$$\hat{p}_I = A_I e^{i\mathbf{k}_I \cdot \underline{x}};$$

assume the wave has "angle of incidence" $\theta_I \Rightarrow$

$$\mathbf{k}_I = k(\cos \theta_I, 0, \sin \theta_I),$$

$$k := |\mathbf{k}_I| = \omega/c$$

$0 \leq \theta_I < \pi/2$ (moving "towards" S).

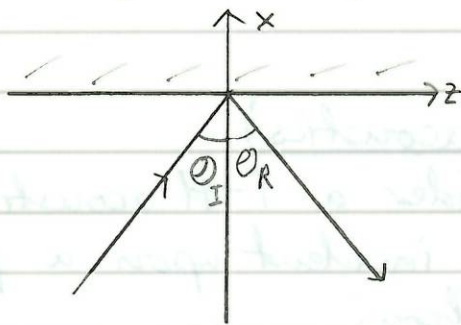
The boundary S will "reflect" the incident wave p_I . What is the reflected wave p_R ? This depends on BC's on S .

Consider acoustically soft BC's: $p=0$ on S (§ 1.31)

$$\Leftrightarrow p = p_I + p_R = 0 \text{ on } S.$$

Seek p_R also a plane T-H wave:

$$p_R = A_R e^{-i\omega_R t + i \underline{k}_R \cdot \underline{x}},$$



$$\underline{k}_R = (k_x, k_y, k_z), \quad |\underline{k}_R| = \omega_R/c = k_R$$

$k_x < 0$ (moving "away" from S')

$$\text{BCs} \Rightarrow \underline{x} \in S' \Leftrightarrow x = 0, \quad \underline{x} = (0, y, z)$$

$$p = p_R + p_I = 0 \Leftrightarrow p_R = -p_I$$

$$\Leftrightarrow A_R e^{-i\omega_R t + i(k_{Ry}y + k_{Rz}z)} = -A_I e^{-i\omega t + i k \sin \theta_I z} \quad \forall t, y, z$$

$$\Rightarrow \omega_R = \omega \text{ (varying } t); \quad k_y = 0 \text{ (varying } y);$$

$$k_z = k \sin \theta_I \text{ (varying } z). \Rightarrow k_R = \omega_R/c = \omega/c = k;$$

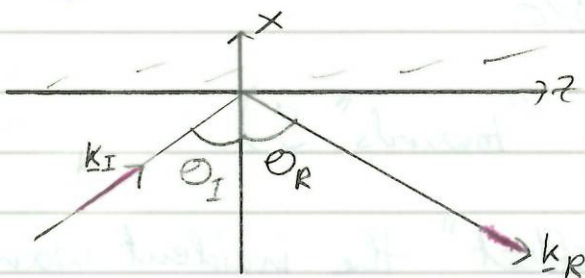
$$A_R = -A_I; \quad k_x^2 + k_z^2 = k^2 \Rightarrow k_x = \pm \sqrt{k^2 - k_z^2}$$

$$\Rightarrow k_x = -\sqrt{k^2 - k^2 \sin^2 \theta_I} \text{ since } k_x < 0.$$

$$= -k \cos \theta_I, \text{ so}$$

$$\underline{k}_R = k(-\cos \theta_I, 0, \sin \theta_I);$$

Interpretation.



$$\underline{\theta}_R = \underline{\theta}_I \text{ (Specular reflection 'law')}$$

$$A_R = -A_I \text{ (Phase "shifts" for } \pi, -1 = e^{-i\pi})$$

Excercise (A) Show that for acoustically hard plane boundaries ($\partial p / \partial n = 0 \Leftrightarrow \partial p / \partial x = 0$)

$$p_R = A_I e^{-i\omega t + i k(-x \cos \theta_I + z \sin \theta_I)}$$

ii) Find the reflected wave for a general impedance BC (1.41)
(For same geometry and incident wave).

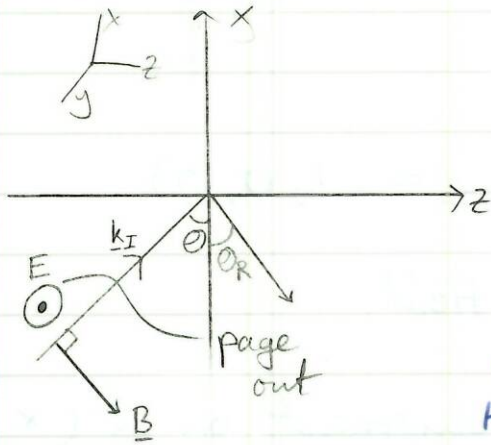
$$i\omega\rho \cdot \hat{\phi} - Z \frac{\partial \hat{\phi}}{\partial n}$$

Conclude that $|A_R| \leq |A_I| \Leftrightarrow \text{Re} Z \geq 0$ (absorbing / neutral boundary).

□

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2.2 Reflection of plane T-H EM waves by perfectly conducting planes:



Consider a plane T-H EM wave propagating within "insulator" medium. (no currents j) medium occupying half-space $x < 0$; with incidence angle θ , unknown angle of reflection θ_R , ϵ, μ given $\rightarrow c = (\epsilon\mu)^{-1/2}$ wave speed.

Let the incident wave be polarised so that \underline{E}_I is parallel to y -axis, pointing out of the plane, i.e.

$$\underline{E}_I = (0, E_I, 0) e^{-i\omega t + ik(x\cos\theta + z\sin\theta)} \quad (1)$$

$$\underline{B}_I = \frac{\underline{E}_I}{c} (-\sin\theta, 0, \cos\theta) e^{-i\omega t + ik(x\cos\theta + z\sin\theta)} \quad (2)$$

$$\underline{k}_I = k(\cos\theta, 0, \sin\theta)$$

The reflected wave is also sought as plane wave:

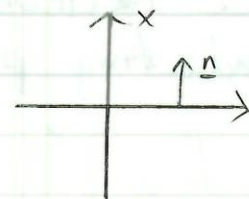
$$\underline{E}_R = (E_x, E_y, E_z) e^{-i\omega_R t + ik_R(-x\cos\theta_R + z\sin\theta_R)} \quad (3)$$

$$\underline{B}_R = (B_x, B_y, B_z) e^{-i\omega_R t + ik_R(-x\cos\theta_R + z\sin\theta_R)} \quad (4)$$

$k_R = \omega_R/c$, $E_x, E_y, E_z, B_x, B_y, B_z$ unknown coeff.

From perfectly conducting BCs (1.47) on

$$S = \{x=0\}, \quad \underline{E} \times \underline{n} = 0.$$



\underline{n} normal vector to $S' \Leftrightarrow \underline{n} = (1, 0, 0)$

$\underline{E} = \underline{E}_I + \underline{E}_R =$ total electric field.

$\Leftrightarrow \underline{E}$ has zero y & z -component on S ($x=0$)

y -component:

$$E_y e^{-i\omega_R t + i k_R z \sin \theta_R} = -E_I e^{-i\omega t + i k z \sin \theta} \quad \forall t, z$$

$$\Rightarrow \omega_R = \omega \Rightarrow k_R = k (= \omega/c) \Rightarrow \theta_R = \theta$$

(i.e. specular reflection law, same as in acoustics)

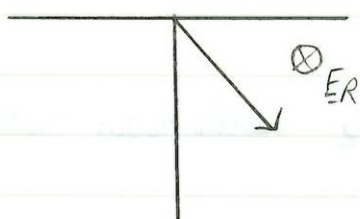
$$E_y = -E_I.$$

z -component: $\Rightarrow E_z = 0$.

Also, since $(\underline{E}_R, \underline{B}_R)$ is a plane wave \Rightarrow

$$\underline{E}_R \perp \underline{k}_R \Leftrightarrow \underline{E}_R \cdot \underline{k}_R = 0 \Leftrightarrow -E_x \cos \theta_R + 0 + 0 = 0$$

$$\Rightarrow E_x = 0.$$



So $\underline{E}_R = (0, -E_I, 0) e^{-i\omega t + i k_R \cdot \underline{x}}$
and points in the plane.

Finally, from (1.39)

$$\underline{B} = \frac{1}{\omega} \underline{k}_R \times \underline{E}_R = \frac{k}{\omega} \begin{vmatrix} \hat{c} & \hat{j} & \hat{k} \\ -\cos\theta & 0 & \sin\theta \\ 0 & -E_I & 0 \end{vmatrix} e^{-i\omega t + i\hat{k}_R \cdot \underline{x}}$$

$$= \frac{E_I}{c} (\sin\theta, 0, \cos\theta) e^{-i\omega t + i\hat{k}_R \cdot \underline{x}}$$

Finally, let's us find "induced electric current" on S:

Notice $\underline{E} \equiv 0$ in $x > 0 \Rightarrow \underline{B} \equiv 0$ in $x > 0$ from ME. (1.25).

$$\frac{\partial \underline{B}}{\partial t} = 0 \Rightarrow \underline{B} = 0 \text{ (since } \underline{B} \text{ T-H)}$$

Now from interface condition (1.46):

$$\frac{1}{\mu} \underline{B}^{(1)} \times \underline{n} - \frac{1}{\mu} \underline{B}^{(2)} \times \underline{n} = \underline{j}_s$$

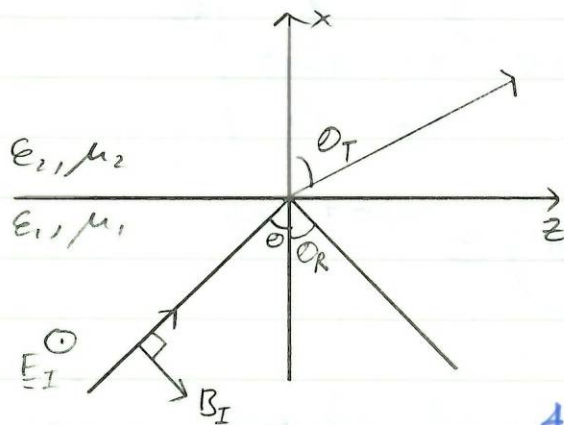
$$\Rightarrow \underline{j}_s = \frac{1}{\mu} (\underline{B}_I + \underline{B}_R) \times (1, 0, 0)$$

$$= \frac{E_I}{c\mu} \begin{vmatrix} \hat{c} & \hat{j} & \hat{k} \\ 0 & 0 & 2\cos\theta \\ 1 & 0 & 0 \end{vmatrix} e^{-i\omega t + i\hat{k}_R \cdot \underline{x}}$$

$$= \frac{2\cos\theta E_I}{c\mu} e^{-i\omega t - i\hat{k}_R \cdot \underline{x}}$$

the induced current \square

2.3 Reflection and refraction by a plane interface between two insulator media.



Consider two insulator EM media with interface $S = \{x=0\}$, ϵ_1, μ_1 medium 1 ($x < 0$), ϵ_2, μ_2 medium 2 ($x > 0$).

A plane wave incident from medium 1 same as in § 2.2, i.e. describable by (1) & (2):

$$(1) \rightarrow \underline{E}_I = E_I(0, 1, 0) e^{-i\omega t + ik(x \cos \theta + z \sin \theta)}$$

$$(2) \rightarrow \underline{B}_I = \frac{E_I}{c} (-\sin \theta, 0, \cos \theta) e^{-i\omega t + ik(x \cos \theta + z \sin \theta)}$$

Seek the reflected wave also as in (3) - (4),

$$(3) \rightarrow \underline{E}_R = (E_x, E_y, E_z) e^{-i\omega_R t + ik_R(-x \cos \theta_R + z \sin \theta_R)}$$

$$(4) \rightarrow \underline{B}_R = (B_x, B_y, B_z) e^{-i\omega_R t + \dots}$$

with (different) unknowns E_x, E_y, \dots, B_z

Additionally, a "transmitted" wave is sought in $x > 0$:
 $(k_R = \omega_R/c_1, k_T = \omega_T/c_2)$

$$\underline{E}_T = (\hat{E}_x, \hat{E}_y, \hat{E}_z) e^{-i\omega_T t + ik_T(x \cos \theta_T + z \sin \theta_T)} \quad (5)$$

$$\underline{B}_T = (\hat{B}_x, \hat{B}_y, \hat{B}_z) e^{-i\omega_T t + ik_T(x \cos \theta_T + z \sin \theta_T)} \quad (6)$$

From (1.43) $\Leftrightarrow \underline{E}^{(1)} \times \underline{n} = \underline{E}^{(2)} \times \underline{n} \Leftrightarrow$ continuity of y and z-components of $\underline{E} \Rightarrow$ ($n \times n = 0$)

y-component: $\Rightarrow \omega_R = \omega_T = \omega \Rightarrow k_R = k = \frac{\omega}{c_1}$, $k_T = \frac{\omega}{c_2}$
 $c_1 = (\epsilon_1 \mu_1)^{-1/2}$, $c_2 = (\epsilon_2 \mu_2)^{-1/2}$
 $\Rightarrow \theta_R = \theta$ (as before).

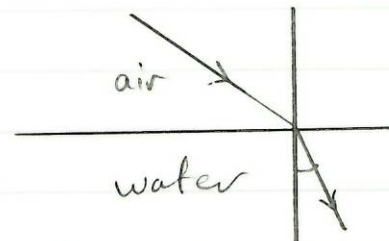
and additionally $k \sin \theta_T = k \sin \theta$

$$\Leftrightarrow \boxed{\frac{\sin \theta_T}{c_2} = \frac{\sin \theta}{c_1}} \quad \left(\frac{\sin \theta_1}{c_1} = \frac{\sin \theta_2}{c_2} \right) \quad (7)$$

Let $n := c_0/c$, c_0 speed of light (vacuum wave speed)
 $\Rightarrow n \geq 1$ ($n=1 \leftrightarrow$ vacuum) $\Rightarrow (7) \Leftrightarrow$

$$n_1 \sin \theta_1 = n_2 \sin \theta_2, \quad n \sin \theta = \text{const} \quad (8)$$

(7) or (8) is the Snell's law of reflection



Further, $E_I + E_y = \hat{E}_y$ (9)

$E_z = \hat{E}_z$ (10)

(z-component of \underline{E})

Also, $\underline{E}_R \cdot \underline{k}_R = 0$, $\underline{E}_T \cdot \underline{k}_T = 0$

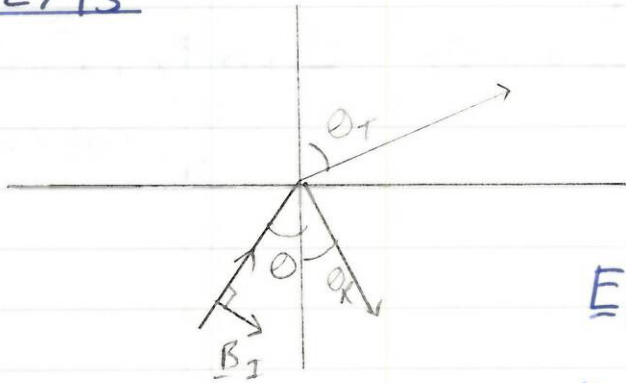
$$\Rightarrow -E_x \cos \theta + E_z \sin \theta = 0 \quad (11)$$

$$\hat{E}_x \cos \theta_T + \hat{E}_z \sin \theta_T = 0 \quad (12)$$

Further, from (1.46), and no surface currents for the two insulators,

$$\frac{1}{\mu_1} (\underline{B}_I + \underline{B}_R) \times \underline{n} = \frac{1}{\mu_2} \underline{B}_T \times \underline{n} \quad (*)$$

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$$\frac{\sin \theta}{c_1} = \frac{\sin \theta_T}{c_2} \quad (7)$$

$$\underline{E}_R = (E_x, E_y, E_z) e^{i \dots}$$

$$\underline{E}_T = (\hat{E}_x, \hat{E}_y, \hat{E}_z) e^{i \dots}$$

$$\underline{E}_I = (0, E_I, 0)$$

$$E_I + E_y = \hat{E}_y \quad (9)$$

$$\dots \quad (10, 11, 12)$$

$$\frac{1}{\mu_1} (\underline{B}_I + \underline{B}_R) \times \underline{u} = \frac{1}{\mu} \hat{\underline{B}}_T \times \underline{u} \quad (*)$$

$$(1.39) \Rightarrow \underline{B}_R = \frac{1}{\omega} \underline{k}_R \times \underline{E}_R$$

$$= \frac{k}{\omega} \begin{vmatrix} \underline{i} & \underline{j} & \underline{k} \\ \cos \theta & 0 & \sin \theta \\ E_x & E_y & E_z \end{vmatrix} e^{-i\omega t + i\underline{k}_R \cdot \underline{x}}$$

$$= \frac{1}{c_1} (-E_y \sin \theta, E_z \cos \theta + E_x \sin \theta, -E_y \cos \theta) e^{i \dots}$$

$$\underline{B}_T = \frac{1}{\omega} \underline{k}_T \times \underline{E}_T = \frac{k_T}{\omega} \begin{vmatrix} \underline{i} & \underline{j} & \underline{k} \\ \cos \theta_T & 0 & \sin \theta_T \\ \hat{E}_x & \hat{E}_y & \hat{E}_z \end{vmatrix}$$

$$= \frac{1}{c_2} \left(-\hat{E}_z \sin \theta_T, \hat{E}_x \sin \theta_T - E_z \cos \theta_T, \hat{E}_y \cos \theta_T \right) \cdot e^{-i\omega t + i\mathbf{k}_T \cdot \mathbf{x}}$$

where $\mathbf{x} = \begin{pmatrix} x \\ y \\ z \end{pmatrix}$

y-component of (*):

$$\frac{1}{\mu_1 c_1} (E_z \cos \theta + E_x \sin \theta) = \frac{1}{\mu_2 c_2} (\hat{E}_x \sin \theta_T - \hat{E}_z \cos \theta_T) \quad (13)$$

z-comp of (*):

$$\frac{1}{\mu_1 c_1} (E_x - E_y) \cos \theta = \frac{1}{\mu_2 c_2} \hat{E}_y \cos \theta_T \quad (14)$$

So (9)-(14) are 6 eqns for $E_x, E_y, \dots, \hat{E}_y, \hat{E}_z$.
Notice (10), (11), (12), (13) are eqns for $E_x, E_z, \hat{E}_x, \hat{E}_z$ only.

$$(13) \Rightarrow \frac{1}{\mu_1 c_1} \left(E_z \cos \theta + \frac{E_z \sin \theta}{\cos \theta} \sin \theta \right)$$

$$= \frac{1}{\mu_2 c_2} \left(-\frac{\hat{E}_z \sin^2 \theta_T}{\cos \theta_T} - E_z \cos \theta_T \right)$$

$$\Leftrightarrow \frac{1}{\mu_1 c_1 \cos \theta} E_z = - \frac{1}{\mu_1 c_2 \cos \theta_T} E_z \Rightarrow E_z = 0.$$

$$\Rightarrow E_x = \hat{E}_x = \hat{E}_z = E_z = 0.$$

So, like the incident electric field E_I , both E_R and E_T have only y-component (same polarisation)

↔ "transverse - electric" ↔ TE wave.

So we have left (9), (14) for E_y, \hat{E}_y :

$$(9) \rightarrow E_I + E_y = \hat{E}_y$$

$$\frac{1}{\mu_1 c_1} (E_I - E_y) \cos \theta = \frac{1}{\mu_2 c_2} \hat{E}_y \cos \theta_T. \quad (14)$$

$$(9) \rightarrow (14) \Rightarrow$$

$$\boxed{\frac{\cos \theta}{\mu_1 c_1}} (E_I - E_y) = \boxed{\frac{\cos \theta_T}{\mu_2 c_2}} (E_I + E_y)$$

!! m_1
!! m_2

$$m_1 E_I - m_1 E_y = m_2 E_I + m_2 E_y$$

$$\Rightarrow E_y = \frac{m_1 - m_2}{m_1 + m_2} E_I, \quad m_1 = \frac{\cos \theta}{\mu_1 c_1}, \quad m_2 = \frac{\cos \theta_T}{\mu_2 c_2}$$

$$\hat{E}_y = E_I + E_y = \frac{2m_1}{m_1 + m_2} E_I.$$

If both media 1 and 2 are "non-magnetic" (as often physically the case) i.e. $\mu_1 = \mu_2 = \mu_0$.

$$(\text{but } \epsilon_1 \neq \epsilon_2) \Rightarrow (n_1 = c_0/c_1, n_2 = c_0/c_2)$$

$$E_y = \frac{n_1 \cos \theta - n_2 \cos \theta_T}{n_1 \cos \theta + n_2 \cos \theta_T} E_I \quad (15)$$

$$\hat{E}_y = \frac{2n_1 \cos \theta}{n_1 \cos \theta + n_2 \cos \theta_T} E_I \quad (16)$$

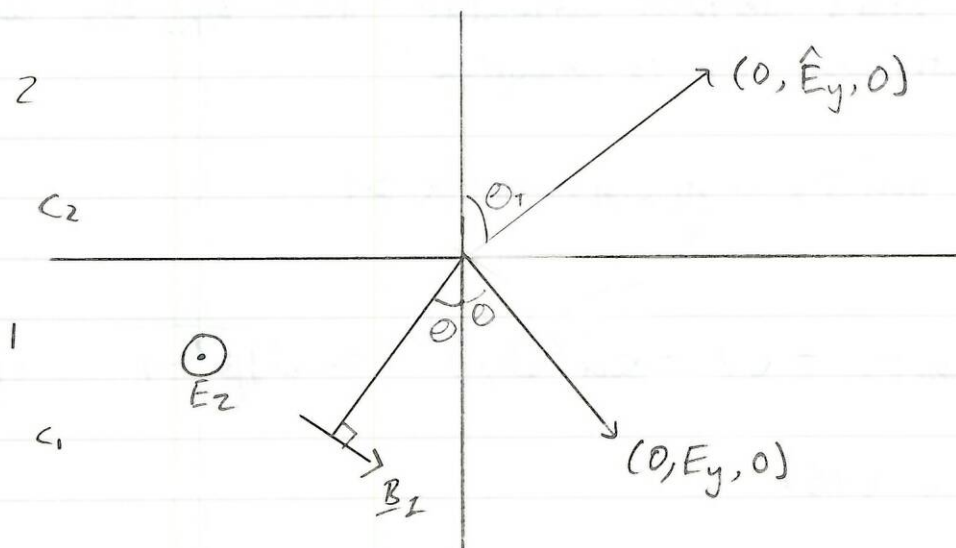
Now via Snell's law (8), $n_1 \sin \theta = n_2 \sin \theta_T$

$$E_y = \frac{\sin \theta_T \cos \theta - \sin \theta \cos \theta_T}{\sin \theta_T \sin \theta + \sin \theta \cos \theta} E_I$$

$$= \frac{\sin(\theta_T - \theta)}{\sin(\theta_T + \theta)} E_I \quad (17)$$

$$\hat{E}_y = \frac{2 \sin \theta_T \cos \theta}{\sin(\theta_T + \theta)} E_I \quad (18)$$

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$$\frac{\sin \theta}{c_1} = \frac{\sin \theta_r}{c_2} \quad (7)$$

$$(E_y, \hat{E}_y) \leftrightarrow (17) - (18) / (15) - (16)$$

After $\underline{E}_R, \underline{E}_T$ found \Rightarrow by (1.39)

$$\underline{B}_R = \frac{1}{\omega} \underline{k}_R \times \underline{E}_R, \quad \underline{B}_T = \frac{1}{\omega} \underline{k}_T \times \underline{E}_T.$$

Total internal reflection

$$\text{Snell's law (7)} \Rightarrow \sin \theta_r = \frac{c_2}{c_1} \sin \theta = \frac{n_1}{n_2} \sin \theta.$$

($n_j = c_0/c_j$ $j=1, 2$ refractive index).

Let $c_2 > c_1$ and $c_2/c_1 \sin \theta > 1 \Leftrightarrow$

$$\theta > \sin^{-1}\left(\frac{c_1}{c_2}\right) =: \theta_c = \text{"the critical angle"}.$$

All the above derived formulae for E_R, E_T are still valid, however θ_T is complex:

$$\sin \theta_T = \frac{n_1}{n_2} \sin \theta =: \alpha > 1.$$

$$\Rightarrow \cos \theta_T = (1 - \sin^2 \theta_T)^{1/2} = i\sqrt{\beta^2 - 1} = i\beta, \beta > 0.$$

\Rightarrow from (15)

$$E_y = \frac{n_1 \cos \theta - n_2 \cos \theta_T}{n_1 \cos \theta + n_2 \cos \theta_T} E_I = \frac{n_1 \cos \theta - i\beta n_2}{n_1 \cos \theta + i\beta n_2} E_I.$$

$\Rightarrow |E_y| = |E_I| \leftrightarrow$ reflected & incident amplitudes same
 \Rightarrow All the energy is reflected. (phases differ \rightarrow phase change upon the reflection).

< For $\theta < \theta_c$, (15) $\Rightarrow |E_y| < |E_I|$
 \leftrightarrow reflected energy less than transmitted. >

For $\theta > \theta_c$, for $x > 0$:

$$\underline{k}_T = k_T (\cos \theta_T, 0, \sin \theta_T)$$

$$= \frac{\omega}{c_2} (i\beta, 0, \alpha)$$

$$E_T = \text{Re} \left\{ \hat{E}_y e^{i\mathbf{k}_T \cdot \mathbf{x} - i\omega t} \right\}$$

$$= \text{Re} \left\{ (0, \hat{E}_y, 0) e^{i k_T \alpha z - k_T \beta x - i\omega t} \right\}$$

$$\begin{aligned} & \text{with } E_y = |E_y| e^{i\psi} \\ & = (0, |E_y| e^{-k_T \beta x} \cos(\omega t - k_T \alpha z - \Psi), 0) \end{aligned}$$

(so \underline{E}_T exponentially decays as $x \rightarrow +\infty$)

$$(1.39) \Rightarrow \underline{B}_T = \frac{1}{\omega} \underline{k}_T \times \underline{E}_T$$

$$= \text{Re} \left\{ \begin{array}{c|c|c} \frac{k_T}{\omega} & \begin{array}{c} i \\ i\beta \\ 0 \end{array} & \begin{array}{c} j \\ 0 \\ \hat{E}_y \end{array} \\ \hline & & \alpha \\ \hline & & 0 \end{array} \middle| e^{i(k_T \alpha z - k_T \beta x - \omega t)} \right\}$$

$$= \frac{e^{-k_T \beta x}}{c_2} \left(-\alpha |\hat{E}_y| \cos(\omega t - k_T \alpha z - \Psi), 0, \beta |\hat{E}_y| \sin(\omega t - k_T \alpha z - \Psi) \right)$$

Now for Poynting vector \underline{S}_T (§1.2) for transmitted wave,

$$\underline{S} = \frac{1}{\mu} \underline{E}_T \times \underline{B}_T$$

$$= \frac{e^{-2k_T \beta x}}{c_2 \mu} |\hat{E}_y|^2 \begin{array}{c|c|c} i & j & k \\ \hline 0 & \cos \phi & 0 \\ \hline -\alpha \cos \phi & 0 & \beta \sin \phi \end{array}$$

$$(\phi := \omega t - k_T \alpha z - \Psi)$$

$$\Rightarrow \langle \underline{S} \rangle_T = \frac{e^{-2k_T \beta x} |\hat{E}_y|^2}{\mu c_2} \left(\beta \langle \sin \phi \cos \phi \rangle_T, 0, \alpha \langle \cos^2 \phi \rangle_T \right)$$

$$\left(\langle \underline{S} \rangle_T := \frac{1}{T} \int_0^T \underline{S}(\underline{x}, t) dt, T = \frac{2\pi}{\omega}, \text{Time averaged } \underline{S} \right)$$

$$= \frac{\alpha e^{-2\beta k_T x} |\hat{E}_y|^2}{2\mu c_2} (0, 0, 1)$$



$$\left(\langle \sin \phi \cos \phi \rangle_T = 0, \langle \cos^2 \phi \rangle = \frac{1}{2} \right)$$

So the EM energy (time averaged one) flows along $z \leftrightarrow$ parallel to the interface S . (So for "transmitted" wave. No energy transport away from S .)

Exercise: Similar analysis can be done for incident plane wave polarised "in x, z -plane" $\leftrightarrow \underline{B}_I(0, B_I, 0) \leftrightarrow$ "transverse magnetic" / TM (cf TE / transverse electric case above).

All same affects / $\theta_R = \theta$, θ_T from Snell's law (7), total internal refln for $\theta > \theta_c$)

Difference! (17) - (18) \rightarrow

$$B_y = \frac{\sin 2\theta - \sin 2\theta_T}{\sin 2\theta + \sin 2\theta_T} B_I.$$

$$\hat{B}_y = \frac{2 \sin 2\theta}{\sin 2\theta + \sin 2\theta_T} B_I,$$

$$\begin{aligned} & \sin \alpha - \sin \beta \\ &= 2 \cos \left(\frac{\alpha + \beta}{2} \right) \sin \left(\frac{\alpha - \beta}{2} \right) \end{aligned}$$

For $\theta = \theta_B$ (the "Brewster angle")

$$B_y = 0, (\Leftrightarrow) \sin 2\theta_B = \sin 2\theta_T (\Leftrightarrow)$$

$$2 \sin \theta \cos \theta = 2 \sin \theta_T \cos \theta_T \text{ etc. . . .}$$

$$\text{using (7)} \quad \theta_B = \tan^{-1} \left(\frac{c_1}{c_2} \right)$$

So for $\theta = \theta_B$ No reflected wave! (Everything is transmitted); Exam 2010 Q2

□

Remark 2: Any other plane T-H EM wave incident upon S can be decomposed into combination of above TE & TM waves. (Since TE & TM have different reflection/transmission rules, e.g. for $\theta \simeq \theta_B \leftrightarrow$ "polarisation filtering effect".

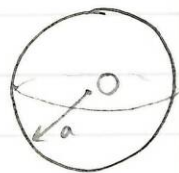
By (7), θ_T depends on C ; often C depends on ω , $c = c(\omega) \Rightarrow \theta_T = \theta_T(\omega) \leftrightarrow$ "dispersion" effect; this is how a ^{$n = n(\omega)$} rainbow or a prism works.

□

2.4 Waves due to spherical sources (acoustics)

Here some fundamental example of radially propagating sound are examined, due to "spherical sources".

2.4.1 A radially pulsating sphere



Consider a sphere centred at origin O which "pulsates": its radius a changes with time, $a = a(t)$.

We assume $|a(t) - a_0| \ll a_0$, a_0 "original" radius, i.e. the perturbations are small (c.f. §1.1)

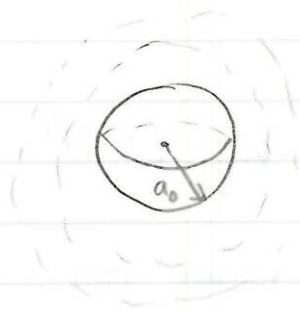
We then expect that an acoustic wave will be produced for $|x| > a_0$ (the sphere is surrounded by acoustic medium with wavespeed c , density ρ_0). Assume

also $a(t) \equiv a_0$ for $t \leq t_0$, cf Causality (§1.11).
The solution is then expected to be radially-symmetric (& causal) $\Rightarrow \underline{u}(x,t) = \nabla \phi$, with
velocity potential $\phi = \phi(r, t)$

$$\phi(r, t) = \frac{f\left(t - \frac{r}{c}\right)}{r},$$

$f = f(\xi)$ to be found from the boundary conditions at $r = a(t) \approx a_0$, $r \approx a_0$.

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$$a(t) \rightarrow v_s = a'(t)$$

$$\rightarrow \phi = \frac{f(t - \frac{r}{c})}{r}$$

$$\Rightarrow \underline{u} = \nabla \phi = \left(-\frac{f(t - \frac{r}{c})}{r^2} - \frac{f'(t - \frac{r}{c})}{cr} \right) \hat{r}$$

Seek:

(unit radial vector)

$$f(t - \frac{r}{c}) = -ca_0 \Psi(t - \frac{r}{c} + \frac{a_0}{c})$$

(Ψ still an arbitrary function, so far)

Boundary condns (BCs): $r = a_0 \Rightarrow \underline{u} = a'(t) \hat{r}$

$$\Leftrightarrow \left. \frac{c}{e^{\frac{c}{a_0} t}} \right|_{a_0} \Psi(t) + \Psi'(t) = a'(t) = v_s(t) \quad (19)$$

\Rightarrow Solving (19) for $\Psi(t)$, using integrating factor $e^{\frac{c}{a_0} t}$

$$\Rightarrow \frac{d}{dt} \left(e^{\frac{c}{a_0} t} \Psi(t) \right) = e^{\frac{c}{a_0} t} v_s(t)$$

$$\Rightarrow e^{\frac{c}{a_0} t} \Psi(t) = \int_{-\infty \text{ or } t_0}^t e^{\frac{c}{a_0} \tau} v_s(\tau) d\tau$$

($v_s = \Psi(t) = 0, t < t_0$ by causality)
 $v_s(t) = 0$

$$\Rightarrow \psi(t) = \int_{t_0}^t e^{-\frac{c}{a}(t-\tau)} v_s(\tau) d\tau.$$

$\psi \rightarrow \phi \rightarrow \underline{u}$ can be found \mapsto by (1.11)

$$p = -\rho_0 \frac{\partial \phi}{\partial t} \text{ pressure can be found too.}$$

For T-H case, $v_s(t) = \operatorname{Re}\left\{ \hat{v}_s e^{-i\omega t} \right\}$ $\hat{v}_s \in \mathbb{C}$.

Seek $\psi(t) = \operatorname{Re}\left\{ B e^{-i\omega t} \right\}$ $B \in \mathbb{C}$ to be found.

Plugging to (19):

$$\frac{c}{a_0} \psi + \psi' = v_s(t) \Rightarrow \frac{c}{a} B e^{-i\omega t} - i\omega B e^{-i\omega t} = \hat{v}_s e^{-i\omega t}$$

$$\Rightarrow B = \frac{a_0 \hat{v}_s}{c(1 - \frac{i\omega a_0}{c})} = \frac{a_0 \hat{v}_s}{c(1 - ika_0)}$$

$$\phi = \operatorname{Re}\left\{ \frac{f(t - \frac{r}{c})}{r} \right\} = \operatorname{Re}\left\{ -\frac{ca_0}{r} \psi\left(t - \frac{r}{c} + \frac{a_0}{c}\right) \right\}$$

$$= \operatorname{Re}\left\{ -\frac{ca_0}{r} B e^{-i\omega(t - \frac{r}{c} + \frac{a_0}{c})} \right\}$$

$$= \operatorname{Re}\left\{ -\frac{a_0^2 \hat{v}_s}{1 - ika_0} x \frac{e^{ikr - ika_0 - i\omega t}}{r} \right\}$$

From (1.41) $\Leftrightarrow p = -i\omega\rho_0\phi$

$$= \operatorname{Re} \left\{ -\frac{i\omega a_0^2 \hat{v}_s \rho_0 e^{ik(r-a_0)-i\omega t}}{(1-ika_0)r} \right\} \quad (20)$$

$$p = \operatorname{Re} \left\{ \frac{A}{r} e^{ikr-i\omega t} \right\}, \quad A = -\frac{i\omega a_0^2 \hat{v}_s \rho_0 e^{-ika_0}}{1-ika_0}$$

So, for time-averaged intensity,

$$\langle \underline{I} \rangle_T = \langle I_r \rangle_T \hat{r}, \quad (1.22)$$

$$\Rightarrow \langle I_r \rangle = \frac{|A|^2}{2\rho_0 c r^2}$$

$$\begin{aligned} \Rightarrow \langle I_r \rangle &= \frac{\omega^2 a_0^4 |\hat{v}_s|^2 \rho_0 c}{(1+(ka_0)^2) 2c^2 r^2} \\ &= \frac{(ka)^2 a_0^2 |\hat{v}_s|^2 c \rho_0}{2(1+k^2 a_0^2) r^2} \end{aligned}$$

\Rightarrow The average radiated power:

$$P = \langle I_r \rangle \times 4\pi r^2$$

$$= \frac{2\pi (ka_0)^2 a_0^2 |\hat{v}_s|^2 c \rho_0}{1+(ka_0)^2}$$

In "low frequency" regime $\Leftrightarrow ka_0 \ll 1$

$$P_{\text{av}} \approx 2\pi (ka_0)^2 a_0^2 |\hat{v}_s|^2 c \rho_0 \quad (21)$$

i.e "of order" $(ka_0)^2 = \left(\frac{\omega}{c} a_0\right)^2$ w.r.t ω .

□

$$(1.1) \quad \langle \underline{I} \rangle = \langle \underline{I} \rangle$$

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2.4.2 Transversely oscillating rigid sphere

Let sphere of radius a be rigid (i.e. $a = \text{constant}$) and oscillate along z -axis, i.e. its centre (and any other point) has velocity

$$v_c(t) = v_c(t) \hat{z} \quad (\hat{z} \text{ unit vector along } z)$$

Oscillations are assumed small. Seek, for $r > a$, acoustic velocity $\underline{u} = \nabla \phi$.

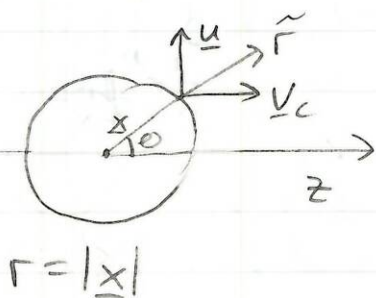
$$\phi = \frac{\partial}{\partial z} \left[\frac{1}{r} \Psi \left(t - \frac{r}{c} + \frac{a}{c} \right) \right] \quad (22)$$

<Check: $\frac{1}{r} \Psi \left(t - \frac{r}{c} + \frac{a}{c} \right)$ solves wave eqn, (1.11)

$$\Rightarrow \frac{\partial}{\partial z} (\nabla^2 + k^2) \left(\frac{1}{r} \Psi \right) = 0.$$

$$\Rightarrow (\nabla^2 + k^2) \left[\frac{\partial}{\partial z} \left(\frac{1}{r} \Psi \right) \right] = 0 \Leftrightarrow \phi \text{ also solves (1.11)}$$

Boundary conditions (BCs): At $r = a$ (on the sphere's surface) acoustic medium's velocity relative to the sphere must be parallel to S :



$$(\underline{u} - \underline{v}_c) \cdot \underline{\hat{x}} = 0$$

$$\Rightarrow \underline{u} \cdot \underline{\hat{x}} = \underline{v}_c \cdot \underline{\hat{x}}$$

$$\Leftrightarrow \frac{\partial \phi}{\partial r} = \nabla \phi \cdot \underline{\hat{r}} = v_c \cos \theta$$

(θ angle between \underline{x} and z -axis)

$$\Rightarrow \frac{\partial \phi}{\partial r} = v_c(t) \cos \theta, \quad r = a \quad (23)$$

is the BC.

Notice that $\frac{\partial f(r)}{\partial z} = \frac{\partial r}{\partial z} \frac{\partial f}{\partial r}$,

$$r = (x^2 + y^2 + z^2)^{1/2} \Rightarrow \frac{\partial r}{\partial z} = \frac{z}{r} = \cos \theta$$

$$\Rightarrow \frac{\partial f}{\partial z} = \cos \theta \frac{\partial f}{\partial r}$$

$$\Rightarrow \phi = \cos \theta \frac{\partial}{\partial r} \left[\frac{1}{r} \Psi \left(t - \frac{r}{c} + \frac{a}{c} \right) \right]$$

From (23)

$$\Rightarrow \cancel{\cos \theta} \frac{\partial^2}{\partial r^2} \left[\frac{1}{r} \Psi \left(t - \frac{r}{c} + \frac{a}{c} \right) \right]_{r=a} = v_c(t) \cancel{\cos \theta} \quad (23')$$

$$\text{Now } \frac{\partial}{\partial r} \left(\frac{1}{r} \Psi \left(t - \frac{r}{c} + \frac{a}{c} \right) \right)$$

$$= -\frac{1}{r^2} \Psi - \frac{1}{cr} \Psi' \left(t - \frac{r}{c} + \frac{a}{c} \right)$$

$$\Rightarrow \frac{\partial^2}{\partial r^2} \left(\frac{1}{r} \Psi \right) = \frac{2}{r^3} \Psi \left(t - \frac{r}{c} + \frac{a}{c} \right)$$

$$+ \frac{2}{cr^2} \Psi' + \frac{1}{c^2 r} \Psi''$$

So (23')

$$\Rightarrow \frac{2}{a^3} \Psi(t) + \frac{2}{ca^2} \Psi'(t) + \frac{1}{c^2 a} \Psi'' = v_c(t)$$

$$2c^2 \Psi + 2ca \Psi' + a^2 \Psi'' = c^2 a^3 v_c(t) \quad (24)$$

which is an ODE for $\Psi(t)$.

Restricting to T-H case, Let

$$v_c(t) = \operatorname{Re} \left\{ \hat{v}_c e^{-i\omega t} \right\}, \quad \hat{v}_c \in \mathbb{C}.$$

Seek $\Psi(t)$ also as $\Psi(t) = \operatorname{Re} [A e^{-i\omega t + ika}]$, $A \in \mathbb{C}$
Then from (24)

$$\begin{aligned} \Rightarrow (2c^2 A - 2i\omega ca A - \omega^2 a^2 A) e^{-i\omega t + ika} \\ = c^2 a^3 \hat{v}_c e^{-i\omega t} \end{aligned}$$

$$\Rightarrow A = \frac{c^2 a^3 \hat{V}_c e^{-ika}}{2c^2 - 2i\omega ca - \omega^2 a^2}$$

$$\left(\frac{\omega}{c} = k\right)$$

$$\Rightarrow A = \frac{a^3 \hat{V}_c e^{-ika}}{2 - 2ika - (ka)^2} \quad (25)$$

$$\Rightarrow \Psi(t) = A e^{-i\omega t + ika} \Rightarrow \text{from (22)}$$

$$\phi = (\text{Re}) \cos \theta \frac{\partial}{\partial r} \left[\frac{A}{r} e^{-i\omega(t - \frac{r}{c} + \frac{a}{c}) + ika} \right]$$

$$= \text{Re} \cos \theta \frac{\partial}{\partial r} \left[\frac{A}{r} e^{-i\omega t + ikr - \cancel{ika} + \cancel{ika}} \right]$$

$$\phi = (\text{Re}) \left(\cos \theta \frac{\partial}{\partial r} \left[\frac{A}{r} e^{ikr - i\omega t} \right] \right) \quad \hat{=} \hat{\phi}$$

$$\phi = \text{Re}(\hat{\phi} e^{-i\omega t}), \quad \hat{\phi} = \cos \theta \frac{\partial}{\partial r} \left(\frac{A}{r} e^{ikr} \right) \quad (26)$$

$$\hat{p} = i\omega \rho_0 \hat{\phi} = i\omega \rho_0 \cos \theta \frac{A}{r} e^{ikr} \left(-\frac{1}{r} + ik \right)$$

$$= -k\omega \rho_0 \cos \theta \frac{A}{r} e^{ikr} \left(1 + \frac{i}{kr} \right)$$

\Rightarrow In the "far field" ($\Leftrightarrow kr \gg 1$):

$$\hat{p} = -k^2 c \rho_0 \cos \theta \frac{A}{r} e^{ikr} \quad (28)$$

For radial velocity $u_r := \underline{u} \cdot \hat{r} = \frac{\partial \phi}{\partial r}$

$$= \operatorname{Re} \left[\cos \theta \frac{\partial^2}{\partial r^2} \left(\frac{A}{r} e^{ikr} \right) \right];$$

In the far field:

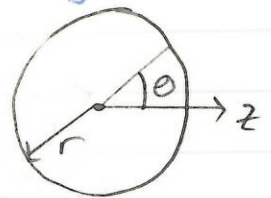
$$u_r = \operatorname{Re} \left\{ \cos \theta \frac{A}{r} e^{ikr} \left(-k^2 + \frac{2}{r} - \frac{2ik}{r} \right) \right\}$$

$$\approx -\frac{k^2 \cos \theta}{r} \operatorname{Re} \left(A e^{ikr - i\omega t} \right)$$

For time-averaged radial acoustic intensity

$$\langle I_r \rangle = \langle p u_r \rangle$$

$$\approx \frac{k^4 c \rho_0 \cos^2 \theta |A|^2}{r^2} \langle \cos^2(\omega t - kr - \psi) \rangle_T$$



$$A = |A| e^{i\psi}$$

$$= \frac{k^4 c \rho_0 \cos^2 \theta |A|^2}{2r^2}$$

(in the far field, $kr \gg 1$)

The average radiated power: (S_r ; sphere of large radius r)

$$P_{av} = \iint_{S_r} \langle I_r \rangle dS_r$$

$$= \frac{k^4 c \rho_0 |A|^2}{2r^2} \int_0^{2\pi} \int_0^\pi \underbrace{\cos^2 \theta r^2 \sin \theta}_{dS} d\theta d\varphi$$

$$= \pi k^4 \rho_0 |A|^2 \left[-\frac{1}{3} \cos^3 \theta \right]_0^\pi$$

$$= \frac{2}{3} \pi \rho_0 |A|^2 k^4$$

$$(25) \Rightarrow |A|^2 = \frac{a^6 |\hat{V}_c|^2}{(2 - (ka)^2)^2 + 4k^2 a^2}$$

$$= \frac{a^6 |\hat{V}_c|^2}{4 + (ka)^4}$$

$$\Rightarrow P_{av} = \frac{2}{3} \pi \rho_0 k^4 \frac{a^6 |\hat{V}_c|^2}{4 + (ka)^4}$$

$$= \frac{2}{3} \pi \frac{(ka)^4 \rho_0 a^2 |\hat{V}_c|^2}{4 + (ka)^4} \quad (29)$$

In "low-frequency" regime, $ka \ll 1$

$$P_{av} \approx \frac{\pi}{6} (ka)^4 \rho_0 a^2 |\hat{V}_c|^2 \quad (30)$$

So the radiated power is "order 4" with respect to frequency.

Notice also (25) \rightarrow (low frequency, $ka \ll 1$)

$$A = \frac{1}{2} a^3 \hat{V}_c$$

So (28)

$$\Rightarrow \hat{p} \approx -k^2 \rho_0 c A \cos \theta \frac{e^{ikr}}{r}$$


$$\hat{p} \approx -\frac{1}{2} \rho_0 c \hat{v}_c (ka)^2 a \cos \theta \frac{e^{ikr}}{r} \quad (31)$$

□

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3. Green's function & Multipole expansion.

3.1 Monopoles & free-space Green's function.

 (a_0, \hat{V}_s) - We've seen that any spherically symmetric TH source produces a (sph-symmetric) outgoing TH wave of the form $p = \text{Re}(\hat{p}(r)e^{-i\omega t})$,

$$\hat{p} = \hat{S} \frac{e^{ikr}}{r}, \quad r > a_0; \quad (1)$$

$$\hat{S} \in \mathbb{C}$$

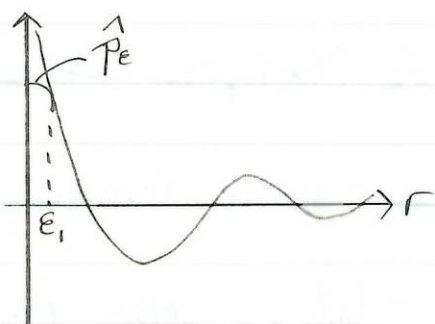
Definition: \hat{S} is referred to as Monopole amplitude.

For radially pulsating sphere (2.20) \Rightarrow

$$\hat{S} = - \frac{i\omega \rho_0 a_0^2 \hat{V}_s e^{-ika_0}}{1 - ika_0}$$

Let $a_0 \rightarrow 0$ and \hat{V}_s increases, so that \hat{S} remains constant. In the "limit" we have $\hat{p}(r)$ defined by (1) everywhere except $r=0$, with a "singularity" at $r=0$.

Such \hat{p} is called a "point source." (In reality a source can be regarded as a point source if
(i) $ka \ll 1$ ($\Leftrightarrow a_0 \ll \lambda$)
(ii) observation point $r \gg a_0$)



For a small $\epsilon > 0$, let $\hat{p}_\epsilon(r) := \hat{p}(r)$ for $r \geq \epsilon$, but \hat{p}_ϵ is "smoothed out" for $r < \epsilon \Rightarrow$

As $(\nabla^2 + k^2)\hat{p} = 0$, $r \neq 0$, $(p = \hat{S} \frac{e^{ikr}}{r})$

$$(\nabla^2 + k^2)\hat{p}_\epsilon = -\hat{S} \Delta_\epsilon(r) \quad (2)$$

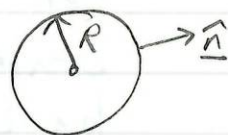
where the RHS $\Delta_\epsilon(r) \neq 0$ only for $r < \epsilon$ (\Leftrightarrow has support in ball B_ϵ). Let $\epsilon \rightarrow 0$, fix $R > 0$ and integrate (2) over ball centred at 0 of radius R , B_R :

$$\int_{B_R} \nabla^2 \hat{p}_\epsilon dV + \int_{B_R} k^2 \hat{p}_\epsilon dV = -\hat{S} \int_{B_R} \Delta_\epsilon(r) dV$$

$$\Leftrightarrow I_1 + I_2 = I_3$$

$$I_1 = \int_{B_R} \nabla \cdot (\nabla \hat{p}_\epsilon) dV \stackrel{\text{Divergence theorem}}{=} \int_{S_R} \nabla \hat{p}_\epsilon \cdot \underline{n} dS$$

$S_R \leftarrow$ sphere of radius R



$$= \int_{S_R} \left(\frac{\partial}{\partial r} \right) \hat{p}_\epsilon dS = \int_{S_R} \frac{\partial}{\partial r} \left(\hat{S} \frac{e^{ikr}}{r} \right) dS$$

$$= \hat{S} \int e^{ikr} \left(\frac{ik}{r} - \frac{1}{r^2} \right) dS = \hat{S} e^{ikR} \left(\frac{ik}{R} - \frac{1}{R^2} \right) \times 4\pi R^2$$

$$= 4\pi \hat{S} e^{ikR} (ikR - 1)$$

$$I_2 = \int_{B_R} k^2 \hat{p}_E dV \xrightarrow{E \rightarrow 0} \int_{B_R} k^2 \hat{p} dV$$

$$= \hat{S} k^2 \int_0^{2\pi} \int_0^\pi \int_0^R \frac{e^{ikr}}{r} \underbrace{r^2 \sin\theta dr d\theta d\varphi}_{dV}$$

$$\int_0^{2\pi} \int_0^\pi \int_0^R \sin\theta d\theta d\varphi = 4\pi \quad (\text{= area of sphere of radius 1})$$

$$\Rightarrow I_2 = 4\pi k^2 \hat{S} \int_0^R \frac{\partial}{\partial r} e^{ikr} r dr$$

$$= -4\pi i k \hat{S} \int_0^R \frac{\partial}{\partial r} (e^{ikr}) r dr$$

$$= -4\pi i k \hat{S} e^{ikR} R + 4\pi i k \hat{S} \int_0^R \frac{\partial}{\partial r} (e^{ikr}) dr$$

$$= -4\pi i \hat{S} e^{ikR} kR + 4\pi \hat{S} (e^{ikR} - 1)$$

$$= 4\pi \hat{S} e^{ikR} (1 - ikR) - 4\pi \hat{S}$$

$$\Rightarrow I_1 + I_2 \xrightarrow{E \rightarrow 0} 4\pi \hat{S} e^{ikR} (ikR - 1 + 1 - ikR) - 4\pi \hat{S}$$

$$= -4\pi \hat{S}$$

$$\Rightarrow I_3 = - \int_{B_R} \Delta_\epsilon dV \xrightarrow{\epsilon \rightarrow 0} -4\pi \delta$$

Hence, denoting $\delta_\epsilon := \frac{1}{4\pi} \Delta_\epsilon$, we observe:

$$(i) \int \delta_\epsilon dV \rightarrow 1 \text{ as } \epsilon \rightarrow 0.$$

(ii) $\delta_\epsilon \neq 0$ only within $r < \epsilon$.

This implies

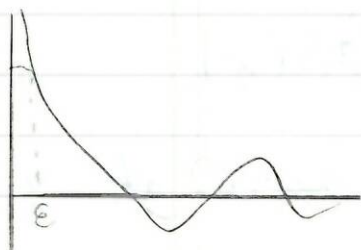
$$\lim_{\epsilon \rightarrow 0} \delta_\epsilon(x) = \delta(x)$$

called "Dirac delta-function", with the "limit" understood appropriately.

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$$\hat{p} = \hat{S} \frac{e^{ikr}}{r}, \quad r > 0 \quad (1)$$



$$(\nabla^2 + k^2) \hat{p}_\epsilon = \Delta_\epsilon$$

$$\delta_\epsilon := \frac{1}{4\pi} \Delta_\epsilon, \quad \int_{\mathbb{R}^3} \delta_\epsilon \rightarrow 1$$

$$\lim_{\epsilon \rightarrow 0} \delta_\epsilon = \delta(\underline{x})$$



So $\delta(\underline{x})$ can be viewed as a "function", such that

$$(i) \quad \delta(\underline{x}) = \begin{cases} 0 & \underline{x} \neq 0 \\ +\infty & \underline{x} = 0 \end{cases}$$

$$(ii) \quad \int_{\mathbb{R}^3} \delta(\underline{x}) dV(\underline{x}) = 1$$

< Rigorous defn' requires more advanced Analysis \leftrightarrow "distribution theory" >

One important property of δ -function, often called "sifting property": \forall "reasonable" (e.g. continuous) function f

$$\int_{(\mathbb{R}^3)} \delta(\underline{x} - \underline{x}') f(\underline{x}') dV(\underline{x}') = \int_{(\mathbb{R}^3)} \delta(\underline{x}' - \underline{x}) f(\underline{x}') dV(\underline{x}') = f(\underline{x}) \quad (3)$$

(follows from "definition" of f)

Hence, for \hat{p} given by (1), as $\epsilon \rightarrow 0$,

$$(\nabla^2 + k^2)\hat{p} = -4\pi \hat{S} \delta(\underline{x})$$

More generally, for a point source at \underline{x}_s ,

$$\hat{p}(\underline{x}) = \hat{S} \frac{e^{ikR}}{R}, \quad R := |\underline{x} - \underline{x}_s| \quad (1')$$

$$\Rightarrow (\nabla^2 + k^2)\hat{p} = -4\pi \hat{S} \delta(\underline{x} - \underline{x}_s) \quad (4)$$

Green's functions: The solution of (4) with RHS $\delta(\underline{x} - \underline{x}_s)$ i.e. $\hat{S} = -1/4\pi$, is called free-space Green's function.

$$G(\underline{x}, \underline{x}') = G_f(\underline{x}, \underline{x}')$$

$$(\nabla^2 + k^2)G(\underline{x}, \underline{x}') = \delta(\underline{x} - \underline{x}') \quad (5)$$

Hence, for Helmholtz

$$G_f(\underline{x}, \underline{x}') = -\frac{e^{ik|\underline{x} - \underline{x}'|}}{4\pi|\underline{x} - \underline{x}'|} \quad (6)$$

(Free space Green fns for general linear differential operators are often also called "fundamental solutions")

3.2 Green's functions

Green's fns play fundamental role mathematically. Consider a field due to arbitrary "distributed" source f :

$$\mathcal{L}p = (\nabla^2 + k^2)p = f(\underline{x}) \quad (7)$$

(In (7), \mathcal{L} is Helmholtz "differential operator"
 $\mathcal{L}: p \rightarrow f$; generally \mathcal{L} could be wave eqn, Maxwell's system related operator.

Normally (7) is supplemented by boundary/initial etc (e.g. Sommerfeld radiation condn) to form boundary-value problem (BVP), to determine p from f uniquely: symbolically $f = \mathcal{L}p \rightarrow p = \mathcal{L}^{-1}f$. Assume \mathcal{L} is invertible. Since \mathcal{L} is differential, its natural to seek \mathcal{L}^{-1} as an "integral operator"

$$p(\underline{x}) = (\mathcal{L}^{-1}f)(\underline{x}) = \int_{\Omega} K(\underline{x}, \underline{x}') f(\underline{x}') dV(\underline{x}')$$

Ω is "domain" of \mathcal{L} .

The kernel $K(\underline{x}, \underline{x}')$ is to be found. Now;

$$f(\underline{x}) = (\mathcal{L}p)(\underline{x}) = \mathcal{L}(\mathcal{L}^{-1}f)(\underline{x})$$

$$= \mathcal{L}_x \int K(\underline{x}, \underline{x}') f(\underline{x}') dV(\underline{x}')$$

$$= \int \underbrace{\mathcal{L}_x K(\underline{x}, \underline{x}')}_{\substack{\text{"G}(\underline{x}, \underline{x}') \\ \text{S}(\underline{x}-\underline{x}')}} f(\underline{x}') dV(\underline{x}') \stackrel{(*)}{=} f(\underline{x}).$$

Notice that taking $K(\underline{x}, \underline{x}') = G(\underline{x}, \underline{x}')$ serves the purpose: by (5) $\mathcal{L}_x G(\underline{x}, \underline{x}') = \delta(\underline{x} - \underline{x}')$, and by (3) the (*) then holds

$$\Rightarrow (\mathcal{L}^{-1}f)(\underline{x}) = \int G(\underline{x}, \underline{x}') f(\underline{x}') dV(\underline{x}'). \quad (8)$$

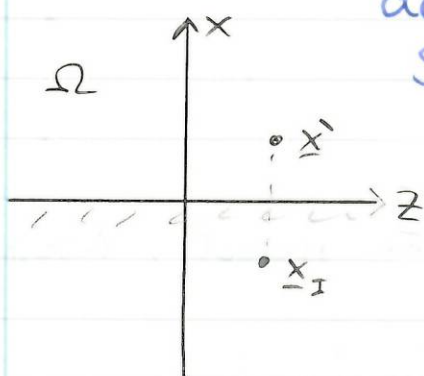
For Helmholtz eqn in domains with boundaries the Green's function is generally of the form:

$$G(\underline{x}, \underline{x}') = G_f(\underline{x}, \underline{x}') + G_B(\underline{x}, \underline{x}')$$

with G_f the free-space Green's fn (GF) given by (6), and G_B solves:

$(\nabla^2 + k^2)G_B = 0$ in Ω , and $G_f + G_B$ satisfy the boundary conditions (BCs).

Example: Ω half-space $x \geq 0$, with acoustically-soft boundary $S' = \{x = 0\}$.



Let $\underline{x}' = (x', y', z')$, $x' > 0$, a "source" point. Consider an "image" $\underline{x}_I = (-x', y', z')$
 $\Rightarrow \underline{x}_I \notin \Omega$; and take $G_B(\underline{x}, \underline{x}') = -G_f(\underline{x}, \underline{x}_I)$
 \Rightarrow (i) $(\nabla^2 + k^2)G_B = -\delta(\underline{x} - \underline{x}_I) = 0$ in Ω ;

BC; $x \in S \Leftrightarrow \underline{x} = (0, y, z)$

$$\begin{aligned} \Rightarrow G(\underline{x}, \underline{x}') &= G_f(\underline{x}, \underline{x}') - G_f(\underline{x}, \underline{x}'_{\perp}) \\ &= G_f(|\underline{x} - \underline{x}'|) - G_f(|\underline{x} - \underline{x}'_{\perp}|) = 0 \end{aligned}$$

Since $|\underline{x} - \underline{x}'| = |\underline{x} - \underline{x}'_{\perp}| \quad \forall \underline{x} \in S$

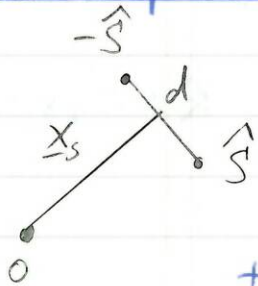
\Leftrightarrow acoustically soft BC.

Exercise: For acoustically hard BC, $\frac{\partial p}{\partial n} = \underline{n} \cdot \nabla p = 0$ on S , show that

$$G(\underline{x}, \underline{x}') = G_f(\underline{x}, \underline{x}') + G_f(\underline{x}, \underline{x}'_{\perp})$$

Exam 2012 Q4(b) \square .

3.3 Dipoles and Quadrupoles.



Take two monopole (= point) sources a small distance apart of amplitudes \hat{S} and $-\hat{S}$, i.e. oscillating "in anti-phase". Let $\pm \hat{S}$ be at $\underline{x}_s \pm \frac{1}{2} \underline{d}$, $|\underline{d}| = d$.

\Rightarrow Resulting acoustic field is

$$\begin{aligned} \hat{p}(\underline{x}) &= -4\pi \hat{S} G_f(\underline{x}, \underline{x}_s + \frac{1}{2} \underline{d}) \\ &\quad + 4\pi \hat{S} G_f(\underline{x}, \underline{x}_s - \frac{1}{2} \underline{d}) \end{aligned}$$

Let $kd \ll 1$; using Taylor Series:

$$\hat{p}(\underline{x}) = -4\pi \hat{S} G_{\text{rf}}(\underline{x}, \underline{x}_s) - 4\pi \hat{S} \nabla_s G_{\text{rf}}(\underline{x}, \underline{x}_s) \cdot \frac{1}{2} \underline{d} \\ + 4\pi \hat{S} G_{\text{rf}}(\underline{x}, \underline{x}_s) - 4\pi \hat{S} \nabla_s G_{\text{rf}}(\underline{x}, \underline{x}_s) \cdot \frac{1}{2} \underline{d} \\ + O(d^2)$$

$$= -4\pi \hat{\underline{\mathcal{D}}} \cdot \nabla_s G_{\text{rf}}(\underline{x}, \underline{x}_s) + O(d^2)$$

where $\hat{\underline{\mathcal{D}}} := \hat{S} \underline{d}$ is the dipole amplitude vector.
 Since $G_{\text{rf}}(\underline{x}, \underline{x}_s) = G_{\text{rf}}(\underline{x} - \underline{x}_s)$, see (6),
 $\Rightarrow \nabla_s G_{\text{rf}} = -\nabla G_{\text{rf}}$

\Rightarrow To main order in small d ,

$$\hat{p}(\underline{x}) = 4\pi \hat{\underline{\mathcal{D}}} \cdot \nabla G_{\text{rf}}(\underline{x}, \underline{x}_s) \\ = 4\pi \hat{\underline{\mathcal{D}}} \cdot \nabla \left(-\frac{1}{4\pi} \frac{e^{ik|\underline{x}-\underline{x}_s|}}{|\underline{x}-\underline{x}_s|^2} \right) \\ = -\hat{\underline{\mathcal{D}}} \cdot \frac{d}{dR} \left(\frac{e^{ikR}}{R} \right) \hat{\underline{R}}, \quad R := |\underline{x} - \underline{x}_s|$$

$$\hat{\underline{R}} = \frac{\underline{x} - \underline{x}_s}{R}$$

$$\Rightarrow \hat{p} = -\hat{\underline{\mathcal{D}}} \cdot \hat{\underline{R}} \frac{d}{dR} \left(\frac{e^{ikR}}{R} \right)$$

Exercise: Show that transversely oscillating rigid sphere (§ 2.4.2) produces a dipole field, and determine related $\hat{\underline{\mathcal{D}}}$. Hint: let $\underline{x}_s = 0$,

$$\hat{\mathcal{D}} = \mathcal{D} \hat{\underline{z}}$$

$$(*) \Rightarrow \hat{\underline{p}} = -\mathcal{D} \underbrace{\hat{\underline{z}} \cdot \underline{r}}_{\cos \theta} \frac{d}{dr} \left(\frac{e^{ikr}}{r} \right)$$

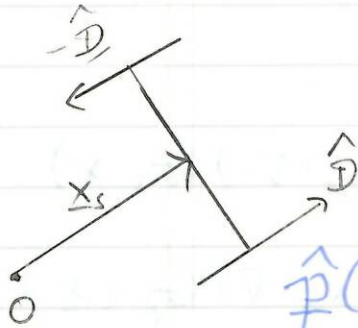
$$= -\mathcal{D} \cos \theta \frac{d}{dr} \left(\frac{e^{ikr}}{r} \right) \text{ cf (2.27) } \square$$

Morale so far : while radially pulsating sphere produces a monopole field, transversely oscillating one produces a dipole field.

\square

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Quadrupoles



Consider now two dipoles small distance $d = |d|$ apart, with opposite dipole vectors $\pm \hat{D} \Rightarrow$ (for d small $\Leftrightarrow kd \ll 1$)

$$\hat{p}(\underline{x}) = -4\pi \hat{D} \cdot \nabla_s G_f(\underline{x}, \underline{x}_s + \frac{1}{2} \underline{d})$$

$$+ 4\pi \hat{D} \cdot \nabla_s G_f(\underline{x}, \underline{x}_s - \frac{1}{2} \underline{d})$$

Taylor series

$$\approx -4\pi (\underline{d} \cdot \nabla_s) (\hat{D} \cdot \nabla_s) G_f(\underline{x}, \underline{x}_s)$$

$\nabla_s = -\nabla$

$$= +4\pi (\underline{d} \cdot \nabla) (\hat{D} \cdot \nabla) \left(\frac{1}{4\pi} \frac{e^{ikR}}{R} \right)$$

$$= \sum_{\alpha, \beta=1}^3 Q_{\alpha\beta} \frac{\partial^2}{\partial x_\alpha \partial x_\beta} \left(\frac{e^{ikR}}{R} \right) \quad (10)$$

where $R = |\underline{x} - \underline{x}_s|$, $\underline{x} = (x_1, x_2, x_3)$

$$Q_{\alpha\beta} = d_\alpha \hat{D}_\beta \quad (x, y, z)$$

(10) is a "quadrupole" field \leftrightarrow combination of 2nd derivatives of "monopole" field e^{ikR}/R .

3.4 Multipole expansions.

Take a cluster of monopole (point) sources!



$$p_n = \frac{\hat{S}_n e^{ik|\underline{x} - \underline{x}_n|}}{|\underline{x} - \underline{x}_n|} = -4\pi \hat{S}_n G_f(\underline{x} - \underline{x}_n)$$

such that $k|\underline{x}_n| \ll 1$ (i.e. close enough to each other), $n=1, 2, \dots, N$.

Using Taylor series:

$$G_{\underline{f}}(\underline{x} - \underline{x}_n) = G_{\underline{f}}(\underline{x}) - (\underline{x}_n \cdot \nabla) G_{\underline{f}}(\underline{x}) + \frac{1}{2} (\underline{x}_n \cdot \nabla) (\underline{x}_n \cdot \nabla) G_{\underline{f}}(\underline{x}) + \dots$$

\Rightarrow The resulting field $\hat{\underline{p}}$:

$$\hat{\underline{p}} = \sum_{n=1}^N \underline{p}_n = -4\pi \left(\sum_n \hat{S}_n \right) G_{\underline{f}}(\underline{x}) + 4\pi \sum_n \hat{S}_n (\underline{x}_n \cdot \nabla) G_{\underline{f}}(\underline{x}) - 4\pi \sum_n \frac{1}{2} \hat{S}_n (\underline{x}_n \cdot \nabla)^2 G_{\underline{f}}(\underline{x}) + \dots$$

$$= \underbrace{\hat{S}}_{\text{monopole}} \frac{e^{i\mathbf{k}\cdot\mathbf{r}}}{r} - \underbrace{\hat{\mathcal{D}}}_{\text{dipole}} \cdot \nabla \left(\frac{e^{i\mathbf{k}\cdot\mathbf{r}}}{r} \right) + \sum_{\alpha, \beta} \underbrace{Q_{\alpha\beta}}_{\text{quadrupole}} \frac{\partial^2}{\partial x_\alpha \partial x_\beta} \left(\frac{e^{i\mathbf{k}\cdot\mathbf{r}}}{r} \right) \quad (11)$$

where $\hat{S} := \sum_n \hat{S}_n$, $\hat{\mathcal{D}} = \sum_n \hat{S}_n \underline{x}_n$,

$$Q_{\alpha\beta} = \frac{1}{2} \sum_n \hat{S}_n x_{n_\alpha} x_{n_\beta}, \quad \underline{x}_n = (x_{n_1}, x_{n_2}, x_{n_3})$$

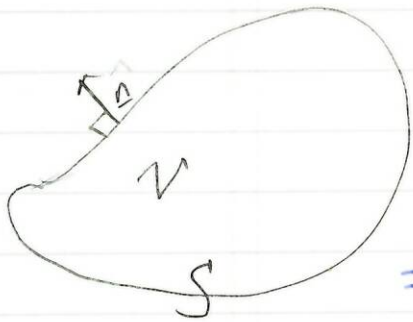
(11) is an example of a multipole expansion: monopole, dipole, quadrupole terms are the first 3 terms in an infinite hierarchy of "multipoles". So, monopole term dominates unless $\sum \hat{S}_n = 0$, then dipole term dominates etc.

3.5 Kirchhoff - Helmholtz integral Theorem

This is an important integral representation for solutions of Helmholtz eqn.

Based on "Green's formula (identity)": Let f, g be two functions in domain / volume V with boundary S .

Apply divergence theorem to $\underline{F} = f \nabla g - g \nabla f$.



$$\int_V \nabla \cdot \underline{F} dV = \int_S \underline{F} \cdot \underline{n} dS$$

$$\nabla \cdot \underline{F} = \nabla \cdot (f \nabla g) - \nabla \cdot (g \nabla f)$$

$$= \cancel{\nabla f \cdot \nabla g} + f \nabla^2 g - \cancel{\nabla g \cdot \nabla f} - g \nabla^2 f$$

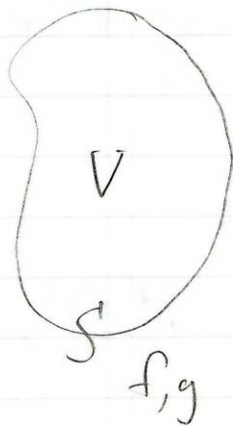
$$\underline{F} \cdot \underline{n} = f \nabla g \cdot \underline{n} - g \nabla f \cdot \underline{n} = f \frac{\partial g}{\partial n} - g \frac{\partial f}{\partial n}$$

$$\Rightarrow \int_V (f \nabla^2 g - g \nabla^2 f) dV = \int_S \left(f \frac{\partial g}{\partial n} - g \frac{\partial f}{\partial n} \right) dS \quad (12)$$

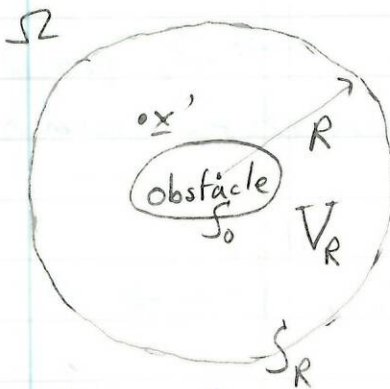
(12) is the Green's formula / Green's identity:

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$$\int_V (f \nabla^2 g - g \nabla^2 f) dV$$



$$= \int_S \left(f \frac{\partial g}{\partial n} - g \frac{\partial f}{\partial n} \right) dS \quad (12)$$



Choose in (12):
 $f(x) = \Psi = \text{soln of Helmholtz eqn (1.9),}$
 $(\nabla^2 + k^2)\Psi = 0.$

in a domain Ω exterior to a bounded "obstacle" with boundary S_0 ;

$$g(\underline{x}) = G(\underline{x}, \underline{x}') = \frac{e^{i|\underline{x}-\underline{x}'|}}{4\pi |\underline{x}-\underline{x}'|} \quad (6),$$

free-space Green's function, $G = G_f$, $\underline{x}' \in \Omega$;

$V = V_R = \text{part of } \Omega \text{ within ball } B_R \text{ of a large radius } R,$
 $\underline{x}' \in V_R, \text{ bounded } S_R \text{ (sphere of radius } R)$

Plugging into (12):

$$\int_{V_R} [\Psi(\underline{x}) \nabla^2 G(\underline{x}, \underline{x}') - G(\underline{x}, \underline{x}') \nabla^2 \Psi(\underline{x})] dV(\underline{x})$$

$$= \int_{S_0 \cup S_R} \left[\Psi(\underline{x}) \frac{\partial G(\underline{x}, \underline{x}')}{\partial n(\underline{x})} - G(\underline{x}, \underline{x}') \frac{\partial \Psi(\underline{x})}{\partial n} \right] dS(\underline{x}) = \text{RHS'}$$

$$\text{LHS} \stackrel{(5), (1.9)}{=} \int_{V_R} [\cancel{\Psi(\underline{x})(-k^2 G)} + \delta(\underline{x} - \underline{x}') + \cancel{G k^2 \Psi}] \, dV(\underline{x}) \quad \left| \begin{array}{l} \text{Recall (5) is} \\ \nabla^2 G + k^2 G = \delta(\underline{x} - \underline{x}') \end{array} \right.$$

$$= \int_{V_R} \Psi(\underline{x}) \delta(\underline{x} - \underline{x}') \, dV(\underline{x})$$

$$\stackrel{(3)}{=} \Psi(\underline{x}'), \text{ by (3)} \Rightarrow \text{LHS} = \Psi(\underline{x}');$$

RHS: Show that, as $R \rightarrow \infty$, $\int_{S_R} \dots \rightarrow 0$; provided Ψ satisfies also Sommerfeld radiation condn (1.19):

$$(1.19) \Leftrightarrow r \left(\frac{\partial \Psi}{\partial r} - ik\Psi \right) \rightarrow 0, \quad r \rightarrow \infty$$

$$\Leftrightarrow \frac{\partial \Psi}{\partial r} = ik\Psi + o(r^{-1})$$

↑ small "0"

Notice G given by (6) also satisfies (1.19)

$$G = -\frac{e^{ik|\underline{x} - \underline{x}'|}}{4\pi|\underline{x} - \underline{x}'|} \approx -\frac{e^{ikr}}{4\pi r};$$

Also $G = O(r^{-1})$, $\Psi = O(r^{-1})$, $r \rightarrow \infty$ (can be shown to follow from (1.19))

↑ capital "O"


$$\Rightarrow \int_{S_R} \left(\Psi \frac{\partial G}{\partial r} - G \frac{\partial \Psi}{\partial r} \right) dS$$

$$\begin{aligned}
 & \stackrel{(R \rightarrow \infty)}{=} \int_{S_R} [\cancel{\Psi_i k G} + \underbrace{\Psi_o(R^{-1})}_{= Q(R^{-1})} - \cancel{G_i k \Psi} - \underbrace{G_o(R^{-1})}_{= Q(R^{-1})}] dS \\
 & = \int_{S_R} Q(R^{-1}) dS \\
 & = o(R^{-2}) \times 4\pi R^2 \rightarrow 0.
 \end{aligned}$$

$R \rightarrow \infty$, as required.

So (12) reduces to:

$$\Psi(\underline{x}') = \int_{S_0} \left[\Psi(\underline{x}) \frac{\partial G(\underline{x}, \underline{x}')}{\partial n(\underline{x})} - G(\underline{x}, \underline{x}') \frac{\partial \Psi(\underline{x})}{\partial n} \right] dS(\underline{x}) \quad (13)$$


 Notice in (13), \underline{n} is exterior to $\Omega \leftrightarrow$ interior to obstacle O , so for $\underline{n}_s = -\underline{n}$, exterior to O , specialising (13) to $\Psi = \hat{p}$, acoustic pressure, swapping $\underline{x}' \rightarrow \underline{x}$, $\underline{x} \rightarrow \underline{x}_s$, also $G(\underline{x}', \underline{x}) = G(\underline{x}, \underline{x}')$, (6).

$$\hat{p}(\underline{x}) = \int_{S_0} \left[G(\underline{x}, \underline{x}_s) \frac{\partial \hat{p}(\underline{x}_s)}{\partial n} - \hat{p}(\underline{x}_s) \frac{\partial G(\underline{x}, \underline{x}_s)}{\partial n_s} \right] dS(\underline{x}_s) \quad (14)$$

(13), (14) are called Kirchhoff - Helmholtz integral theorem

Notice (1.41) $\leftrightarrow \hat{p} = i\omega \rho_0 \hat{\phi}$

$$\Rightarrow \frac{\partial \hat{p}}{\partial n} = i\omega \rho_0 \underline{n} \cdot \nabla \phi \stackrel{= \hat{u}}{=} = i\omega \rho_0 \hat{v}_n,$$

\hat{v}_n = "normal ^{component} velocity" of S_0 ;

$$(6) \Rightarrow G = \frac{-e^{ikR}}{4\pi R}, \quad R = |\underline{x} - \underline{x}_s|$$

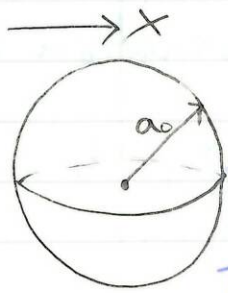
$$\frac{\partial G}{\partial n_s} = \underline{n}_s \cdot \nabla_s G = -\underline{n}_s \cdot \nabla G.$$

$$\Rightarrow \hat{p}(\underline{x}) = -\frac{1}{4\pi} \int_{S_0} i\omega \rho_0 \hat{v}_n(\underline{x}_s) \frac{e^{ikR}}{R} + \hat{p}(\underline{x}_s) (\underline{n}_s \cdot \nabla) \left(\frac{e^{ikR}}{R} \right) dS(\underline{x}_s) \quad (15)$$

Notice (15), (14) etc, express the unknown \hat{p} anywhere away from S_0 ($\underline{x} \in \Omega$) in terms of boundary only values of \hat{p} but also \hat{v}_n ($\leftrightarrow \frac{\partial \hat{p}}{\partial n}$).

However, we normally know, from BC's, only either \hat{p} or \hat{v}_n . Nevertheless, we normally know, from BC's, only either \hat{p} or \hat{v}_n . Nevertheless, if e.g. \hat{v}_n we know, let in (15) $\underline{x} \rightarrow S_0$ (\underline{x} to put on the boundary S_0) \Rightarrow (15) becomes a boundary integral equ (B.I.E) which needs to be solved by separate means. Then (15) constructs \hat{p} (and \hat{u}) everywhere. \square

3.6 Example: scattering by air bubbles in water.



Let air bubble of unperturbed radius a_0 in water, be subjected to an incident acoustic field; a_0 "small": $a_0 \ll \lambda = 2\pi/k$ the wavelength. Then the "incident field's pressure", p_i is: $p = Ae^{ikx - i\omega t} \approx Ae^{-i\omega t}$ ($|x| = a_0 \Rightarrow |kx| \ll 1$)

The bubble "scatters" p_i : the scattered field p_s , sough as outgoing spherically symmetric (§1)

$$\Rightarrow p_s = \frac{B}{r} e^{ikr - i\omega t}, \quad (B \in \mathbb{C}) \quad (16)$$

$$\text{Then (1.16)} \Leftrightarrow \underline{\hat{u}} = -\frac{c}{\rho_0 \omega} \nabla \hat{p} \Rightarrow$$

$$(17) v_s = \frac{1}{i\rho_0 \omega} \frac{\partial p_s}{\partial r} = \frac{B}{i\rho_0 \omega r^2} (ikr - 1) e^{ikr - i\omega t},$$

$v_s =$ "radial velocity".

We assume the bubble is made of an ideal gas, responding "adiabatically" $\Leftrightarrow PV^\gamma = \text{constant}$. ($\gamma > 1$ adiabatic constant).

Differentiating in time.

$$\frac{dp}{dt} V^\gamma + \gamma P V^{\gamma-1} \frac{dV}{dt} = 0.$$

pressure is not dependant on r since the bubble is so small.

$$\text{Let } P = P_0 + P', \quad V = \frac{4}{3}\pi(a_0 + a')^3$$

$P' \ll P_0$, $a' \ll a_0$ (perturbation small \Rightarrow)

$$\frac{dP'}{dt} = -\sigma P V^{-1} \frac{4}{3}\pi 3a_0^2 \left(\frac{da'}{dt}\right) = v_s$$

Now P' time-harmonic T-H $\Rightarrow \frac{dP'}{dt} = -i\omega P'$

$$\Rightarrow -i\omega P' = -\frac{\sigma P_0}{\frac{4}{3}\pi a_0^3} 4\pi a_0^2 v_s \quad \left(v_s = \frac{da'}{dt} : (17)\right)$$

$$P' = \frac{3\sigma P_0}{i\omega a_0} v_s \quad (18)$$

On the bubble surface, $r = a_0$, in absence of surface tension, etc, balance of pressures:

$$P_i|_{r=a_0} + P_s|_{r=a_0} = P'$$

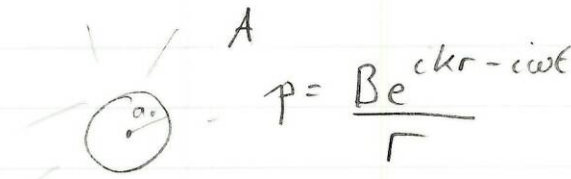
\Leftrightarrow via (16) - (18):

$$A e^{-i\omega t} + \frac{B}{a_0} e^{ika_0 - i\omega t} = \frac{3\sigma P_0}{i\omega a_0} \cdot \frac{B(ika_0 - 1)}{i\omega \rho a_0^2} \dots$$

$$\cdot e^{ika_0 - i\omega t}$$

$$a_0 A = B e^{i k a_0} \left[-1 + \frac{3 \mu P_0}{\omega^2 \rho_0 a_0^2} (1 - i k a_0) \right]$$

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$$p = \frac{B e^{i(kr - \omega t)}}{r}$$

$$a_0 A = B e^{i k a_0} \left[-1 + \frac{3 \gamma P_0}{\omega^2 \rho_0 a_0^2} (1 - i k a_0) \right]$$

-/-

Since $ka_0 \ll 1 \Rightarrow e^{i k a_0} = 1 + i k a_0 + \mathcal{O}((ka_0)^2)$

Also denoting $\omega_0^2 := \frac{3 \gamma P_0}{\rho_0 a_0^2}$

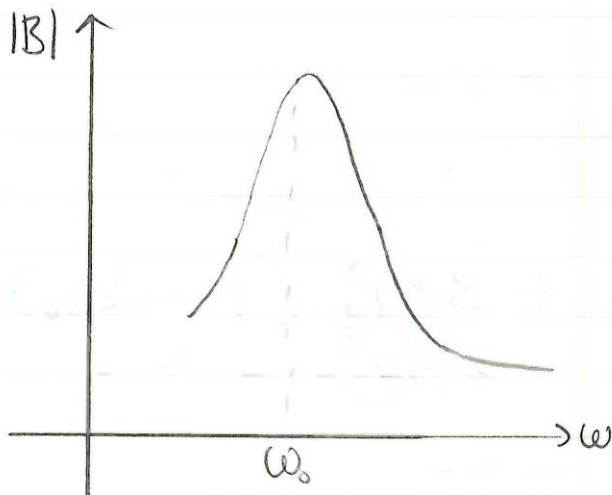
$$\Rightarrow a_0 A = B \left[-1 - i k a_0 + \frac{\omega_0^2}{\omega^2} \overbrace{(1 + i k a_0) \times (1 + k a_0)}^{1 - i k a_0} \right]$$

$$+ \mathcal{O}((ka_0)^2)$$

$$= B \left[\frac{\omega_0^2}{\omega^2} - 1 - i k a_0 + \mathcal{O}((ka_0)^2) \right]$$

$$\Rightarrow B \approx \frac{a_0 A}{\left(\frac{\omega_0}{\omega}\right)^2 - 1 - i k a_0}, \quad B \in \mathbb{C}.$$

$$\text{So } |B|^2 = \frac{a_0^2 |A|^2}{\left[\left(\frac{\omega_0}{\omega}\right)^2 - 1\right]^2 + (ka_0)^2}$$



$|B|$ has a peak at $\omega \approx \omega_0$.

Moral: ω_0 is a "resonance frequency" of the bubble: the scattered wave's amplitude surges for $\omega \approx \omega_0$.

$$\omega_0 \approx \left(\frac{3\sigma P_0}{\rho_0 a_0^3} \right)^{1/2}$$

4. High frequency waves, WKB method, Waveguides.

4.1 WKB method.

The WKB (Wentzel, Kramers, Brillouin) method is a method for constructing highly-oscillatory solutions of differential eqns containing a "large parameter" (e.g. ω , frequency).

Consider an ODE.

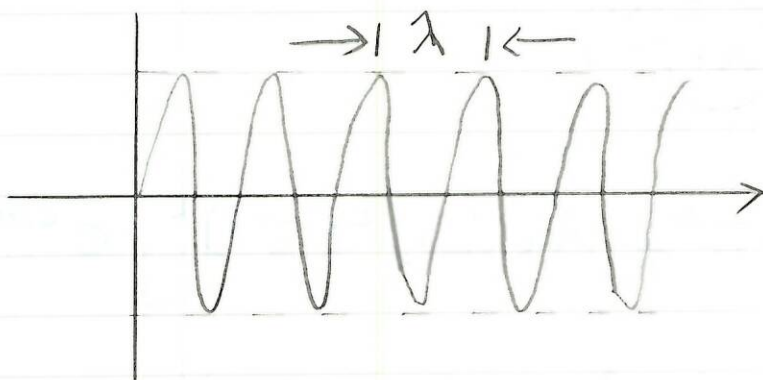
$$\frac{d^2 \Psi}{dx^2} + q(x) \Psi = 0.$$

where $|q(x)|$ is "large", in particular if $q(x) = k^2(x) > 0$.

$$\frac{d^2\psi}{dx^2} + k^2(x)\psi = 0. \quad (1)$$

In particular, if $k(x) \equiv k = \text{constant}$, (1) describes plane T-H waves (1-D Helmholtz eqn; with solutions $\psi = Ae^{ikx} + Be^{-ikx}$)

Eg: $\psi_1 = \text{Re}(Ae^{ikx - i\omega t}) = |A|\cos(\omega t - kx - \psi)$



When k large ($k \gg 1$)
 $\lambda = \frac{2\pi}{k} \ll 1$ small
 \rightarrow rapid oscillations;

$$k = \omega/c, \text{ so } k \gg 1 \Leftrightarrow \omega \gg 1$$

\Leftrightarrow High frequency.

Mathematically $k(x)$ "large", can be expressed via $k(x) = \omega\mu(x)$, where $\omega \gg 1$ is a large constant ($\omega = \text{frequency for waves}$) and $\mu = \mathcal{O}(1)$ "not large".

Then (1)

$$\Rightarrow \frac{d^2\psi}{dx^2} + \omega^2\mu^2(x)\psi = 0 \quad (2).$$

[For waves $k(x) = \omega/c(x)$, corresponding to "inhomogeneous media" with varying wave speed $c(x)$; $\Rightarrow \mu(x) = c^{-1}(x)$.

We seek the "WKB approximation" to a solution of (2) in the form.

$$\Psi(x, \omega) \approx A e^{i\omega\tau(x, \omega)}, \quad (3)$$

with A constant, $\omega \gg 1$. Then (3) \rightarrow (2)

$$\frac{d\Psi}{dx} = A i\omega \left(\frac{d\tau}{dx} = \tau' \right) e^{i\omega\tau}.$$

$$\frac{d^2\Psi}{dx^2} = \left[A i\omega \tau'' + A (i\omega)^2 (\tau')^2 \right] e^{i\omega\tau}.$$

$$\Rightarrow \cancel{i\omega} \tau'' - \omega^2 (\tau')^2 + \omega^2 \mu^2(x) = 0 \quad (4)$$

Now seek $\tau(x, \omega)$ as so-called "regular perturbation form" with respect to small ω^{-1} :

$$\tau(x, \omega) \sim \tau_0(x) + \omega^{-1} \tau_1(x) + \omega^{-2} \tau_2(x) + \dots \quad (5)$$

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$$\psi \approx A e^{i\omega \tau(x)} \quad (3)$$

$$i\omega \tau'' - \omega^2 (\tau')^2 + \omega^2 \mu^2 = 0 \quad (4)$$

$$\tau(x, \omega) \sim \tau_0(x) + \omega^{-1} \tau_1(x) + \dots \quad (5)$$

Plugging (5) into (4), to main order in ω , $\mathcal{O}(\omega^2)$

$$-\cancel{\omega^2} (\tau_0')^2 + \cancel{\omega^2} \mu^2(x) = 0$$

$$\Rightarrow \tau_0' = \pm \mu(x)$$

$$\tau_0(x) = \pm \int_{x_0}^x \mu(x') dx' \quad (6)$$

Next, for $\mathcal{O}(\omega)$ in (4):

$$i\omega \cancel{\tau_0''} - \cancel{\omega^2} 2\tau_0' \omega^{-1} \tau_1' = 0$$

$$\Rightarrow \tau_1' = \frac{i}{2} \frac{\tau_0''}{\tau_0'}$$

$$\Rightarrow \tau_1(x) = \frac{i}{2} \ln |\tau_0'| + C_{\pm}$$

$$= \frac{i}{2} \ln \mu(x) + C_{\pm}$$

Hence, from (3), the WKB approximation is:

$$\psi(x, \omega) \approx A e^{i\omega(\tau_0 + \omega^{-1} \tau_1)}$$

$$= A e^{i\omega t_0 + iT},$$

$$= A \exp \left\{ \pm i\omega \int_{x_0}^x \mu(x') dx' - \frac{1}{2} \ln \mu + \tilde{C}_{\pm} \right\}$$

$$= \tilde{A}_{\pm} \mu^{-1/2}(x) \exp \left\{ \pm i\omega \int_{x_0}^x \mu(x') dx' \right\} \text{ or}$$

$$\Psi(x, \omega) \approx \mu^{-1/2}(x) \left[A e^{i\omega \int_{x_0}^x \mu(x') dx'} + B e^{-i\omega \int_{x_0}^x \mu(x') dx'} \right]$$

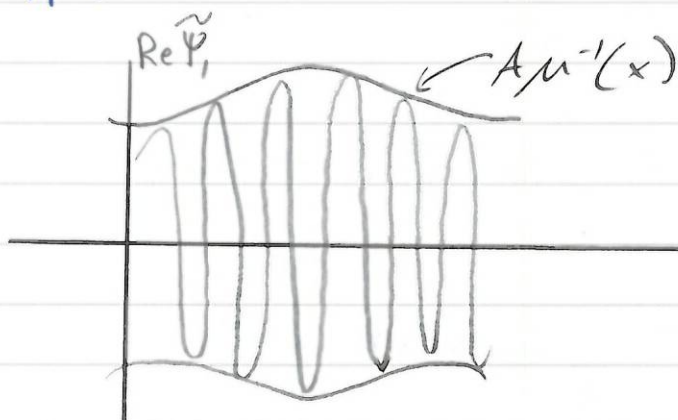
$$=: \tilde{\Psi}(x, \omega) \quad (7).$$

where A and B arbitrary constants.

The WKB approximation (7) contains two ^{high frequency} "waves", travelling to the right and left resp;

$$\tilde{\Psi} = \tilde{\Psi}_1 + \tilde{\Psi}_2$$

e.g. for $\tilde{\Psi}_1$:



Remark: If we substitute $\tilde{\Psi}$ into (2), then (Exercise):

$$\frac{d^2 \tilde{\Psi}}{dx^2} + \omega^2 \mu^2(x) \tilde{\Psi}(x) = \left[\frac{3}{4} \frac{(\mu')^2}{\mu^2} - \frac{\mu''}{2\mu} \right] \tilde{\Psi}$$

This suggests the (7) is accurate, provided

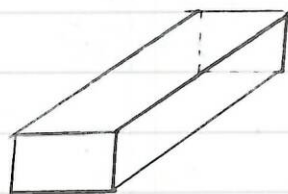
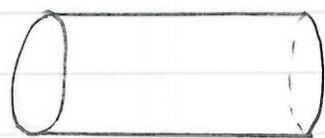
$$\frac{3}{4} \frac{(\mu')^2}{\mu^2} - \frac{\mu''}{2\mu} \ll \omega^2 \mu^2 \Leftrightarrow$$

$$\frac{1}{\omega^2} \left[\frac{3(\mu'(x))^2}{4\mu^4(x)} - \frac{\mu''(x)}{2\mu^3(x)} \right] \ll 1 \quad (8)$$

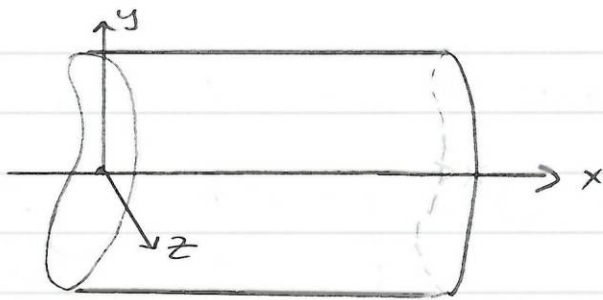
So this holds provided $\omega \gg 1$ and $\mu(x) \geq \mu_0 > 0$ (not too close to zero) and μ', μ'' "not too large" (slowly enough varying).

Exercise: 1. EM wave in layered media (Exam 2011 Qn 5)
2. Exam 2010 Qn 2.

4.2 Waveguides.



A waveguide is typically a tube / pipe of rather arbitrary cross-section, along which waves can propagate. (Physically, the wave bounces / is reflected from the walls, and propagates down the axis. *eg a cylinder.*)



Choose x along the waveguide's axis, seek for T-H solutions. So, both in acoustics and EM, we seek solutions of Helmholtz eqn (1.15) or (1.35) - (1.36), for $\hat{\Psi}$:

$$(\nabla^2 + k^2)\hat{\Psi} = 0 \Leftrightarrow \frac{\partial^2 \hat{\Psi}}{\partial x^2} + \overbrace{\frac{\partial^2 \hat{\Psi}}{\partial y^2} + \frac{\partial^2 \hat{\Psi}}{\partial z^2}}^{\nabla_{\perp}^2} + k^2 \hat{\Psi} = 0.$$

$k = \omega/c$. In both cases $\hat{\Psi}$ is sought in the form

$$\hat{\Psi} = \Psi(y, z) e^{i\beta x} \quad (9)$$

where "propagation constant" β is unknown, so (9) looks like a plane wave along x with "amplitude" Ψ depending on y, z .

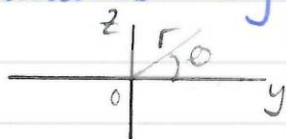
$\Rightarrow \Psi$ solves:

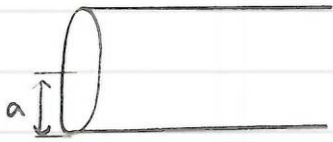
$$\nabla_{\perp}^2 \Psi + (k^2 + \beta^2) \Psi = 0 \quad (10)$$

where

$$\nabla_{\perp}^2 = \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} = \frac{\partial^2}{\partial r^2} + \frac{1}{r} \frac{\partial}{\partial r} + \frac{1}{r^2} \frac{\partial^2}{\partial \theta^2}$$

is "cross-sectional Laplacian" on (y, z) -plane (r, θ) polar coordinates ($y = r \cos \theta, z = r \sin \theta$).





Now specialise to acoustics and circular waveguides: $r \leq a$.

Seek solutions of (10) via separation of variables:

$$\Psi = R(r) \Theta(\theta).$$

$$\Rightarrow \frac{r^2}{R} R'' \Theta + \frac{1}{r} R' \Theta + \frac{1}{r^2} R \Theta'' + (k^2 - \beta^2) R \Theta = 0.$$

$$\Leftrightarrow \frac{r^2 R''}{R} + \frac{r R'}{R} + r^2 (k^2 - \beta^2) = - \frac{\Theta''}{\Theta} = m^2$$

(m^2 separation constant).

$$\Rightarrow \Theta'' + m^2 \Theta = 0$$

$$\Rightarrow \Theta = e^{\pm i m \theta}$$

So since Θ must be 2π -periodic in θ

$$\Rightarrow m \in \mathbb{Z}, \text{ integer};$$

$$\Rightarrow r^2 R'' + r R' + \left[\underbrace{r^2 (k^2 - \beta^2)}_{\tilde{r}^2} - m^2 \right] R = 0 \quad (11)$$

Via change of variables, $\tilde{r} = r(k^2 - \beta^2)^{1/2}$.

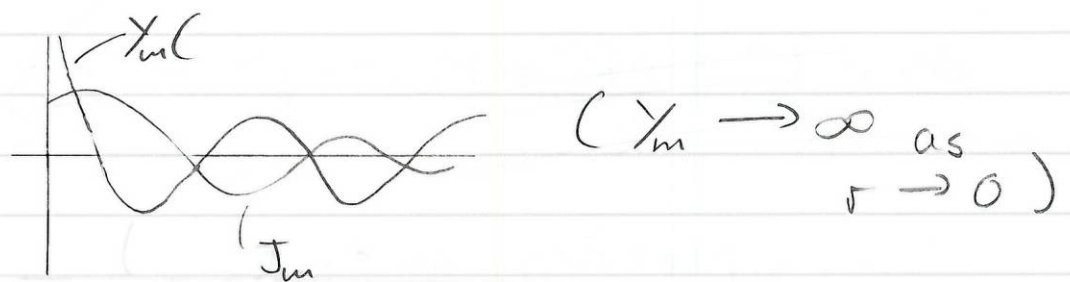
$$\Rightarrow \tilde{r}^2 \frac{d^2 R}{d\tilde{r}^2} + \tilde{r} \frac{dR}{d\tilde{r}} + [\tilde{r}^2 - m^2] R = 0 \quad (11')$$

This is known to be Bessel's differ. eqn' of order m , with general soln:

$$R = A J_m(\tilde{r}) + B Y_m(\tilde{r}) = 0 \quad (12)$$

A, B arbitrary const's, J_m the Bessel function, Y_m (so-called) Neumann function.

The solution has to be a smooth for $0 \leq r \leq a$ in particular for $r \rightarrow 0$, which J_m is but Y_m is not.



$\Rightarrow B=0$, and our "modal" solutions are

$$\hat{\Psi}(x, y, z) = A_m J_m(r(k^2 - \beta^2)^{1/2}) e^{i m \theta - i \beta z} \quad (13)$$

$$m \in \mathbb{Z}, A_m \in \mathbb{C}$$

Let the waveguide's boundary $r=a$ be e.g. acoustically hard i.e.:

$$\frac{\partial \hat{\Psi}}{\partial n} = \frac{\partial \hat{\Psi}}{\partial r} = 0, \text{ as } r=a.$$

\Rightarrow from (13), $J_m'(a(k^2 - \beta^2)^{1/2}) = 0$, so if α_{mn} , $n=1, 2, 3, \dots$ are roots of $J_m(\alpha_{mn}) = 0$.

$$\Rightarrow a(k^2 - \beta^2)^{1/2} = \alpha_{mn}.$$

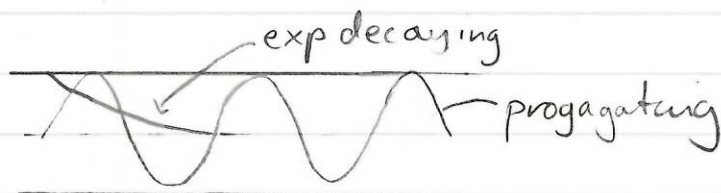
$$\Rightarrow \beta = \beta_{mn} = \pm \left(k^2 - \frac{\alpha_{mn}^2}{a^2} \right)^{1/2}, \quad n=1, 2, 3, \dots$$

So \exists infinitely many "modal solu"

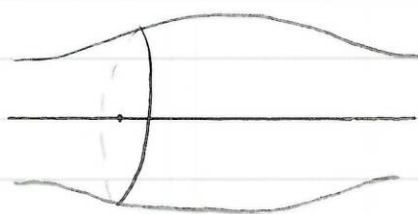
$$\hat{\Psi}_{mn} = A_{mn} J_m \left(\alpha_{mn} \frac{r}{a} \right) e^{im\theta - i\beta_{mn}x}.$$

$$m=0, 1, 2, \dots; \quad n \geq 1.$$

Notice however that for fixed k, m $\alpha_{mn} \rightarrow \infty$ as $n \rightarrow \infty$ so β_{mn} becomes complex for $n > (k, m)$: the corresponding modes are exponentially decaying in x .



4.3 Slowly varying waveguides.

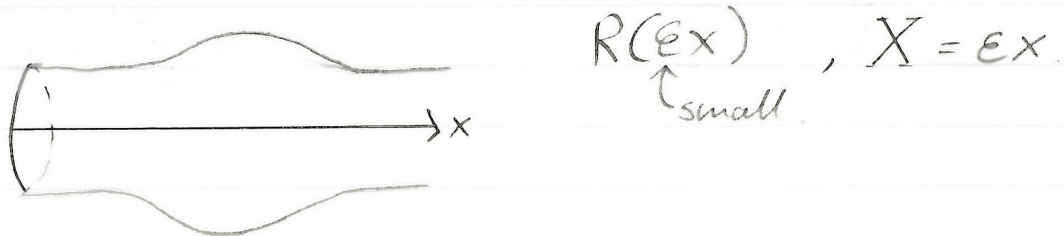


Let the waveguide's cross-section be circular but with a "slowly varying" radius R . Mathematically, let $\epsilon > 0$ be a "small parameter," $\epsilon \ll 1$, and let $R = R(\epsilon x)$: then

$$\frac{dR}{dx} = \epsilon R'(\epsilon x) = \epsilon R'(X) = \mathcal{O}(\epsilon).$$

where $X := \epsilon x$ is the so-called "Slow variable".

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Seek soln in the form, cf (14).

$$\hat{\phi} = A(X, r) e^{im\theta} \exp\left(\frac{i}{\epsilon} \int_0^X \mu(X') dX'\right) \quad (15)$$

Substitute to Helmholtz eqn in (x, r, θ) coord:

$$\frac{\partial^2 \hat{\phi}}{\partial x^2} + \frac{\partial^2 \hat{\phi}}{\partial r^2} + \frac{1}{r} \frac{\partial \hat{\phi}}{\partial r} + \frac{1}{r^2} \frac{\partial^2 \hat{\phi}}{\partial \theta^2} + k^2 \hat{\phi} = 0. \quad (16)$$

$$\Rightarrow \left(X = \epsilon x, \quad \frac{\partial}{\partial x} = \epsilon \frac{\partial}{\partial X} \right)$$

$$\Rightarrow -\mu^2(X) A(X, r) + 2\epsilon i \mu \frac{\partial A}{\partial X} + i \epsilon \mu'(X) A$$

$$+ \epsilon^2 \frac{\partial^2 A}{\partial X^2} + \frac{\partial^2 A}{\partial r^2} + \frac{1}{r} \frac{\partial A}{\partial r} - \frac{m^2}{r^2} A + k^2 A = 0 \quad (17)$$

BCs: $0 = \frac{\partial \Phi}{\partial n} = \underline{n} \cdot \nabla \Phi$, $\underline{n} = \underline{\tilde{\Gamma}} - \frac{dR}{dx} \underline{\tilde{x}}$



$$= \underline{\tilde{\Gamma}} - \epsilon R'(x) \underline{\tilde{x}}$$

$$\Rightarrow \frac{\partial \Phi}{\partial n} = \frac{\partial \Phi}{\partial r} - \epsilon R'(X) \frac{\partial \Phi}{\partial x} = 0, \quad \Gamma = R(X)$$

\Rightarrow via (15)

$$\frac{\partial A}{\partial r} - \epsilon R'(X) \left(i\mu A + \epsilon \frac{\partial A}{\partial X} \right) = 0 \quad (18)$$

Seek $A(X, r)$ as perturbation expansion in small ϵ :

$$A = A_0(X, r) + A_1(X, r) + \dots \quad (19)$$

Plugging into (17), (18), to leading orders:

$$\Delta A := \frac{\partial^2 A_0}{\partial r^2} + \frac{1}{r} \frac{\partial A_0}{\partial r} + \left(k^2 - \mu^2(X) - \frac{m^2}{r^2} \right) A_0 = 0 \quad (20)$$

$$(18) \Rightarrow \frac{\partial A_0}{\partial r} = 0, \quad \Gamma = R(X) \quad (21)$$

(20) - (21) appear identical to § 4.2, $a \rightarrow R(X)$, X parameter, cf (11). \Rightarrow cf (14)

$$A_0(X, r) = N(X) J_m(\gamma_{mn}(X) r), \quad (22)$$

where

$$\gamma_{mn}(X) = (k^2 - \mu^2(X))^{1/2} \stackrel{(18)}{=} \frac{\alpha_{mn}}{R(X)} \quad J_m'(\alpha_{mn}) = 0$$

$$\Rightarrow \mu = \mu_{mn}(X) = \left(k^2 - \frac{\alpha_{mn}^2}{R^2(X)} \right)^{1/2} \quad (23)$$

Notice $N(X)$ still unknown, and is found by equating next-order terms in (17) - (18):

$$\mathcal{L} A_1 = -2i\mu \frac{\partial A_0}{\partial X} - i\mu' A_0$$

$$= -i \frac{1}{A_0} \frac{\partial}{\partial X} (\mu A_0^2) \quad (24)$$

$$(18) \Rightarrow \frac{\partial A_1}{\partial r} = i\mu R'(X) A_0, \quad r = R(X) \quad (25)$$

Take (24) $\times r A_0 - (20) \times A_1 r$, and integrate in A from 0 to $R(X)$:

$$\text{LHS} = \int_0^{R(X)} (\mathcal{L} A_1) r A_0 - (\mathcal{L} A_0) r A_1 \, dr$$

$$\stackrel{\frac{d}{dr} :=}{=} \int_0^R \left(\frac{1}{r} (r A_1')' A_0 - \frac{1}{r} (r A_0')' A_1 \right) dr$$

By parts

$$= \left[\underbrace{\Gamma A_1 A_0}_{(25)} - \underbrace{\Gamma A_0^2 A_1}_{(21)} \right]_0^{R(x)}$$

(25)+(21)

$$= R i \mu R'(x) A_0^2(x, R(x))$$

$$\text{RHS} = -i \int_0^{R(x)} \frac{\partial}{\partial x} (\mu A_0^2 r) dr = i \mu R' R A_0^2.$$

$$\Rightarrow \int_0^{R(x)} \frac{\partial}{\partial x} (\mu A_0^2 r) dr + \mu R'(x) A_0^2 R = 0.$$

$$\Leftrightarrow \frac{d}{dx} \left(\int_0^{R(x)} \mu A_0^2 r dr \right) = 0 \text{ (Leibintz's rule).}$$

$$\Rightarrow \mu(x) \int_0^{R(x)} N^2(x) J_m^2 \left(\alpha_{mn} \frac{\Gamma}{R(x)} \right) r dr.$$

= const.

$$\Leftrightarrow \left(\frac{\Gamma}{R(x)} = \tilde{r} \right) \text{ Change of variable.}$$

$$\Rightarrow \mu(x) N^2(x) R^2(x) \int_0^1 J_m^2(\alpha_{mn} \tilde{r}) \tilde{r} d\tilde{r}$$

= constant.

$$\Rightarrow \mu N^2 R^2(x) = Q^2 \text{ (= constant).}$$

$$\Rightarrow N(X) = \frac{Q}{R(X) \sqrt{\mu_{mn}(X)}}$$

μ_{mn} given by (23)

Exercise: 2-D channels: Exam 2010 Qn 5,
EM version Exam 2012, Qn 5.

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